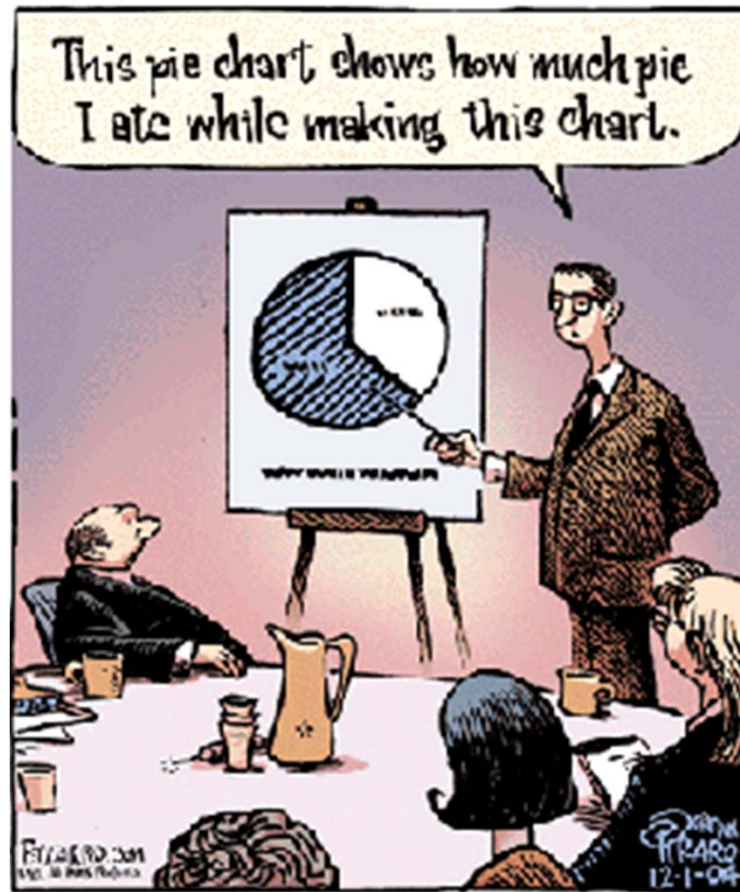


Tips for creating and giving scientific presentations



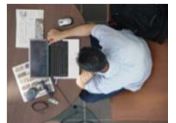
How to get started?

Step 1: Identify your audience: this will control the level of your presentation and the amount of background material you need to orient everyone in the audience

Step 2: Determine how much time you have for your presentation: this will control how much time you have to talk about each part of your outline (see below)

Step 3: Identify the main points you want to convey: you can reasonably convey only 2-3 main points in a 20- or 30-minute talk

Step 4: Create an outline of your talk: this will build in the logical organization of your presentation and help you decide what figures and other supporting evidence you need to make your points



Organizing a 20-minute scientific talk

Background and Introduction (~5 minutes)

⇒ 4–5 slides

~1 Title slide - Your names, date, citation to paper

~1 Outline slide – Organization of talk

~1 Overview slide – Why is this research important?

~1-2 Background slides – Provides essential background for non-experts

Methods (~5 minutes)

⇒ 2–3 slides

Theoretical/experimental methods used in paper

Organizing a 20-minute scientific talk

Results (~7 minutes)

⇒ 4–5 slides

~ What did you (or the authors) find?

Only develop 1-2 key results

*Critique and Citation Summary (~2 minutes)

⇒ 2 slides

1 critique slide – What was wrong with/good about the paper?

1 citation slide – What happened with the result/field after the paper?

Summary (~1 minute)

⇒ 2 slides

1 Summary slide - Review the main points/ criticisms

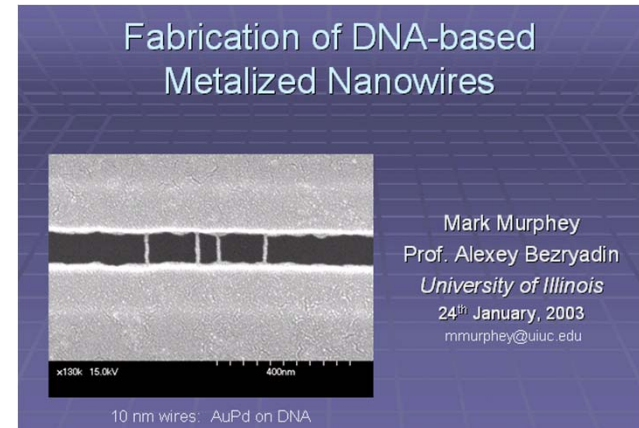
1 Acknowledgment slide – Acknowledge sources of material, help received, etc.

* Journal club only

The title slide and outline prepares the audience to listen and shows organization of talk

Title slide

- Your names and affiliations
- Paper citation (for JC)
- Venue and date
- Attention-getting graphic



Outline or overview of presentation

- Prepares the audience to listen
- Provides a logical structure for your talk
- Provides motivation and context
- Summarizes key points (limit to two or three for a 20- to 30-minute talk)



Particle Physicists Ask ...

1. Why matter?

- CP Violation

2. Why mass?

- Higgs field

3. Why this standard model?

- SUSY or other extensions

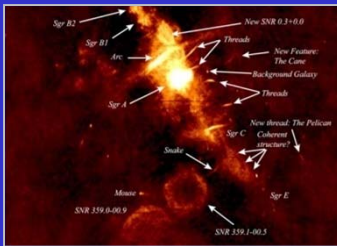


Overview

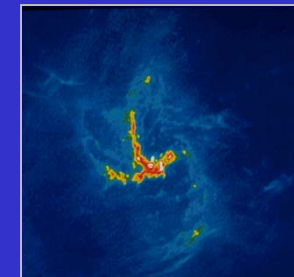
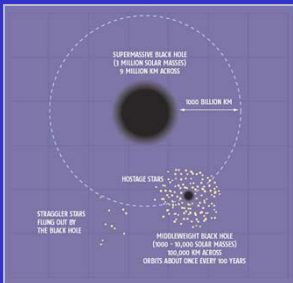
Black holes and star clusters



The galactic center



Intermediate-mass black hole kinematics



Here, we have a VISUAL and WRITTEN outline and it's not too long !

The “body” of your presentation is the intellectual content of your talk

Problem statement, motivation

~1–2 slides

Previous work

~1-2 slides

Methods

~1–3 slides

Key Results

~5–6 slides

What's this all about?

- **Quark Gluon Plasma (QGP)**
 - hypothesized state of deconfined quarks and gluons
 - expected in the early universe, big-bang stage



– this represents a phase change

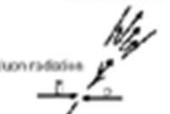


– like “melting” a nucleus

pp vs A-A collisions

hard-scattering
correlations $\langle \dots \rangle$

cone of hadrons




gluon radiation

soft-scattering
correlations $\langle \dots \rangle$


hadron cone? debris scattered?

Energy-loss:
increased medium-induced
gluon radiation



High p_T

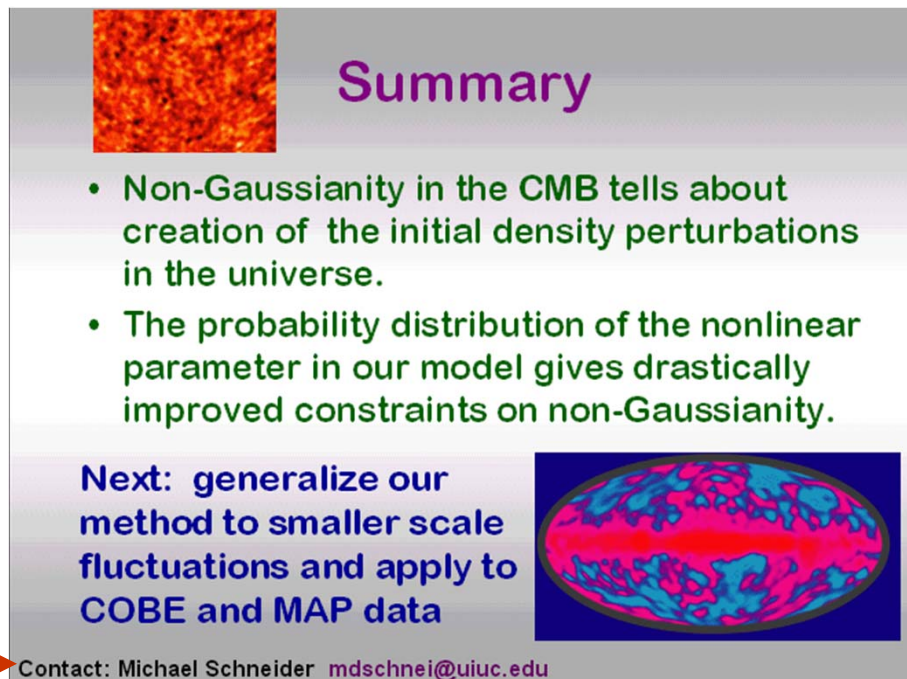
- Hard-scattering occurs at earliest times during a high-energy nuclear collision
 - will **before QGP is expected to form (equilibrates)**
- These “fast” scatterers will experience the strongly interacting medium created in the collision.
- They will lose energy in the hot, dense medium (by **gluon bremsstrahlung**) and their outgoing energy distribution will be modified downward.
- Net effect, depletion in yield of high p_T particles compared to p-p case.
- The game is then to make the comparison.



Provide a “summary” slide

Recap key results and conclusions

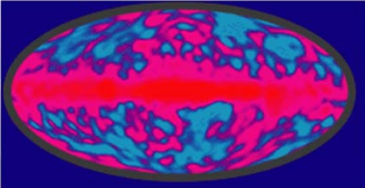
Reiterate main critiques (for JC)



Summary

- Non-Gaussianity in the CMB tells about creation of the initial density perturbations in the universe.
- The probability distribution of the nonlinear parameter in our model gives drastically improved constraints on non-Gaussianity.

Next: generalize our method to smaller scale fluctuations and apply to COBE and MAP data



Contact: Michael Schneider mdschnei@uiuc.edu

An orange arrow points to the contact information at the bottom left of the slide.

This slide will probably stay on the screen during the question period and will thus get the longest audience exposure—make it count!



Summary & Conclusions

Not “exciting” but it has the pieces

- All g-2 data published

- ❖ Systematics lowered again

What was shown

- Consistent results, consistently above theory

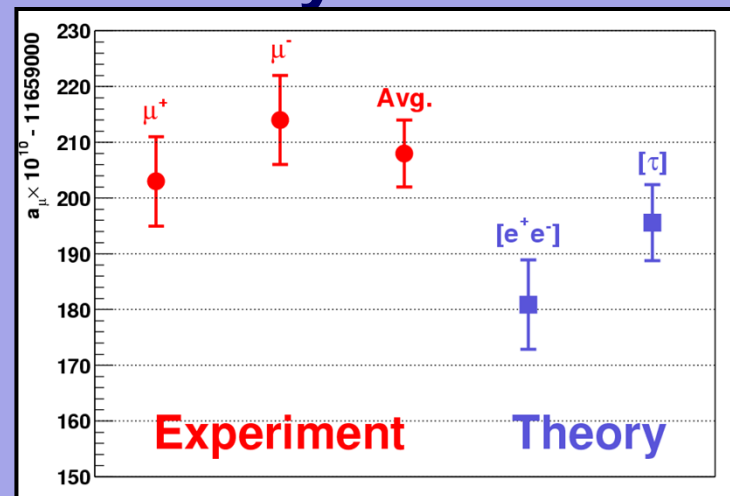
- ❖ ee – tau controversy still quite active
- ❖ considerably more “ee” type data on the way

Where things stand:
summarized nicely
on the plot

- The systematic limit is “far” away ...we should go there

What to do next

Note e-mail and web link



Tips for preparing your talk (cont.)

Have only 1 idea per slide

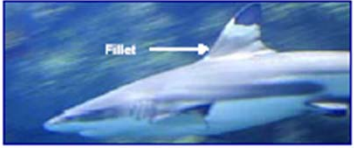
Use the header to state the main idea of the slide, and use the body of the slide to support that idea

Use well-labeled graphs and figures to illustrate your key points...this makes the slide more real and interesting to the audience

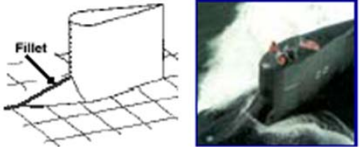
Avoid too much text....

Fillets reduce leading edge vortices in nature and in engineering


Fillet on dorsal fin of shark



Fillet on Seawolf submarine



[Devenport et al., 1991]



Literature Review

- Hefner developed a dynamic model of the thermal behavior of a temperature-dependent IGBT electrical device. The model is expressed in terms of the instantaneous power dissipation and the thermal capacitance and thermal conductance of the silicon chip and the SABER circuit simulation.
- Adam et al. investigated the interactions between the heat sources, the thermal conductivity of the walls and the thermal capacitance of the walls and the thermal behavior of discretely heated enclosures.
- Chen, Wu and ... are modeling of thermal and electrical behavior using several commercial softwares (I-DEAS, Maxwell, Flotherm and Saber) and 3-D, transient approaches.

Too many words

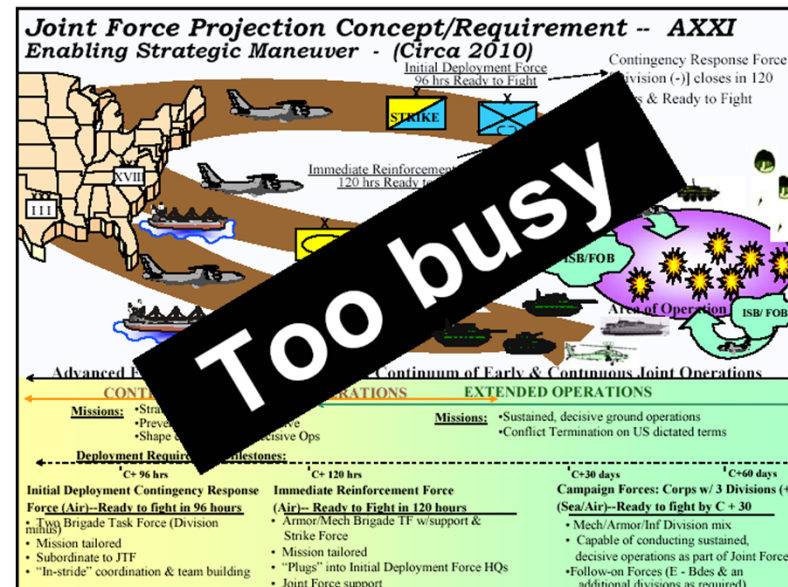
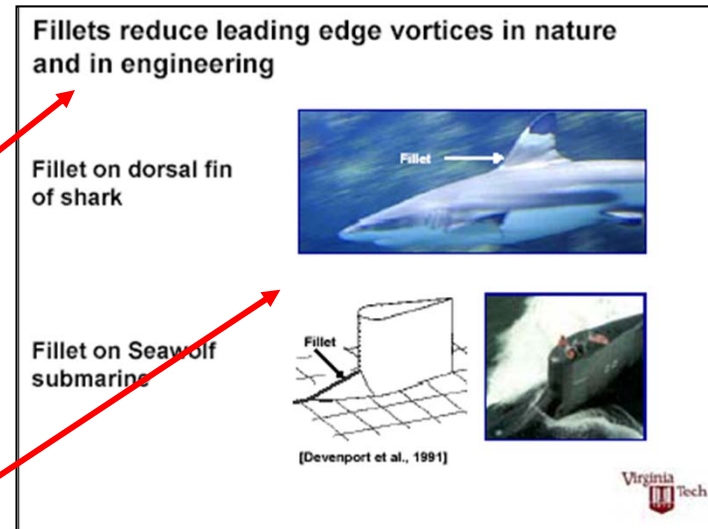
Tips for preparing your talk (cont.)

Have only 1 idea per slide

Use the header to state the main idea of the slide, and use the body of the slide to support that idea

Use well-labeled graphs and figures to illustrate your key points...this makes the slide more real and interesting to the audience

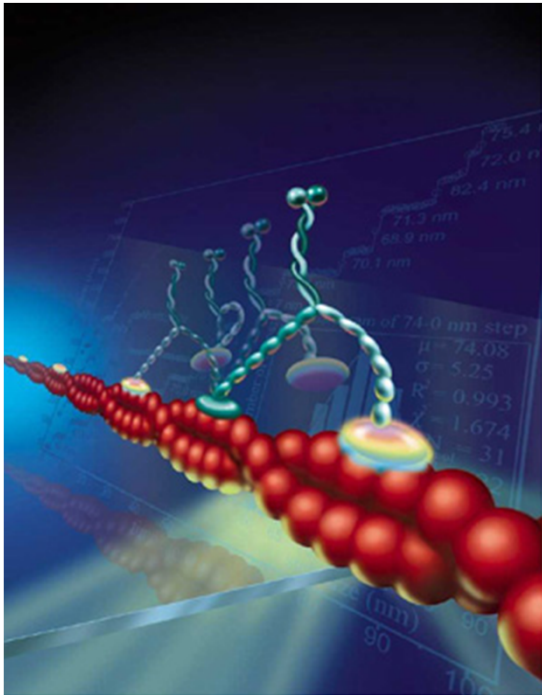
....or too many distracting images



Use figures to illustrate your key points

Figures:

- enliven slides
- promote audience interest
- provide supporting evidence for key points
- help explain complex ideas and relationships quickly
- show how things work, etc.

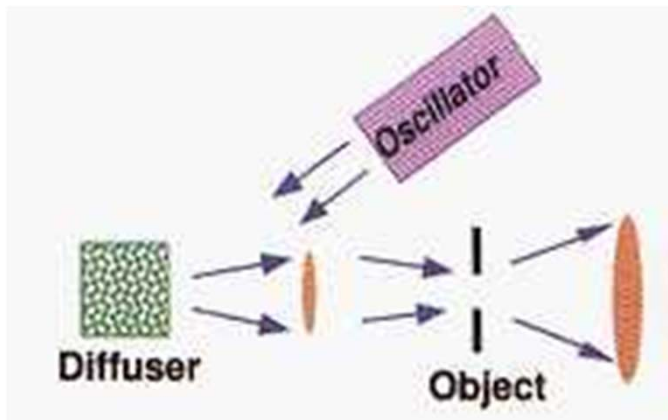


Myosin “walking” on actin
Courtesy of P. Selvin

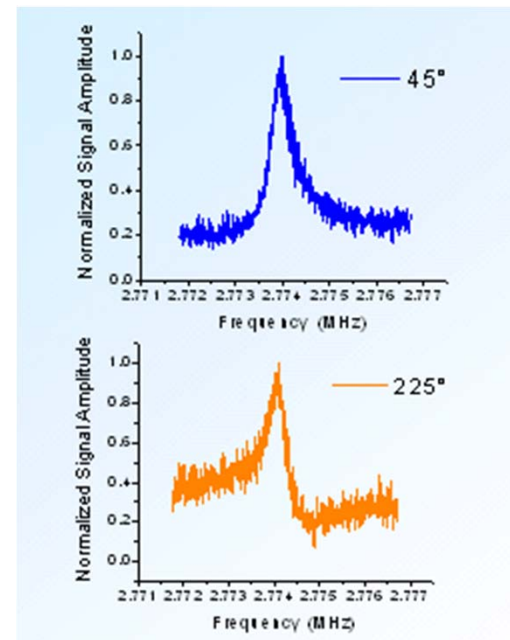


Label all elements in a figure

- Point out important features
- Label both axes of graphs and show units
- Provide a brief caption
- Give credit to source



The Nike laser system uses discharge pre-amplifiers.
(Courtesy US Navy)

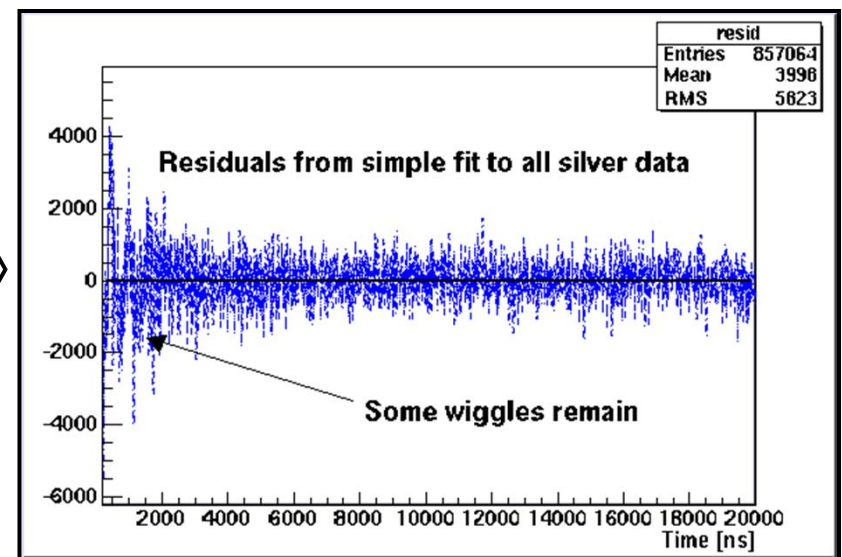
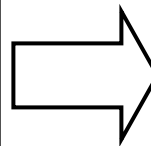
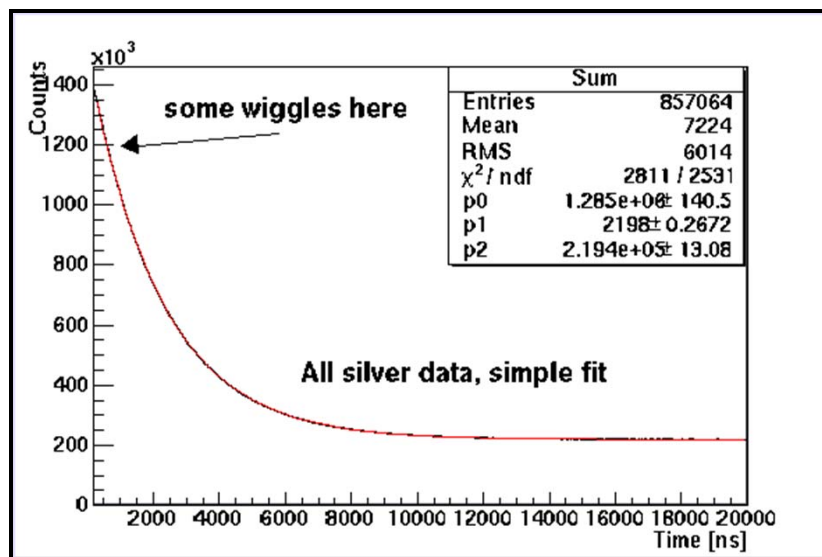


Sample normalized signals from the two-beam optical drive.
(Courtesy C. Michael)



Presenting data is your most important and challenging task

- Avoid copying a graph from a formal article – they have a different style, **e.g., labels are too small**
- Use color and make lines thick, labels legible
- Label axes and annotate important points with arrows and add words
- Use tables sparingly – if used highlight important parts



Show the equipment **IF** it helps as part of your proof – but sparingly, not just because you love it

- **Photographs** give scale and reality – but add labels
- **Schematics** provide concept
- **Diagrams** strip away unnecessary details
- **ALL OF THESE** can be useful in combination

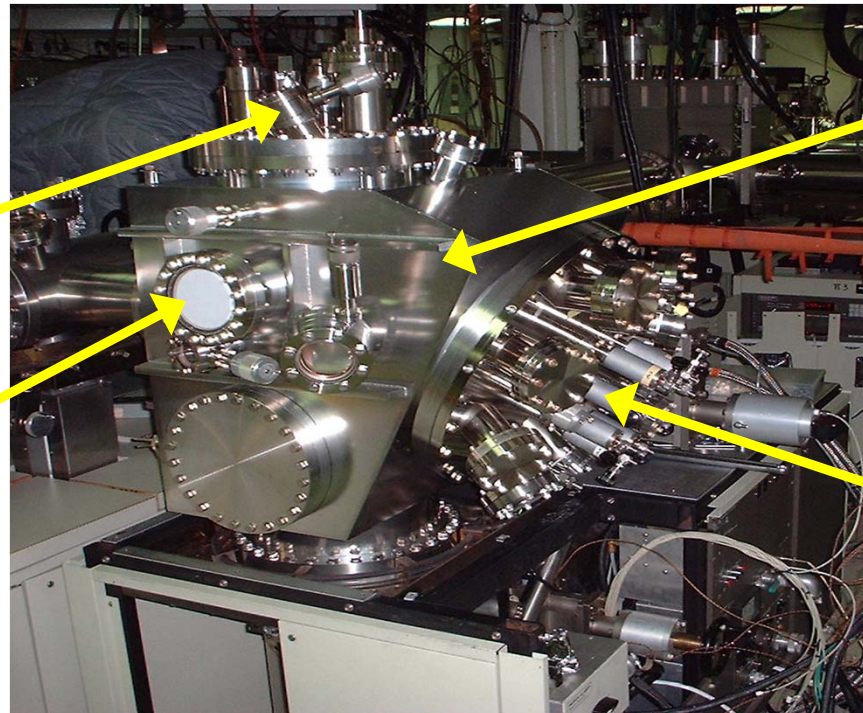
Mass spectrometer

RHEED screen

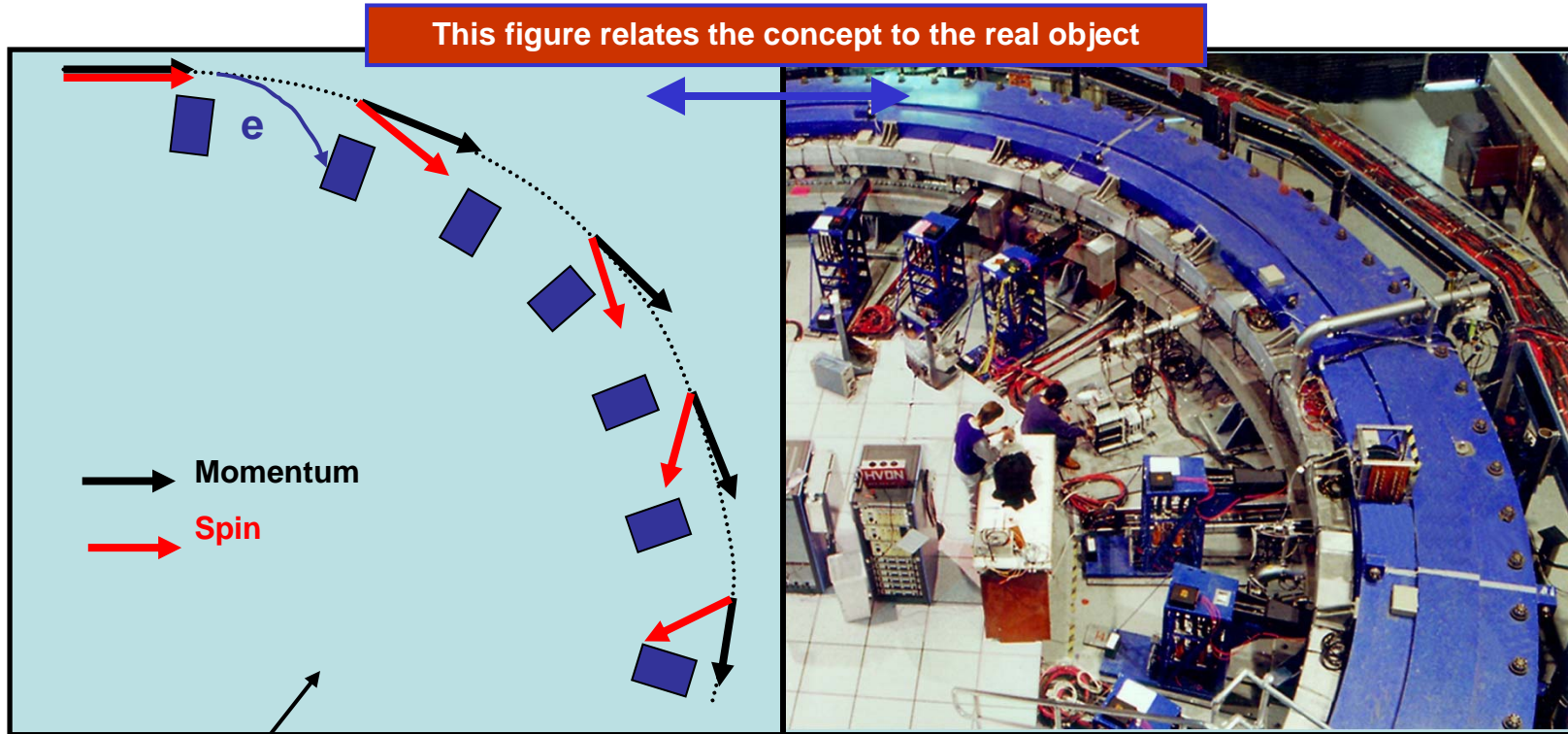
Vacuum chamber

Source flanges

OK, but could be better



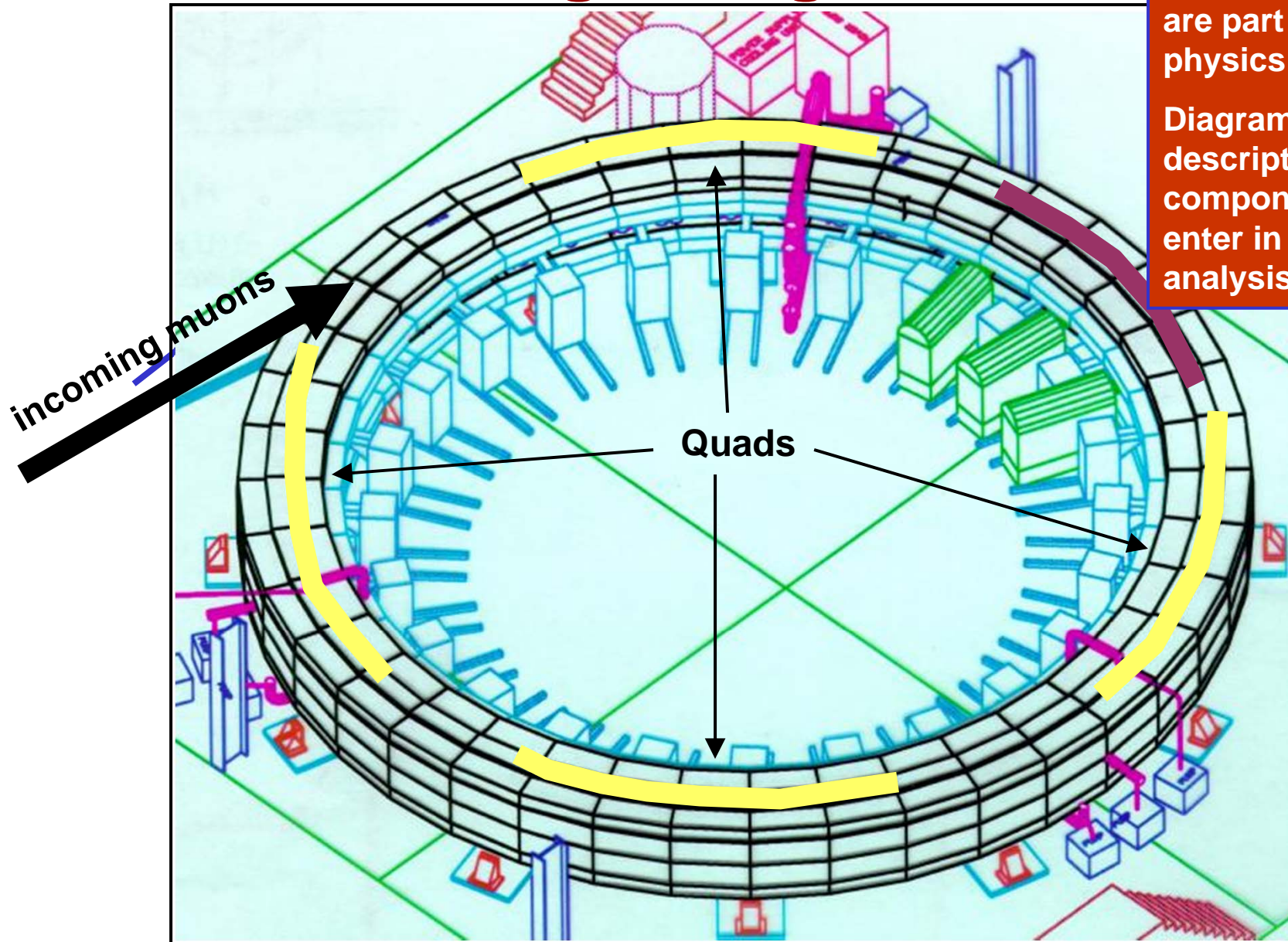
a_μ is proportional to the difference between the spin precession and the rotation rate



This supports assertion in sentence headline

$$\Delta\omega = \omega_a = \left(\frac{g - 2}{2} \right) \frac{eB}{mc}$$

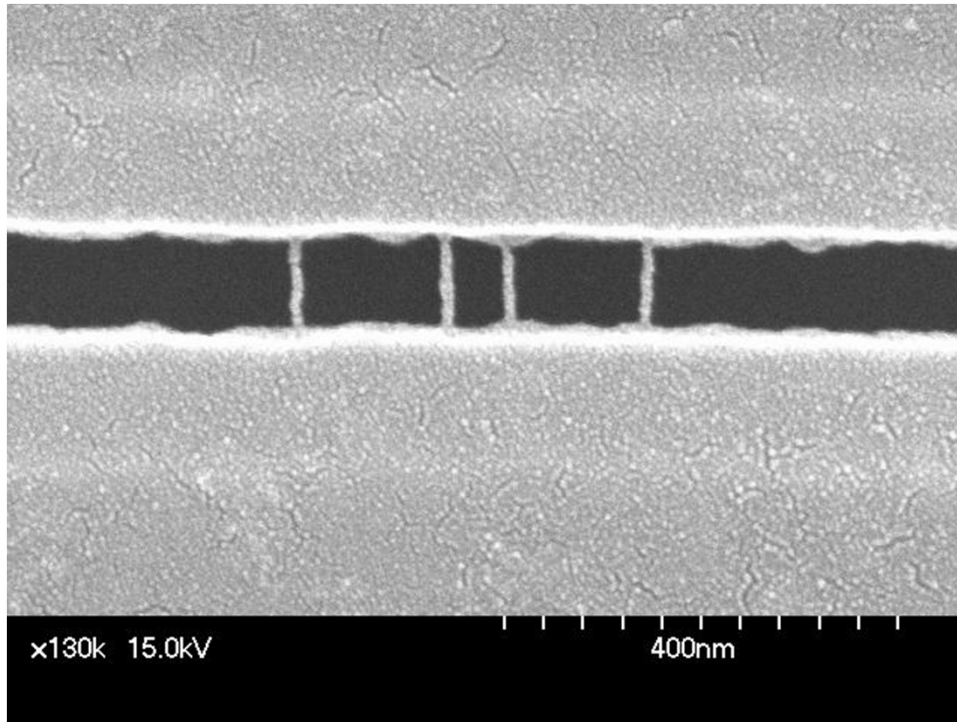
BNL Storage Ring



Features:
Blue/Black circles are part of the physics story
Diagram allows description of components that enter in the data analysis

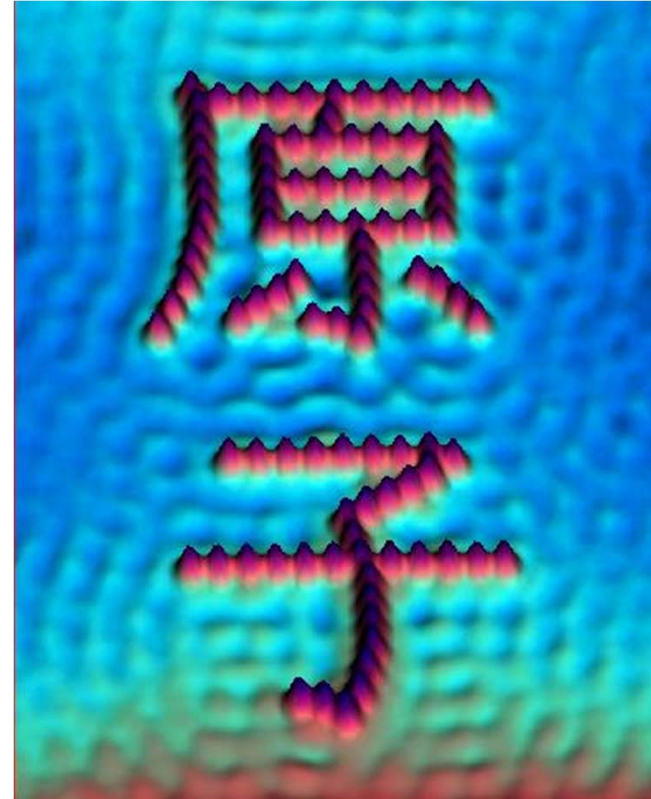
Some more examples of data

A photograph, which reveals the detail



10 nm wires: AuPd on DNA

A photograph, which reveals the detail



Make sure you provide something to show scale, and include a short caption to explain what the audience is looking at

Use equations sparingly

Use equations only when necessary

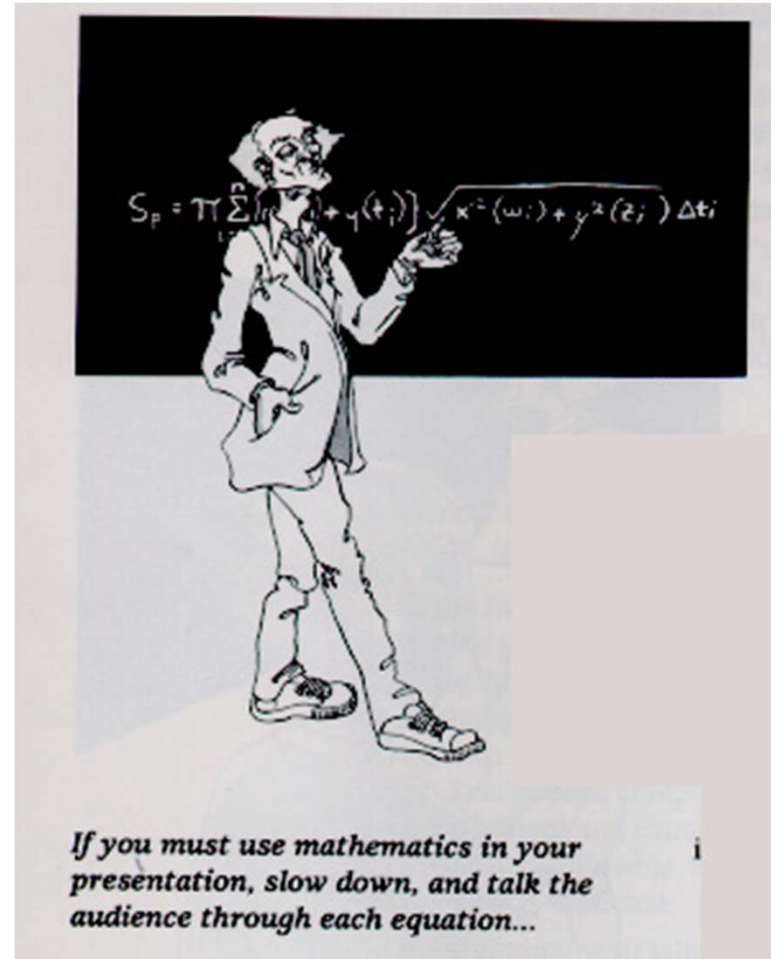
If you use equations

Slow down

Talk through step by step

Explain relevance

Combine with a picture that illustrates the physical principle involved



I think this is a great and effective example of introducing an equation from one of our students

The Radiative Transfer Equation

$$\frac{dI}{ds} = -In(q_a + q_s) + \mathfrak{S}$$

Number of Photons

Density of Dust Grains

Distance Traveled

Absorption Coefficient

Scattering Coefficient

Source Function

(from geometry and composition of dust grains)

The diagram shows the radiative transfer equation $\frac{dI}{ds} = -In(q_a + q_s) + \mathfrak{S}$ with several labels and arrows. 'Number of Photons' has two arrows pointing to 'n' and 'I'. 'Density of Dust Grains' has an arrow pointing to '(q_a + q_s)'. 'Distance Traveled' has an arrow pointing to 'ds'. 'Absorption Coefficient' has an arrow pointing to 'q_a'. 'Scattering Coefficient' has an arrow pointing to 'q_s'. 'Source Function' has an arrow pointing to '\mathfrak{S}'. Below the equation, a note states '(from geometry and composition of dust grains)'.

Requirements to solve analytically:

- n is a constant
- $q_a = 0$ or $q_s = 0$

We want turbulent clouds. n is not a constant

Bad equation example:

Disaster ?

$$\frac{\Delta T}{T} = \frac{T(h\nu, F_o) - T(h\nu, 0)}{T(h\nu, 0)}$$

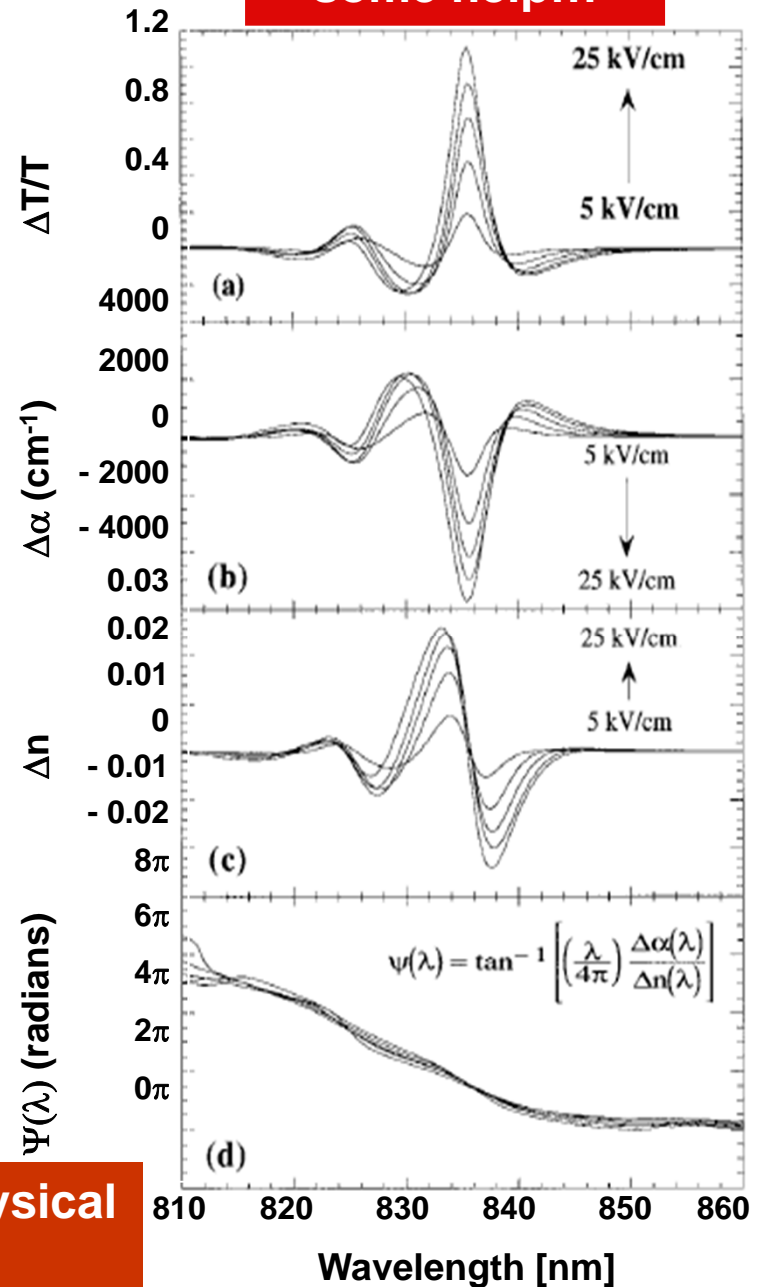
$$\Delta\alpha(h\nu, F_o) = -\frac{1}{L} \ln\left(1 + \frac{\Delta T}{T}\right)$$

$$\Delta n(\lambda) = \frac{\lambda^2}{2\pi^2} P \int_0^\infty \frac{\Delta\alpha(\lambda') d\lambda'}{\lambda^2 - \lambda'^2}$$

$$\psi(\lambda) = \tan^{-1} \left[\left(\frac{\lambda}{4\pi} \right) \frac{\Delta\alpha(\lambda)}{\Delta n(\lambda)} \right]$$

What does this mean? Better to provide a physical interpretation in words next to equations

Data provides some help...



Remember, your goal is to convey your ideas, so avoid distracting text and effects!

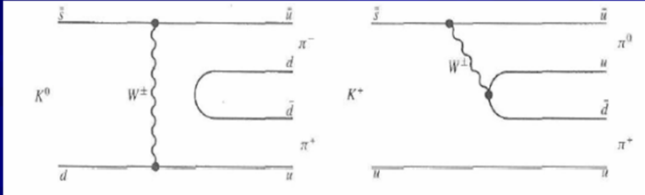
Don't overuse PowerPoint animations and sounds!

Make sure there is good contrast between text and background

Use simple (or no) backgrounds on slides

CP

- Parity invariance fails, combine it with charge conjugation to create a new invariant
- Converts the right-handed anti-neutrino into a left-handed neutrino- exactly what we observe in nature
- Neutral kaon experiment



The image shows two Feynman diagrams illustrating the decay of neutral kaons. The left diagram shows a K^0 meson (quark-antiquark pair $d\bar{u}$) decaying into a W^\pm boson, which then decays into a π^- meson ($d\bar{u}$) and a π^+ meson ($u\bar{d}$). The right diagram shows a K^+ meson (quark-antiquark pair $u\bar{s}$) decaying into a W^\pm boson, which then decays into a π^0 meson ($u\bar{u}$) and a π^+ meson ($u\bar{d}$).



Eschew weird fonts

Don't use *calligraphy*
or serif fonts

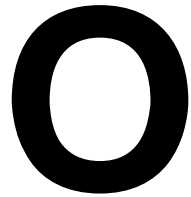
USE THE SAME FONT
THROUGHOUT THE TALK

Make all text at least 20 pt

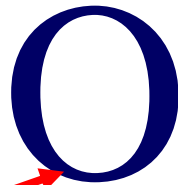


Use San Serif Fonts

Use San Aarif font (e.g., Ariel)



Not Sarif font (e.g., Times New Roman)



Skinny parts disappear when projected

Use “normal” colors

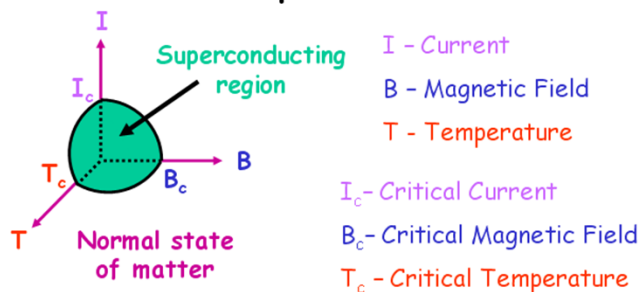
DON'T use red/green or red/blue as contrasting colors

Make sure colors looks the way you expect using an LCD projector!

Avoid neon colors and pastels

Don't use many random colors; people expect color to *mean* something

Superconductivity is an electronic state of matter that exists below certain currents, magnetic fields, and temperatures.



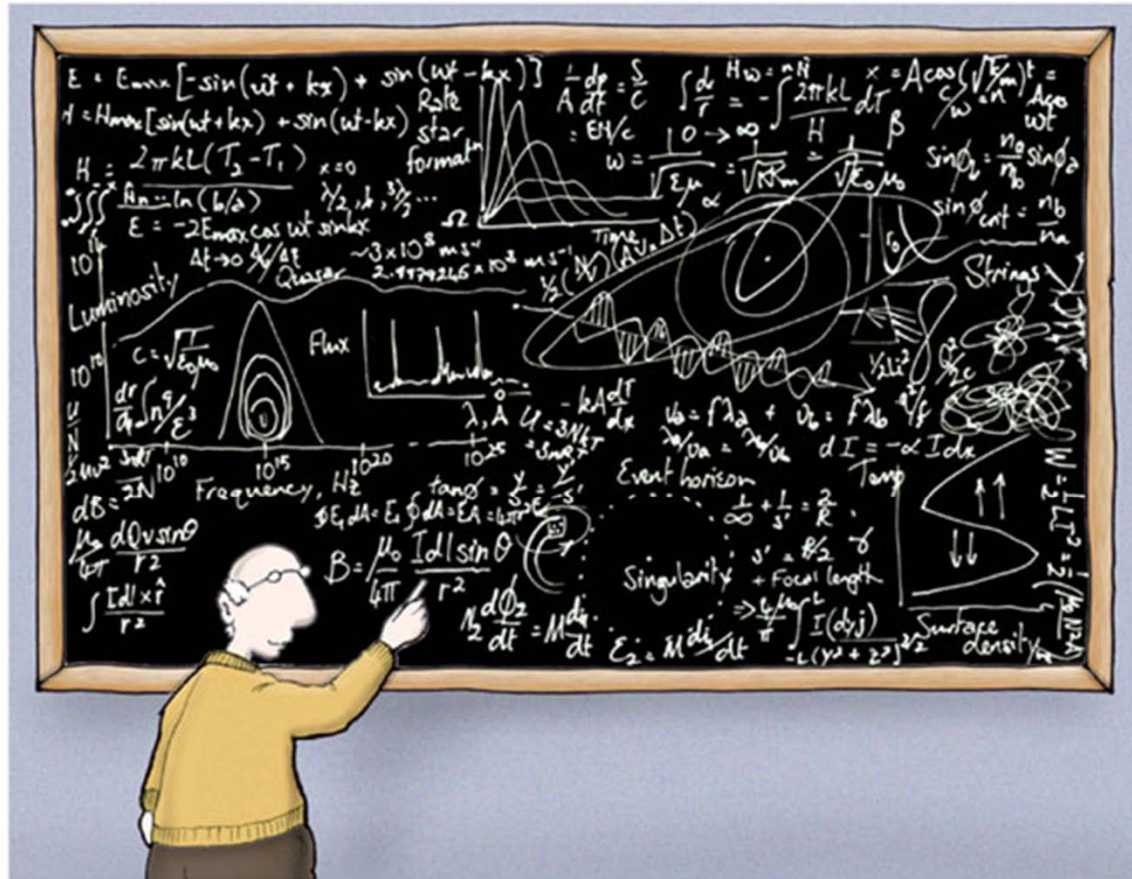
Strive for easy reading

Strive for easy reading

Strive for easy reading



Tips for giving a scientific presentation



Astrophysics made simple



Pointers for giving the best possible talk:

Maintain eye contact with audience

Don't stare at screen or monitor

Do not read your talk!

Avoid nervous mannerisms

Pacing, bobbing, waving arms, jingling coins

Use laser pointer or stick directed at screen

Don't point directly at overhead or projector

Don't block the screen

Train yourself to speak slowly and distinctly— practice!

Avoid “fillers”: “uh”, “like”, “um”, “okay”

Be enthusiastic!

If you don't act excited by your results,
don't expect the audience to be!



Pointers for giving the best possible talk:

Don't show any material on slides (e.g., figures, equations, text, etc.) you can't explain!!

Rehearse how you'll end your talk

Don't end with "Well, I guess
that's it..."

Don't just stop and let the committee guess that you're
done

Thank the audience!



The best way to prepare for a talk is to Know Your Material

Practice, practice, practice

**Focus on communicating,
not performing**

Humor is good, but don't overdo it

Keep it simple

Prepare key phrases

It's okay to write out material first

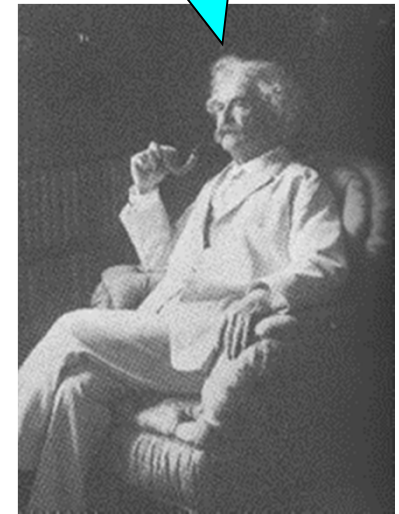
Write the key point to make for each slide

If the slide doesn't have a point, eliminate it!!!

Stay on track

Small (planned) digressions fine if motivated,
but get back on track (shows you are
paying attention to audience)

*It takes three
weeks to prepare
a good ad-lib
speech*



More advice...

Bring a copy of your slides if giving a PowerPoint talk

- this will help you practice
- you can distribute these to interested people

Make appropriate use of the screen:

don't underfill the screen, and don't put key information at the edges of the screen.



Rehearse Your Talk!

A few days before

Practice in front of friends
and check timing
Rehearse likely questions
Solicit feedback about
logic and clarity
Revise (*shorten*)

The night before

Go over one more time
Put all materials *in order*
(number your slides!)



Prof. Per Ahlberg delivering the Presentation Speech for the 2001 Nobel Prize in Chemistry at the Stockholm Concert Hall.

Check *everything* just before your talk

Check the projector

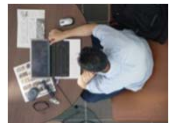
Make sure you know how to turn it on
See that it is plugged in
Check which way to position your slides
Adjust the focus

Check microphones, pointer, other tools

Arrange your slides, notes, and other materials

Be able to reach everything without moving
Be able to go through your slides without fumbling

Have a watch handy to check the time



“Stage Fright”? Be Prepared!

Know your subject thoroughly

Practice in a big room in front of real people

Have all your materials in order

Arrive early

Familiarize yourself with the equipment

Ask a friend to sit in the middle of the audience and speak primarily to him or her

Tell him to look interested and nod frequently

Ask her to smile and nod encouragingly whenever she catches your eye



Non-Native (and Native!) English Speakers:

Do not use slang or 'laboratory' terms

Choose the simplest word

Have a native speaker listen to a rehearsal and review your slides

Speak slowly and distinctly



Handling questions is an essential part of giving a talk

As part of preparing your talk, try to anticipate questions you might get

In each slide, try to identify what the weak points are, what questions you might ask, etc.

Be prepared to repeat simple derivations of equations or estimates presented on your slides

If you don't know the answer?

Say "That's an excellent question. I'm not sure; I'll have to look into it" or "Let's talk about it afterward"



Express your thanks

At the beginning of your talk

Acknowledge colleagues and collaborators who contributed to the work

At the end of the talk

Thank your committee for their attention

