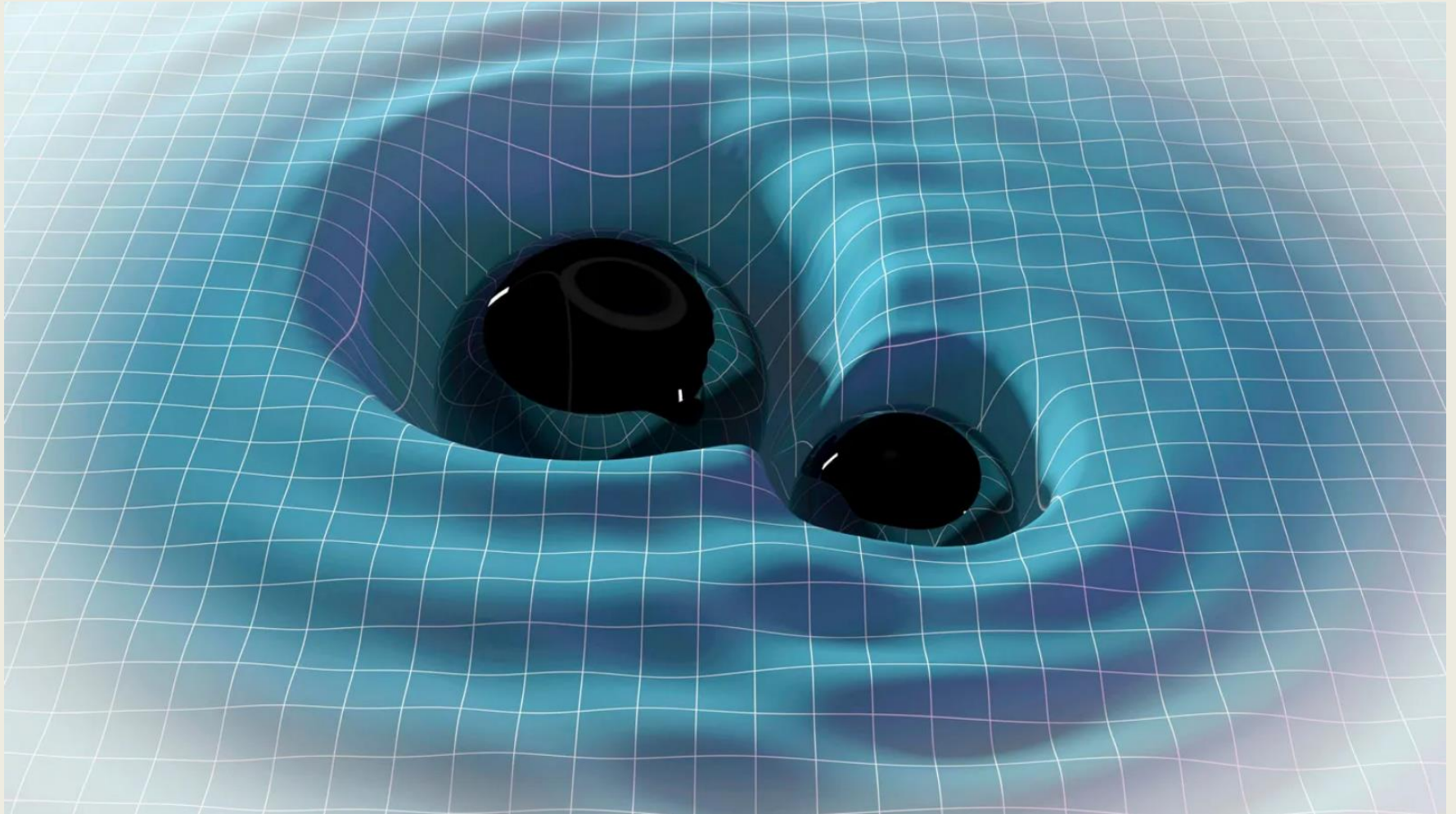


Gravitational Waves



MARK GARLICK/SCIENCE PHOTO LIBRARY/GETTY IMAGES PLUS

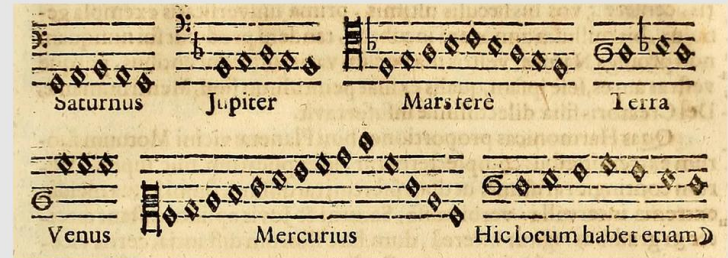
Gravitational Waves



Jorge Luis Borges. *The Library of Babel*. 1941

$$R_{\mu}^{\nu} - \frac{1}{2} \delta_{\mu}^{\nu} R + \Lambda \delta_{\mu}^{\nu} = T_{\mu}^{\nu}$$

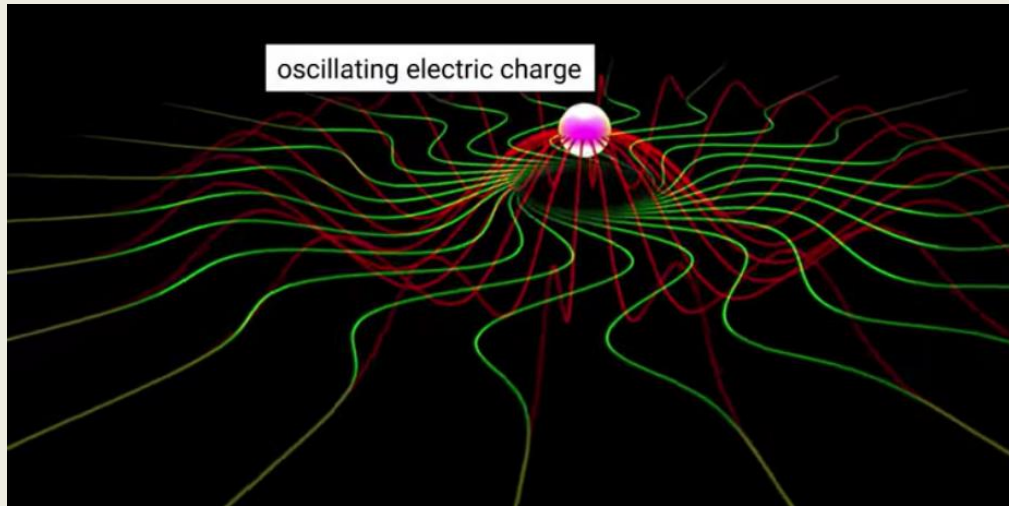
Albert Einstein.
General Relativity Theory.
Gravitational waves. 1916



Celestial harmonics of the planets, from *The Harmony of the World* (1619) by Johannes Kepler, based on the Pythagorean concept of the Music of the Spheres.

Johannes Kepler feat. Pythagoras.
Music of the Spheres. 1619

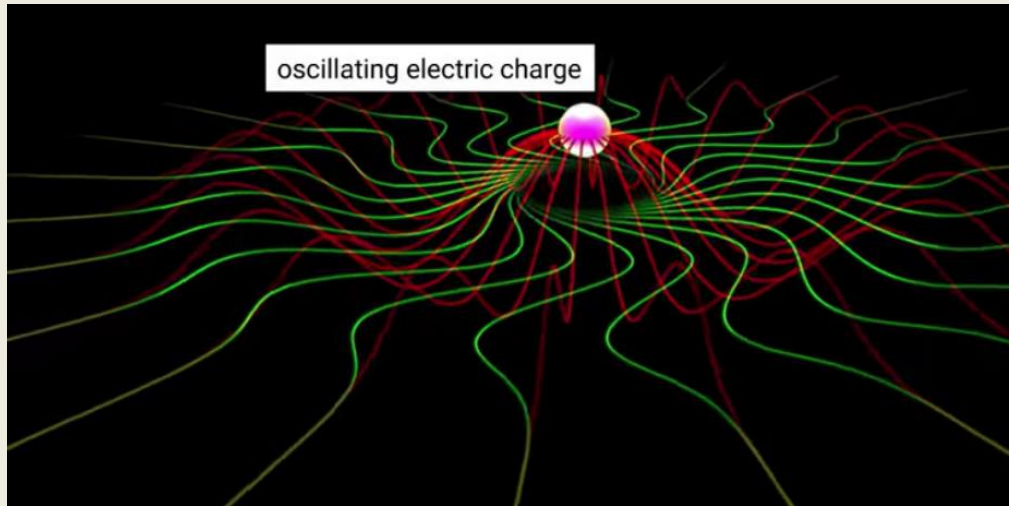
Accelerating charge produces electromagnetic waves



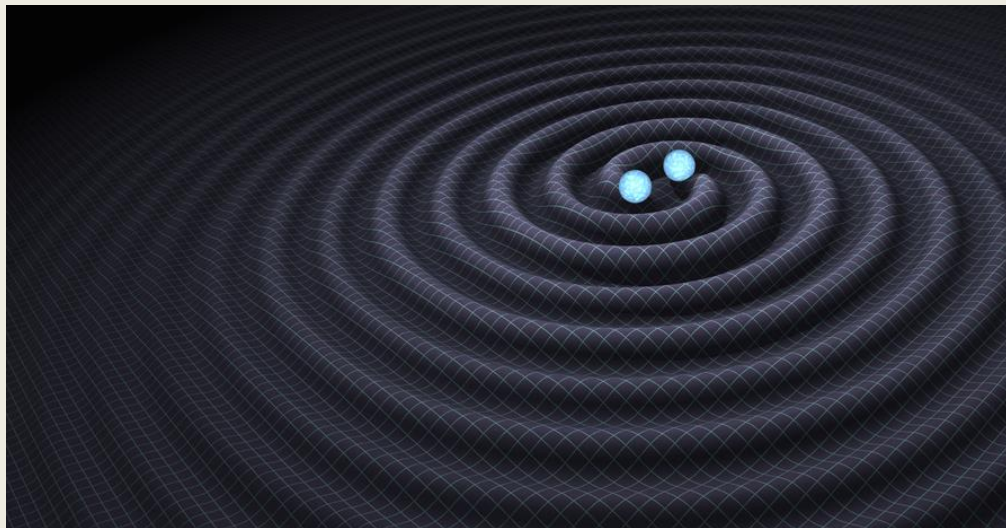
Accelerating mass produces gravitational waves

$$R_{\mu}^{\nu} - \frac{1}{2} \delta_{\mu}^{\nu} R + \Lambda \delta_{\mu}^{\nu} = T_{\mu}^{\nu}$$

Accelerating charge produces electromagnetic waves



Accelerating mass produces gravitational waves

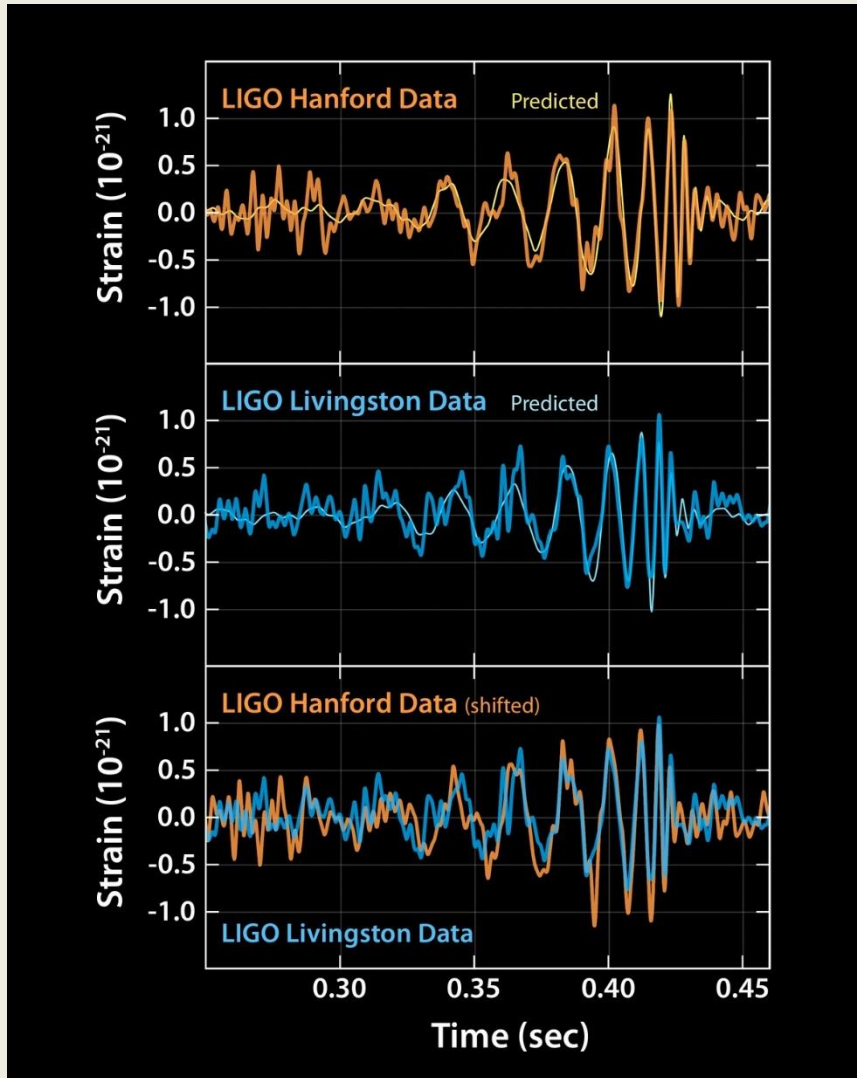


The First Event

September 14, 2015, the **L**aser **I**nterferometer **G**ravitational Wave **O**bservatory (LIGO) made the first direct detection of gravitational waves. These waves were produced by the merger of two black holes about 1.3 billion light-years away.

Strain

$$h \sim \frac{GM}{c^2} \frac{(v/c)^2}{d}$$



Strain

$$h \sim \frac{GM}{c^2} \frac{(v/c)^2}{d}$$

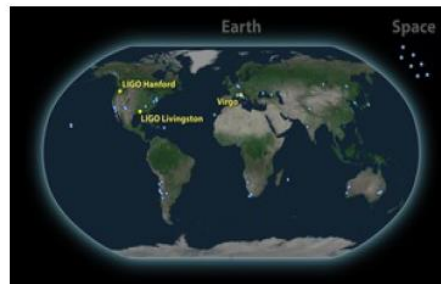
M is the binary's total mass,
 v its velocity near merger and
 d its luminosity distance to Earth

- change in distance: $\frac{\Delta L}{L} \sim \frac{GM}{c^2} \frac{(v/c)^2}{d} \approx 10^{-21}$
- e.g.:
 - $L(\text{London} - \text{NY}) \sim 5500\text{km} \Rightarrow \Delta L \sim \text{proton}$
 - $L(\text{Earth} - \text{Proxima Centauri}) \sim 4.2\text{y} \Rightarrow \Delta L \sim \text{human hair}$
- “tricks:”

really sensitive microphone make microphone enormous keep everything really quiet



(credit: Caltech/MIT/LIGO Laboratory)



detector network



LIGO Livingston

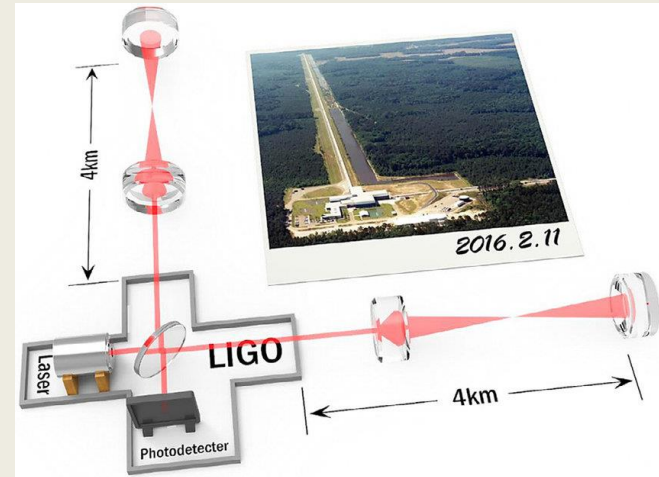
a

$$^a M = 60M_{\odot}, v/c \sim 0.5, d \sim 1.4\text{billionly}; M_{\odot} \sim 2 \cdot 10^{30}\text{kg}; 1\text{y} = 9.4 \cdot 10^{12}\text{km}$$

Credit: Helvi Witek. Lectures for King's College London

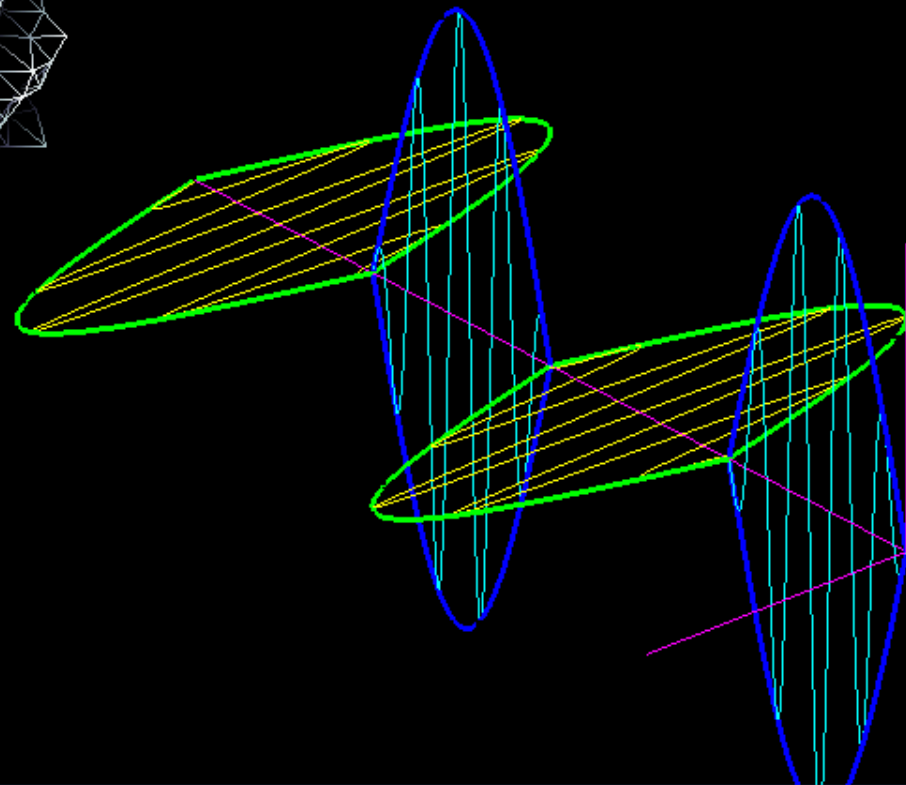
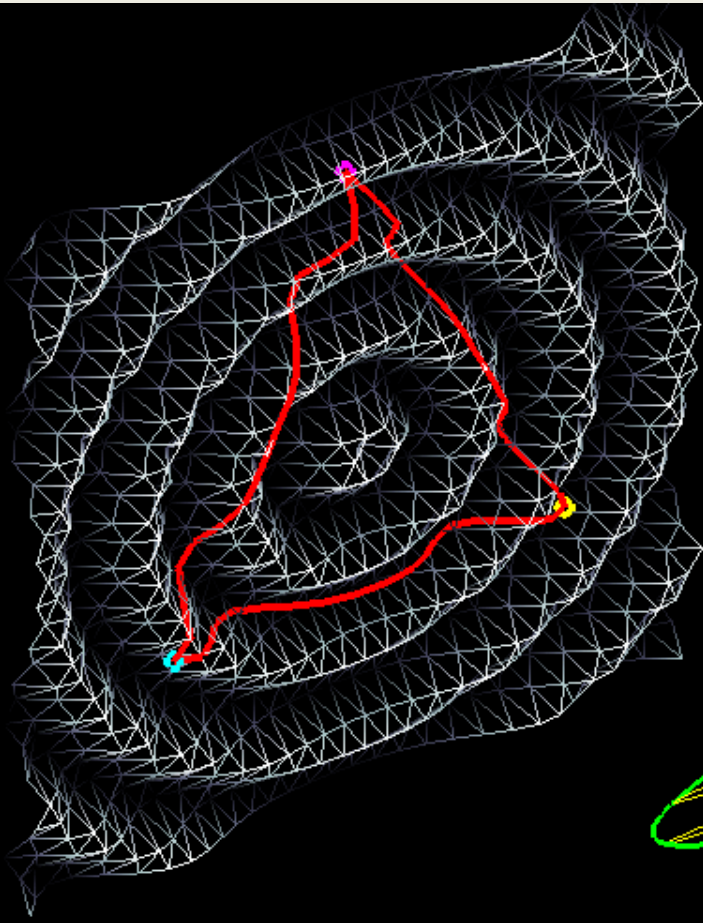
How can we detect it at all?

L
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terferometer
r
a
vitational wave
b
s
ervatory!



<https://www.ligo.org/detections/GW150914.php>

https://youtu.be/tQ_telUb3tE



What can we find from gravitational waves?

Cosmology

- **Primordial Gravitational Waves**
- **Cosmic Microwave Background (CMB) Polarization**
- **Large-Scale Structure of the Universe**
- **Phase Transitions**
- **Cosmic Strings**

Compact Objects

- **Astrophysical Sources**
- **Black Holes**
- **Neutron Stars**
- **Stellar Evolution**
- **Modified Gravity Theories**

BIG BANG

What Powered the Big Bang?

Gravitational Waves can Escape from Earliest Moments of the Big Bang

Inflation
(Big Bang plus 10^{-35} seconds?)

Big Bang plus
300,000 Years

gravitational waves

light

Now

Big Bang plus
15 Billion Years

PGW = Primordial Gravitational Wave

PGW

CMB Polarization



B-mode

E-mode



THE SPECTRUM OF GRAVITATIONAL WAVES



Observatories & experiments

Ground-based experiment



Space-based observatory



Pulsar timing array



Cosmic microwave background polarisation



Timescales

milliseconds

seconds

hours

years

billions of years

Frequency (Hz)

100

1

10^{-2}

10^{-4}

10^{-6}

10^{-8}

10^{-16}

Cosmic fluctuations in the early Universe

Cosmic sources



Supernova



Pulsar



Compact object falling onto a supermassive black hole



Merging supermassive black holes



Merging neutron stars in other galaxies



Merging stellar-mass black holes in other galaxies



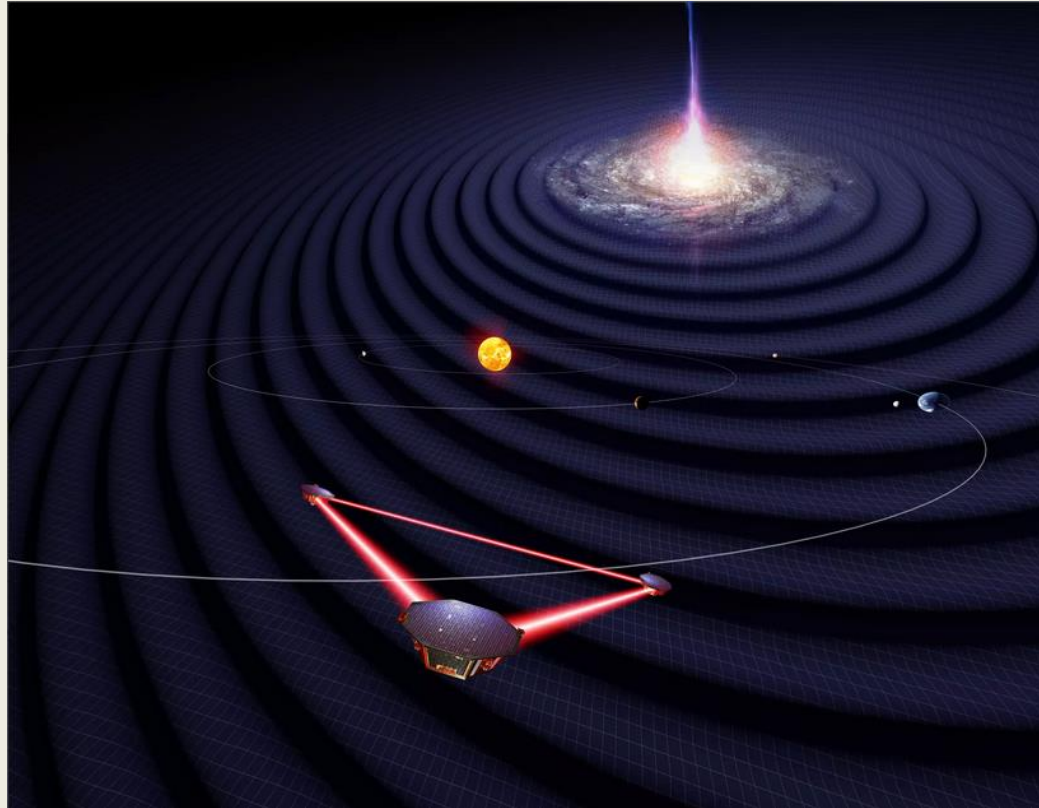
Merging white dwarfs in our Galaxy

#Lisa



Credit: European Space Agency

What's then?



Credit: University of Florida / Simon Barke (CC BY 4.0)

The **L**aser **I**nterferometer **S**pace **A**ntenna (LISA) concept features three spacecraft arranged in an equilateral triangle with each side 2.5 million kilometers long, flying in an Earth-like heliocentric orbit. Launch date 2035 (planned).