













Isaac Newton (1642 - 1727)

Born the year Galileo died

at Woolsthorpe, near Grantham in Lincolnshire, into a poor farming family.

Terrible farmer, sent to Cambridge University in 1661 to become a preacher. Instead, he studied mathematics.

Forced to leave Cambridge from 1665 to 1667 because of the great plague. Newton called this period the "Height of his Creative Power".

Greatest works were accomplished while he was 24- 26 years old!

One of the most influential people who ever lived Newton's Paradigm - now called "classical physics" - dominated Western thought for more than two centuries



Isaac Newton (1642 - 1727) continued

- Suffered a mental breakdown in 1675.
- In 1679 (responding to a letter from Hooke) suggested that particles, when released, would spiral toward the center of the earth. Hooke wrote back claiming the path would be an ellipse.
- Hating to be publicly contradicted, Newton began to work out the mathematics of orbits.
- Urged by Halley to publish his calculations and results, Newton released Principia in 1687. This became one of the most important and influential works on physics of all times

Calculus – Newton vs. Leibnitz

- First known steps ancient Greece
 - Zeno's paradox; Archimedes
- Newton wrote a tract (circulated among mathematicians) in 1666
 - First clear statement of the fundamental theorem of calculus
- Gottfried Wilhelm Leibnitz (1646 1716)
 - From a poor family
 Objied Breading
 - Child Prodigy
 - Famous German Mathematician and Philsopher
 - Invented Calculus 1674-5; published 1684 Controversial whether he had seen Newton's work

Newton's Three Laws

• Inertia:

"Every body continues in its state of rest, or of uniform motion in a right line, unless it is compelled to change that state by a force impressed on it."

• Force, Mass, Acceleration (F=ma):

 "The change in motion [rate of change of momentum] is proportional to the motive force impressed; and is made in the direction of the right line in which that force is impressed."

"Action = Reaction":

 "To every action [change of momentum] there is always opposed an equal reaction; or, the mutual actions of two bodies are always equal, and directed to contrary parts."

Newton's First Law

"Every body continues in its state of rest, or of uniform motion in a right line, unless it is compelled to change that state by a force impressed on it."

Same as Galileo's law of inertia.

- If a body moves with constant velocity in a straight line, then there is NO net Force acting on the body.
- If the body is moving in any other way (i.e. accelerating), then there MUST be a Force acting on the body.

Galilean Relativity revisited:

- "Rest" and "Uniform Motion" really are the same! No net force
 on the object
- As Galileo argued, no experiment in a steadily moving ship will show that is is moving. Only by looking outside can you detect relative motion.





Exercise

- Suppose you are on an airplane travelling at constant velocity with a speed of 500 miles per hour (roughly 200 m/s)
- If you throw a ball straight up, does it return to you?
- How does it appear to you?
- How does the path of the ball look to an observer on the ground?
- Can you think of any experiment done inside the airplane that would detect the motion of the airplane at constant velocity?









Force

- What is a force?
- Force is the tendency to cause acceleration.
- Operationally defined by measuring accelerations.
- Is this just a circular definition?
- No! Forces can be related to physical systems. Compressed springs, gravitational forces, This is the basis for the predictive power of Newton's equations.
- More later on Forces

This is the new idea not present in Galileo's work

Force is a Vector The "Net Force" or "Total Force" on an object is the vector sum of all the forces on it due to other objects This what goes in Newton's Equation Force = mass x acceleration F = m a $\overrightarrow{f_1}$ $\overrightarrow{f_2}$ $\overrightarrow{f_1}$ $\overrightarrow{f_2}$ Net Force F is the vector sum of the three applied forces





Demonstration: Newton's Third Law: Action/Reaction

- Examples of equal and opposite forces
 - Does not matter which body "caused" the force
- · Person pushing on a table
- How does a rocket accelerate?
- Rocket Cart! ---- DEMONSTRATION!
- Note that the total momentum does not change (We will come back to this as an example of a "conservation law" -- momentum is "conserved")



Exercise: Action/Reaction solution

Suppose a tennis ball (m= 0.1 kg) moving at a velocity v = 40 m/sec collides head-on with a truck (M = 500 kg) which is moving with velocity V = 10 m/sec.

· During the collision, the tennis ball exerts a force on the truck which is smaller than the force which the truck exerts on the tennis ball. TRUE or FALSE ? Equal and opposite forces!

The tennis ball will suffer a larger acceleration during the collision than will the truck. Acceleration = Force / mass

· Suppose the tennis ball bounces away from the truck after the collision. How fast is the truck moving after the collision? < 10 m/sec = 10 m/sec > 10 m/sec ?

To conserve total momentum, the truck's speed must decrease since the tennis ball moves in the opposite direction after the collision.

Summary – to this point

Definitions: displacement, velocity, acceleration, momentum are vectors that describe motion

Newton's three laws:

- 1. A body moves with constant velocity unless acted upon by a force -- equivalent to principle of inertia
- 2. F=ma
- 3. Equal and opposite forces -- action/reaction (equivalent to conservation of momentum -- more later)

Concept of Force, Mass

- Mass is a scalar measure of "inertia" or resistance to acceleration
- · Force is a vector tends to cause acceleration
- The force in Newton's equation is the "Net Force" -- the
- vector sum of all forces on a body **Demonstrations of Laws**
- **Curved Motion & Circular Motion** Curved motion is accelerated motion!



















Newton Has Said More than Kepler!

- Kepler's Laws describe the motion of a planet about the Sun.
- Newton uses same laws that apply to all motion!
- Newton's framework (forces & masses) allows him to generalize from the Sun-planet system to all bodies in the universe! This is "universal" gravitation!
- Newton's Third Law implies that each body exerts equal and opposite forces on the other. Depends upon both masses.
 - Describes the moon orbiting the earth
 - The moons of Jupiter, and much more!
 - Totally different from Kepler's approach.

Exercise: Kepler's Laws Suppose you know that the radius of Saturn's orbit is about 9 AU. (the radius of the Earth's orbit = 1AU). • Can you predict the average speed of Saturn in its orbit in terms of the average speed of the Earth in its orbit? • If you can, do it; if you can't, what other information would you need? • Can you predict the acceleration of Saturn in its orbit in terms of the acceleration of the Earth in its orbit? • If you can, do it; if you can't, what other information would you need? • Can you predict the acceleration of Saturn in its orbit in terms of the acceleration of the Earth in its orbit? • If you can, do it; if you can't, what other information would you need? • Can you predict the force that the Sun exerts on Saturn in terms of the force that the Sun exerts on the Earth? • If you can, do it; if you can't, what other information would you need?

The Apple and the Moon I

- Is Newton's Gravitation Force Law really "universal"? Does the same force law describe apples falling to the Earth and the Moon's orbit about the Earth? Can we predict the acceleration due to gravity on the surface of the Earth from the Period & Radius of the Moon's orbit?
 - Acceleration of the moon: $a_{moon} = v^2 / R = 4\pi^2 R / P^2$
 - + If due to gravitation, then also: $a_{moon} = F / m_{moon} = GM_{earth} / R^2$
 - P (period of moon = 1 lunar month), R (distance to the moon) , G (unversal constant measured much later)

We can determine the mass of the earth!

If we know

But Newton Did not know the value of G!



Effects of gravity

- Seen everywhere around us
 Falling objects
- Planets, Moons orbiting larger bodies
- Double star systems rotating around each
 other
- Galaxies millions of stars clustered due to gravitational forces
- See Feyman, Chapter 5



Gravity is a VERY Tiny force

• Force between two objects each 1 Kg at a distance of 1 meter is

 $\mathbf{F} = \mathbf{G} \, \mathbf{M}_1 \, \mathbf{M}_2 / \mathbf{R}^2$

= 6.67 x 10⁻¹¹ N

- 1 N is about the weight of one apple on the earth
- The reason the effects of gravity are so large is that the masses of the earth, sun, stars, are so large -- and gravity extends so far in space

Additional Comments

- Newton's Theory of gravitation contains one deeply unsatisfying aspect
 Newton recognized the problem
- · ·
- The law f = G M m /r² means "action at a distance"
- Instantaneous force due on one object due to another object no matter how far they are away from one another
- What should a scientist do?

Summary Newton's 3 laws Circular Motion Centripetal (toward center) accel. a = v²/r Centripetal force f = ma Example: Ball on a string moving in a circle Kepler's Laws explained by gravitational force in Newton's laws Universal law of gravitation: f = G M m /r² The falling Apple and the Moon: each accelerates toward the earth obeying the same law! Motion on Earth and in the heavens obeying the same simple laws! Force of gravity – extremely weak – large effects for large objects

Next Time

- Conservation Laws
- MORE important than Newton's Equations! still valid in modern physics even though Newton's laws are not !
- The most useful conclusions without solving any equations!
 - Conservation of momentum: Follows from Newton's third law. (Chapt. 2 in Text)
 - Conservation of energy: The most important and useful law. (Chapt. 5 in Text, Chapter 4 in Feynman)



