Last Name: $\qquad$ First Name: $\qquad$ Net ID: $\qquad$
Discussion Section: $\qquad$ Discussion TA: $\qquad$

## Academic Integrity:

The following actions or activities during a University Examination are grounds for disciplinary action up to and including expulsion:

- Giving assistance to or receiving assistance from another student.
- Using unauthorized materials, including unauthorized electronic devices (phones, tablets, smart watches, laptops, etc.).

Turn off and put away all internet-capable electronic devices.

## Calculators may not be shared.

## Exam Duration:

You have 1.5 hours ( 90 minutes) to complete this examination.

## Examination Type:

This is a Midterm Exam.

This is a closed-book examination. Put away all notes, textbooks, and study aids.

## Instructions:

## Use only a \#2 (HB) pencil to fill in the answer sheet.

1. Name:
a. Print your Last Name in the designated spaces on the answer sheet.
b. Fill in the corresponding circles below each character in your Last Name.
c. Print your First Initial in the designated spaces on the answer sheet.
d. Fill in the corresponding circle below you First Initial.
2. NetID:
a. Print your NetID in the designated spaces on the answer sheet.
b. Fill in the corresponding circles below each character in your NetID.
3. Your UIN:
a. Print you UIN in the STUDENT NUMBER spaces on your answer sheet.
b. Fill in the corresponding circle below each number in your UIN.
4. Sign your full name on THE STUDENT SIGNATURE line on the answer sheet.
5. Print your Discussion Section on the SECTION line of your answer sheet. Do not fill in the SECTION box.
6. Your Exam Test Form A. Mark the circle for $\mathbf{A}$ in the TEST FORM box on your answer sheet.
7. Your exam should have $\mathbf{1 2}$ exam pages and $\mathbf{2}$ formula sheets. Count the number of pages in your exam booklet now.

## Grading Policy:

This exam is worth 114 points. There are
three types of questions:

- MC5: multiple-choice-five-answer questions, each worth 6 points.

Credit will be granted as follows:
(a) If you mark only one answer and it is the correct answer, you earn 6 points.
(b) If you mark two answers, one of which is the correct answer, you earn $\mathbf{3}$ points.
(c) If you mark three answers, one of which is the correct answer, you earn $\mathbf{2}$ points.
(d) If you mark no answers, or more than three, you earn $\mathbf{0}$ points.

- MC3: multiple-choice-three-answer questions, each worth 3 points.

No partial credit. Mark only one answer
(a) If your marked answer is the correct answer, you earn $\mathbf{3}$ points.
(b) If your marked answer is wrong answer or if you mark no answers, you earn $\mathbf{0}$ points.

- TF: true-false questions, each worth 2 points.

No partial credit. Mark only one answer.
(a) If your marked answer is the correct answer, you earn $\mathbf{2}$ points.
(b) If your marked answer is wrong or you marked neither answer, you earn 0 points.

Two identical loudspeakers produce sound of equal intensity and frequency $=1200 \mathrm{~Hz}$. The sound waves travel at a speed of $340 \mathrm{~m} / \mathrm{s}$. The speakers are driven in phase and are located at $\mathrm{x}=-5 \mathrm{~m}$ and $\mathrm{x}=+5 \mathrm{~m}$. A listener at $\mathrm{x}=0$ hears a total sound intensity $\mathrm{I}_{\text {Total }}$ If the listener moves to the left by a distance $d$, further from one speaker and closer to the other, the total intensity that she hears drops to $\mathrm{I}_{\text {Total }} / 2$.


- 5 m
+5 m

1) What is the minimum value of $d$ ?
a. 0.567 m
b. 0.142 m
c. 0.283 m
d. 0.0708 m
e. 0.0354 m

Microwaves of frequency $5 \times 10^{10} \mathrm{~Hz}$ are incident on a circular hole that is located 20 meters from a wall, as shown below. At the wall, the microwave intensity is zero at a radius of 1 m from the central maximum.

2) What is the diameter of the hole?
a. 0.29 m
b. 0.71 m
c. 0.52 m
d. 0.12 m
e. 0.15 m

Six loudspeakers are located on a circle. Assume that the sound wave from each speaker has a $\pi / 2$ phase shift (in radians) relative to the previous one. That is, $\phi_{1}=0, \phi_{2}=\pi / 2, \phi_{3}=\pi$, etc. Speakers 1-5 each separately generate sound at the center of the circle with intensity of $1 \mathrm{~W} / \mathrm{m}^{2}$. Speaker 6 generates a sound intensity of $5 \mathrm{~W} / \mathrm{m}^{2}$.

3) What is the total sound intensity at the center of the circle in $\mathrm{W} / \mathrm{m}^{2}$ ? Hint: Use phasors.
a. 10
b. 2.5
c. 10
d. 0
e. 6

A water-wave disturbance travelling along the surface of a lake takes the form, $h(x, t)=0.2 \exp \left[-(2 x+5 t)^{2} / 3\right]$. Here, $x$ has units of meters and $t$ has units of seconds. Suppose, instead, that a generator produces sinusoidal water waves of frequency 8 Hz on the same lake.
4) What would be the wavelength of these water waves?
a. 0.438 m
b. 0.312 m
c. 0.625 m
d. 0.5 m
e. 0.104 m

## The next three questions pertain to the situation described below.

Consider the wavefunction shown below, relating to an excited eigenstate of an electron in an infinite 1D quantum well. The 1D well has a length $\mathrm{L}=0.5 \mathrm{~nm}$.

5) If the electron decays to a lower energy level, what is the longest wavelength of photon that may be emitted?
a. 206 nm
b. 91.6 nm
c. 103 nm
d. 165 nm
e. 26.2 nm
6) Imagine that particle in the well is a muon instead of an electron, where a muon has the same charge as an electron but a rest mass that is 207 times greater. Which of the following relationships would describe the relative energies of the electron's second-
excited state $\left(\mathrm{E}_{\mathrm{e}}{ }^{\mathrm{n}=3}\right)$ and the muon's second-excited state $\left(\mathrm{E}_{\mathrm{mu}}{ }^{\mathrm{n}=3}\right)$ ?
a. $\mathrm{Eel}^{\mathrm{n}=3}>\mathrm{E}_{\mathrm{mu}}{ }^{\mathrm{n}=3}$
b. $\mathrm{E}_{\text {el }}{ }^{\mathrm{n}=3}<\mathrm{E}_{\mathrm{mu}}{ }^{\mathrm{n}=3}$
c. $\mathrm{E}_{\mathrm{el}}{ }^{\mathrm{n}=3}=\mathrm{E}_{\mathrm{mu}}{ }^{\mathrm{n}=3}$
7) Consider now a muon in the ground state $(\mathrm{n}=1)$. A measurement of its position determines that it is within the region between $x=0$ and $x=0.05 \mathrm{~nm}$. If you subsequently performed a
measurement of the muon's speed, what is roughly (to an order of magnitude) the maximum value that you would expect to observe?
a. $\sim 10^{6} \mathrm{~m} / \mathrm{s}$
b. $\sim 0 \mathrm{~m} / \mathrm{s}$
c. $\sim 10^{4} \mathrm{~m} / \mathrm{s}$

Consider the quantum state wavefunctions shown below (labeled \#1-\#5).
Note that the first 4 relate to energy eigenstates, while the final plot is for a superposition of energy eigenstates.

Also note that the axes (both horizontal and vertical) of the various plots are not necessarily scaled the same.

8) What is the correct ordering of the above wavefunctions, according to their probability density $P(0)$ to be found at the point $\mathrm{x}=0$ ?
a. $\mathrm{P}_{5}(0)>\mathrm{P}_{1}(0)=\mathrm{P}_{2}(0)>\mathrm{P}_{4}(0)>\mathrm{P}_{3}(0)$
b. $\mathrm{P}_{2}(0)>\mathrm{P}_{1}(0)>\mathrm{P}_{4}(0)>\mathrm{P}_{5}(0)>\mathrm{P}_{3}(0)$
c. $\mathrm{P}_{5}(0)>\mathrm{P}_{2}(0)>\mathrm{P}_{3}(0)>\mathrm{P}_{4}(0)>\mathrm{P}_{1}(0)$

Consider the quantum eigenstate wavefunctions shown below (labeled \#1-\#5).
Note that the axes (both horizontal and vertical) of the various plots are not necessarily scaled the same.




(4)

9) What is the correct ordering of the above wavefunctions according to their energy E?
a. $\mathrm{E}_{1}>\mathrm{E}_{3}>\mathrm{E}_{4}>\mathrm{E}_{5}>\mathrm{E}_{2}$
b. $\mathrm{E}_{3}>\mathrm{E}_{4}>\mathrm{E}_{5}>\mathrm{E}_{2}>\mathrm{E}_{1}$
c. $\mathrm{E}_{5}>\mathrm{E}_{3}>\mathrm{E}_{1}>\mathrm{E}_{4}>\mathrm{E}_{2}$
d. $\mathrm{E}_{1}>\mathrm{E}_{3}>\mathrm{E}_{4}>\mathrm{E}_{2}>\mathrm{E}_{5}$
e. $\mathrm{E}_{3}>\mathrm{E}_{1}>\mathrm{E}_{4}>\mathrm{E}_{2}>\mathrm{E}_{5}$

Consider a particle in the 5th excited state $(\mathrm{n}=6)$ of the potential well shown below. The numbering is such that $\mathrm{n}=1$ is the ground state.

10) Which of the following best represents the probability distribution of the particle? (The horizontal axis corresponds to probabilty $=0$.)

a

b

C

d

e
a. .
b. .
c. .
d. .
e. .

## The next three questions pertain to the situation described below.

A screen with multiple slits is fully illuminated by a red laser (wavelength 700 nm ) and the following interference pattern is observed on a screen 3 m away.

11) How many slits are there on the screen?
a. 4
b. 5
c. 3
12) Consider now the second principal maximum of the intensity. At the point of this maximum, what is the path difference for the waves coming from the first slit and the third slit of the diffraction grating?
Assume the wave length is $\lambda$.
a. $4 \lambda$
b. $\lambda / 2$
c. $2 \lambda$
13) Suppose now the distance between the zero principal maximum and the first principal maximum on the screen is 0.001 m . Calculate the distance between the slits of the diffraction grating.
a. 0.0021 m
b. 0.0105 m
c. 0.042 m
d. 0.021 m
e. $7 \times 10^{-4} \mathrm{~m}$

## The next three questions pertain to the situation described below.

Traveling to interstellar space is difficult, in large part because it's very expensive to carry along all the necessary fuel. An alternative is to put an optical reflector on the back of the space ship, and then push it using laser beams from the earth. Initially, scientists (including Steven Hawking) envision using very small "nanocraft", each the size of a postage stamp (mass = 1 gram) and each equipped with a thin reflective circular sail that is 4 m in diameter. The purpose of the sail is to reflect all of the collected photons from the lasers back to the source, which doubles the momentum transfer to the space ship.
14) The original laser beam will be very large ( $\sim 1 \mathrm{~km}$ in diameter). Initially neglecting spreading of the beam (valid when a nanocraft is close to earth), estimate the acceleration that the nanocraft will feel, assuming the total laser power over the entire 1-km diameter is a sustained 100 GigaWatts ( 1 GigaWatt $=10^{9} \mathrm{~W}$ ). Don't forget to account for the finite size of the sail, or the fact that it is $100 \%$ reflective.
a. $330,000 \mathrm{~m} / \mathrm{s}^{\wedge} 2$
b. $10.7 \mathrm{~m} / \mathrm{s}^{\wedge} 2$
c. $3.2 \mathrm{~m} / \mathrm{s}^{\wedge} 2$
d. $0.23 \mathrm{~m} / \mathrm{s}^{\wedge} 2$
e. $670,000 \mathrm{~m} / \mathrm{s}^{\wedge} 2$
15) Of course when the ship gets far enough away, the diffraction of the beam will reduce the intensity of the laser radiation hitting the sail. Which of the following would reduce the diffraction spreading of the beam?

1. Use a shorter wavelength for the 100 GW laser.
2. Use a longer wavelength for the 100 GW laser.
3. Decrease the size of the laser beam at the source.
a. 2 and 3 would both work
b. 2
c. 3
d. 1 and 3 would both work
e. 1
16) Assuming that the wavelength of the laser beam is 400 nm and that the "sail" is made of gold (which has a work function of 5.1 eV ) what is the kinetic energy of electrons that may be released from the gold by the photoelectric effect?
a. 1.0 eV
b. No electrons are released.
c. 3.1 eV

The next three questions pertain to the situation described below.
A beam of electrons (initially nearly at rest) is accelerated through an electostatic potential of 5000 V , resulting in electrons that have a kinetic energy of $5.0 \pm 0.1 \mathrm{keV}$. These electrons are then directed onto a nanofabricated structure with two small slits separated by a distance of 10 nm . An imaging detector is placed 2 m away from the slits.
17) Assuming that only one electron passes through the slits at a time, what is the separation of adjacent interference maxima on the screen?
a. 5 cm
b. There is no interference if only one electron passes the slits at a time.
c. 3.5 mm
18) A beam of photons is now directed onto the same pair of slits and generates the same interference probability distribution interference pattern for the arrival probability distribution on the screen. What is the energy of the photons?
a. 24.5 eV
b. 71.5 keV
c. 5 keV
19) The top slit is now slightly charged to slow the electrons going through that slit to $90 \%$ of the speed of the electrons going through the other slit. What will happen to the overall pattern of the electrons on the screen, assuming the total number of electrons reaching the screen from each slit is not changed?



a. Figure c: There will no longer be interference fringes, but the overall envelope will be slightly wider.
b. Figure b: The entire pattern will shift up slightly, i.e., the central principal maximum will occur at a (small) positive angle; the top of the screen corresponds to the right side of the diagram.
c. Figure a: The entire pattern will shift down slightly, i.e., the central principal maximum will occur at a (small) negative angle; the top of the screen corresponds to the right side of the diagram.


The Cosmic Microwave Background (CMB) is a primordial glow of electromagnetic radiation, which comes to us with almost equal intensity from every direction in the sky. The CMB shows a characteristic pattern of very faint bright and dim patches, with a typical patch separation of one degree. We would like to build a telescope that can resolve these bright and dim patches, i.e. distinguish two sources in the sky separated by 1 degree.
20) Suppose that we build our telescope to measure the CMB at a frequency of 120 GHz . What diameter must the telescope's (circular) aperture have in order to just barely resolve these features?
a. 0.0025 m
b. 5.73 m
c. 0.175 m
d. 0.00205 m
e. 0.349 m
21) Suppose that this telescope collects an average power flow of 0.15 pW , again at a frequency of 120 GHz . On average, how many photons does it collect each second?
a. 800 per second
b. $3.1 \times 10^{19}$ per second
c. $3.1 \times 10^{17}$ per second
d. $1.9 \times 10^{21}$ per second
e. $1.9 \times 10^{9}$ per second

The next two questions pertain to the situation described below.
A spectrometer is used to analyze the light coming from a distant source. The spectrometer has a $9-\mathrm{mm}$ wide diffraction grating with $3 \times 10^{3}$ lines. Light is normally incident on the grating.
22) The third-order principle maximum is observed at an angle of $45^{\circ}$ from normal incidence. What is the wavelength of the light coming from the source? Assume the entire grating is illuminated.
a. 109 nm
b. 1014 nm
c. 707 nm
d. 105 nm
e. 100 nm
23) Light from two distant sources now reaches the spectrometer. The stationary source emits light of wavelength 595 nm . Due to the Doppler shift, the receding source emits a longer wavelength, $595+\Delta \lambda$ nm . What is the smallest value of $\Delta \lambda$ that can be detected by this spectrometer, now assuming that only half the number of lines in the grating are illuminated? (Hint: what is the highest order that can be used?)

a. 0.32 nm
b. 0.4 nm
c. 0.08 nm

## The next two questions pertain to the situation described below.

The LIGO interferometer is used to detect the very very small compression and expansion of space as a gravity wave passes by. It is essentially a Michelson interferometer, each of whose arms $L_{1}$ and $L_{2}$ is 4 km long. The input is a $200-\mathrm{W}$ laser at a wavelength of 1064 nm . Assume that at the start the interferometer arms are adjusted so that all the light exits to the detector shown.

24) $\mathrm{L}_{1}=\mathrm{L}_{2}$ gives the full 200 W coming to the detector. What is the minimum increase in $\mathrm{L}_{2}$ required so that only 100 W goes to the detector?
a. 266 nm
b. 133 nm
c. 532 nm
25) Assume that $L_{2}$ has been increased by the amount calculated in the previous question, so 100 W is reaching the detector. The LIGO interferometer arms are kept under high vacuum. If a small amount of air now leaks into the system (filling both arms equally), what will initially happen to the power on the detector? Neglect any absorption of the light by the air.
a. The power will increase slightly.
b. The power will stay the same.
c. The power will decrease slightly.

## Physics 214 Common Formulae

| SI Prefixes |  |  |
| :--- | :--- | :--- |
| Power | Prefix | Symbol |
| $10^{9}$ | giga | G |
| $10^{6}$ | mega | M |
| $10^{3}$ | kilo | k |
| $10^{0}$ |  |  |
| $10^{-3}$ | milli | m |
| $10^{-6}$ | micro | $\mu$ |
| $10^{-9}$ | nano | n |
| $10^{-12}$ | pico | p |


| Physical Data and Conversion Constants |  |
| :---: | :---: |
| speed of light | $\mathrm{c}=2.998 \times 10^{8} \mathrm{~m} / \mathrm{s}$ |
| Planck constant | $\begin{aligned} \mathrm{h} & =6.626 \times 10^{-34} \mathrm{~J} \cdot \mathrm{~s} \\ & =4.135 \times 10^{-15} \mathrm{eV} \cdot \mathrm{~s} \end{aligned}$ |
| Planck constant / $2 \pi$ | $\begin{aligned} \hbar & =1.054 \times 10^{-34} \mathrm{~J} \cdot \mathrm{~s} \\ & =0.658 \times 10^{-15} \mathrm{eV} \cdot \mathrm{~s} \end{aligned}$ |
| electron charge | $\mathrm{e}=1.602 \times 10^{-19} \mathrm{C}$ |
| energy conversion | $1 \mathrm{eV}=1.602 \times 10^{-19} \mathrm{~J}$ |
| conversion constant | $h c=1240 \mathrm{eV} \cdot \mathrm{nm}=1.986 \times 10^{-25} \mathrm{~J}-\mathrm{m}$ |
| useful combination | $\mathrm{h}^{2} / 2 \mathrm{~m}_{\mathrm{e}}=1.505 \mathrm{eV} \mathrm{nm}^{2}$ |
| Bohr radius | $a_{o}=\left(4 \pi \varepsilon_{o}\right) \hbar^{2} / m_{e} e^{2}=0.05292 \mathrm{~nm}$ |
| Rydberg energy | $h c R_{\infty}=m_{e} e^{4} / 2\left(4 \pi \varepsilon_{o}\right)^{2} \hbar^{2}=13.606 \mathrm{eV}$ |
| Coulomb constant | $\kappa=1 /\left(4 \pi \varepsilon_{o}\right)=8.99 \times 10^{9} \mathrm{~N} \cdot \mathrm{~m}^{2} / \mathrm{C}^{2}$ |
| Avagadro constant | $\mathrm{N}_{\mathrm{A}}=6.022 \times 10^{23} / \mathrm{mole}$ |
| electron mass | $\mathrm{m}_{\mathrm{e}}=9.109 \times 10^{-31} \mathrm{~kg}=0.511 \mathrm{MeV} / \mathrm{c}^{2}$ |
| proton mass | $\mathrm{m}_{\mathrm{p}}=1.673 \times 10^{-27} \mathrm{~kg}=938.3 \mathrm{MeV} / \mathrm{c}^{2}$ |
| neutron mass | $\mathrm{m}_{\mathrm{n}}=1.675 \times 10^{-27} \mathrm{~kg}=939.6 \mathrm{MeV} / \mathrm{c}^{2}$ |
| hydrogen atom mass | $\mathrm{m}_{\mathrm{H}}=1.674 \times 10^{-27} \mathrm{~kg}$ |
| Electron magnetic moment | $\begin{aligned} \mu_{\varepsilon} & =9.2848 \times 10^{-24} \mathrm{~J} / \mathrm{T} \\ & =5.795 \times 10^{-5} \mathrm{eV} / \mathrm{T} \end{aligned}$ |
| Proton magnetic moment | $\begin{aligned} \mu_{\mathrm{p}} & =1.4106 \times 10^{-26} \mathrm{~J} / \mathrm{T} \\ & =8.804 \times 10^{-8} \mathrm{eV} / \mathrm{T} \end{aligned}$ |

## Trigonometric identities

$$
\begin{gathered}
\sin ^{2} \theta+\cos ^{2} \theta=1 \\
\cos \theta+\cos \phi=2 \cos \left(\frac{\theta+\phi}{2}\right) \cos \left(\frac{\theta-\phi}{2}\right) \\
\sin \theta+\sin \phi=2 \sin \left(\frac{\theta+\phi}{2}\right) \cos \left(\frac{\theta-\phi}{2}\right)
\end{gathered}
$$

$$
\cos (\theta+\phi)=\cos \theta \cos \phi-\sin \theta \sin \phi
$$

$$
\sin (\theta+\phi)=\sin \theta \cos \phi+\cos \theta \sin \phi
$$

$A_{1} \sin \left(\omega t+\phi_{1}\right)+A_{2} \sin \left(\omega t+\phi_{2}\right)=A_{3} \sin \left(\omega t+\phi_{3}\right)$
$A^{2}+B^{2}+2 A B \cos \phi=C^{2}(\phi$ here is the external angle $)$

## Waves, Superposition

$$
k \equiv \frac{2 \pi}{\lambda} \quad \omega \equiv 2 \pi f \quad T \equiv \frac{1}{f} \quad v=\lambda f=\frac{\omega}{k}
$$

General relation for I and A: $I \propto A^{2}, A=A_{1}+A_{2}+\ldots$
Two sources: $\mathrm{I}_{\text {max }}=\left|\mathrm{A}_{1}+\mathrm{A}_{2}\right|^{2}, \mathrm{I}_{\text {min }}=\left|\mathrm{A}_{1}-\mathrm{A}_{2}\right|^{2}$
Two sources, same $\mathrm{I}_{1}: I=4 I_{1} \cos ^{2}(\phi / 2)$ where

$$
\phi=2 \pi \delta / \lambda
$$

## Interference: Slits, holes, etc.

Far-field path-length difference: $\delta \equiv r_{1}-r_{2} \approx d \sin \theta$
Phase difference: $\frac{\phi}{2 \pi} \equiv \frac{\delta}{\lambda}=\frac{d \sin \theta}{\lambda} \approx \frac{d \theta}{\lambda} \approx \frac{d}{\lambda} \frac{y}{L}$ if $\theta$ small
Principal maxima: $d \sin \theta_{\max }= \pm m \lambda \quad m=0,1,2, \ldots$
N slit: $I_{N}=I_{1}\left\{\frac{\sin (N \phi / 2)}{\sin (\phi / 2)}\right\}^{2}$ where $\phi=2 \pi d \sin \theta / \lambda$
Single slit: $\delta_{a}=a \sin \theta \quad a \sin \theta_{\min }= \pm m \lambda \quad$ with $m=1,2,3 \ldots$

$$
\frac{\beta}{2 \pi} \equiv \frac{\delta_{a}}{\lambda}=\frac{a \sin \theta}{\lambda} \approx \frac{a \theta}{\lambda} \approx \frac{a}{\lambda} \frac{y}{L}
$$

Single slit: $I_{1}=I_{0}\left\{\frac{\sin (\beta / 2)}{\beta / 2}\right\}^{2}$ with $\beta=2 \pi a \sin \theta / \lambda$
slit: $\theta_{0} \approx \lambda / a$ or hole: $\theta_{0} \approx 1.22 \lambda / D \approx \alpha_{c}$
Approx. grating resolution: $\frac{\Delta \lambda}{\lambda} \geq \frac{1}{N m}$

| Quantum laws, facts.... |  |
| :--- | :---: |
| UNIVERSAL: $p=\hbar k=h / \lambda \quad E=h f=\hbar \omega$ |  |
| Light: $E=h f=\hbar \omega=h c / \lambda=p c$ |  |
| Slow particle: $K E=m v^{2} / 2=p^{2} / 2 m=h^{2} / 2 m \lambda^{2}$ |  |
| Photoelectric effect: $K E_{\text {max }}=e V_{\text {stop }}=h f-\Phi$ |  |
| UNIVERSAL: $\Delta x \Delta p_{x} \geq \hbar \quad \Delta E \Delta t \geq \hbar$ |  |
| $\psi^{*}(x) \psi(x) \equiv\|\psi(x)\|^{2}$ |  |
| $P_{a b}=\int_{a}^{b}\|\psi(x)\|^{2} d x, \quad a \leq x \leq b$ |  |
| (Slow) particle in fixed potential $\mathrm{U}:$ <br> $-\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \psi(x, t)}{\partial x^{2}}+U(x) \psi(x)=i \hbar \frac{\partial \psi(x, t)}{\partial t}$ |  |

## Physics 214 Common Formulae

Quantum stationary states (energy eigenstates):

$$
\begin{gathered}
\Psi(x, t)=\psi(x) e^{-i \omega t} \quad \text { where } \mathrm{E}=\hbar \omega \\
-\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \psi(x)}{\partial x^{2}}+U(x) \psi(x)=\hbar \omega \psi(x)=E \psi(x)
\end{gathered}
$$

In 1-D box: $n \lambda=2 L$ where $n=1,2 \ldots$
$\psi_{n}(x)=\sqrt{\frac{2}{L}} \sin \left(\frac{n \pi}{L} x\right) \quad$ for $\quad 0 \leq x \leq L$
$E_{n}=\frac{\hbar^{2}}{2 m}\left(\frac{n \pi}{L}\right)^{2}=\left(\frac{h^{2}}{8 m L^{2}}\right) n^{2}=E_{1} n^{2} \quad$ (*last part*)

## Box, 3-D:

$\psi(x, y, z)=\sqrt{\frac{8}{a b c}} \sin \left(\frac{n_{1} \pi}{a} x\right) \sin \left(\frac{n_{2} \pi}{b} y\right) \sin \left(\frac{n_{3} \pi}{c} z\right)$
$E\left(n_{1}, n_{2}, n_{3}\right)=\frac{h^{2}}{8 m}\left(\frac{n_{1}{ }^{2}}{a^{2}}+\frac{n_{2}{ }^{2}}{b^{2}}+\frac{n_{3}{ }^{2}}{c^{2}}\right)$
Simple Harmonic Oscillator (SHO):

$$
\begin{aligned}
& E_{n}=\left(n+\frac{1}{2}\right) \hbar \omega \text { where } n=0,1,2 \ldots \\
& \omega=\sqrt{k / m}
\end{aligned}
$$

Free slow particle with definite p:
$\Psi(x, t)=A e^{i(k x-\omega t)}$ with $\hbar \omega=\hbar^{2} k^{2} / 2 m$

| H-like atom |
| :---: |
| potential $U(r)=-\frac{\kappa Z e^{2}}{r}$ |
| $E_{n}=\frac{-1}{4 \pi \varepsilon_{o}} \frac{(Z e)^{2}}{2 a_{o}} \frac{1}{n^{2}}=-\frac{1}{\left(4 \pi \varepsilon_{o}\right)^{2}} \frac{m e^{4} Z^{2}}{2 \hbar^{2} n^{2}}$ |
| $=-13.606 \mathrm{eV} \frac{Z^{2}}{n^{2}}$ |
| Ground state: $\psi_{1 s}(r, \theta, \phi)=\frac{1}{\sqrt{\pi a_{o}^{3}}} e^{-r / a_{o}}$ |
| Radial density for s-state: $P(r) d r=4 \pi r^{2}\|\psi(r)\|^{2} d r$ |
| Form of $n, l, m$ eigenstate: |
| $\psi_{n t m}(r, \theta, \phi)=R_{n \ell}(r) Y_{\ell m}(\theta, \phi)$ |
| $Y_{00}=\frac{1}{\sqrt{4 \pi}}, \quad Y_{10}=\sqrt{\frac{3}{4 \pi}} \cos \theta$, |
| $Y_{1 \pm 1}=\mp \sqrt{\frac{3}{8 \pi}} \sin \theta e^{ \pm i \rho}$ |


| Tunneling |
| :---: |
| $T \approx G e^{-2 K L} \quad$ where $\quad G=16 \frac{E}{U_{0}}\left(1-\frac{E}{U_{0}}\right)$ |
| $K=\sqrt{\frac{2 m}{\hbar^{2}}\left(U_{0}-E\right)}=2 \pi \sqrt{\frac{2 m}{h^{2}}\left(U_{0}-E\right)}$ |

Angular momentum and magnetism
Orbital: $L_{z}=m \hbar$ where $m=0, \pm 1, \pm 2, \ldots \pm \ell$

$$
L^{2}=\ell(\ell+1) \hbar^{2} \text { where } \ell=0,1,2, \ldots
$$

Spin: $S_{z}=m_{s} \hbar$ where $m_{s}= \pm 1 / 2$
Magnetic energy: $U=-\vec{\mu} \cdot \vec{B}$
Force: $F_{z}=\mu_{z} \frac{d B_{z}}{d z}$ where $\mu_{z} \approx-\frac{e}{m_{e}} S_{z}$


