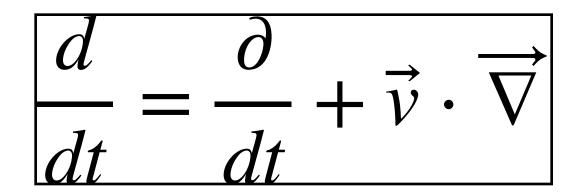
# **Introduction to Fluid Dynamics**

#### Yuk Tung Liu **University of Illinois at Urbana-Champaign April 2024**

#### **Convective Derivatives and Partial Derivatives**

Partial time derivative 
$$\frac{\partial q}{\partial t}$$
: rate of change of  $q(t, x)$   
Convective time derivative  $\frac{dq}{dt}$ : rate of change of  
 $\frac{dq}{dt} = \frac{\partial q}{\partial t} + \frac{\partial q}{\partial x}\frac{dx}{dt} + \frac{\partial q}{\partial y}\frac{dy}{dt} + \frac{\partial q}{\partial z}\frac{dz}{dt} = \frac{\partial q}{\partial t} + \frac{\partial q}{\partial t} + \frac{\partial q}{\partial t}\frac{dx}{dt} = \frac{\partial q}{\partial t} + \frac{\partial q}{\partial t}\frac{dx}{dt}$ 



x,y,z) at a fixed location.

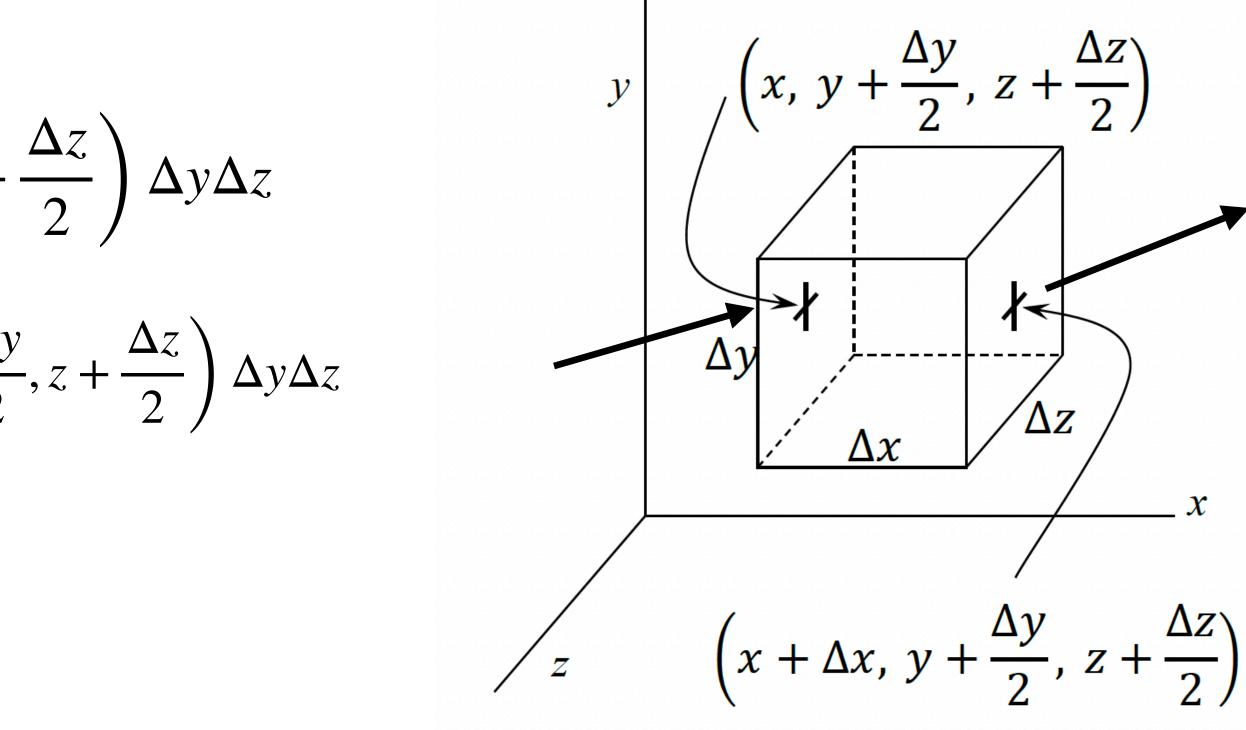
of q along a path.

 $+\frac{\partial q}{\partial x}v_x + \frac{\partial q}{\partial y}v_y + \frac{\partial q}{\partial z}v_z$ 

## **Continuity Equation I**

Net mass flow rate in the x-direction:

$$\Delta \dot{m}_{x} = \rho \left( x, y + \frac{\Delta y}{2}, z + \frac{\Delta z}{2} \right) v_{x} \left( x, y + \frac{\Delta y}{2}, z + \frac{\Delta y}{2} \right)$$
$$-\rho \left( x + \Delta x, y + \frac{\Delta y}{2}, z + \frac{\Delta z}{2} \right) v_{x} \left( x + \Delta x, y + \frac{\Delta y}{2} \right)$$
$$= -\frac{\partial}{\partial x} (\rho v_{x}) \Delta x \Delta y \Delta z$$
$$= -\frac{\partial}{\partial x} (\rho v_{x}) \Delta V$$





## **Continuity Equation II**

Similarly, net mass flow rate in the y and z directions are

$$\Delta \dot{m}_y = -\frac{\partial}{\partial y}(\rho v_y)\Delta V$$
 ,  $\Delta \dot{m}_z = -\frac{\partial}{\partial z}(\rho v_y)$ 

Total mass flowing into the volume/time is

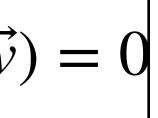
$$\Delta \dot{m} = \frac{\partial}{\partial t} (\rho \Delta V) = -\left[\frac{\partial}{\partial x} (\rho v_x) + \frac{\partial}{\partial y} (\rho v_y) + \frac{\partial}{\partial y} (\rho$$

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v})$$

This is called the *continuity* equation.



 $+ \frac{\partial}{\partial z} (\rho v_z) \bigg| \Delta V = - \overrightarrow{\nabla} \cdot (\rho \overrightarrow{v}) \Delta V$ 



### **Continuity Equation III**

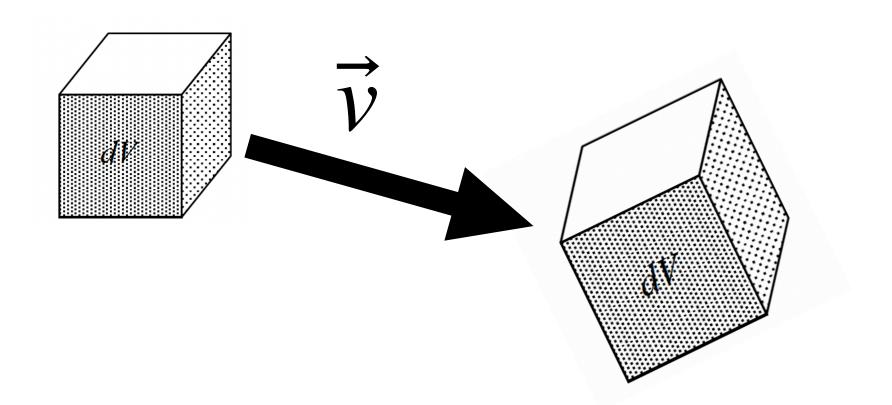
Suppose we follow the motion of the fluid.

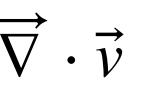
Recall: 
$$\frac{d\rho}{dt} = \frac{\partial\rho}{\partial t} + \vec{v} \cdot \vec{\nabla}\rho$$

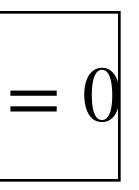
$$\frac{d\rho}{dt} = -\overrightarrow{\nabla}\cdot(\rho\overrightarrow{v}) + \overrightarrow{v}\cdot\overrightarrow{\nabla}\rho = -\rho\overrightarrow{\nabla}$$

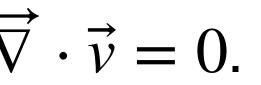
$$\frac{d\rho}{dt} + \rho \overrightarrow{\nabla} \cdot \overrightarrow{v}$$

For incompressible fluid,  $d\rho/dt = 0$ . Hence  $\vec{\nabla} \cdot \vec{v} = 0$ .





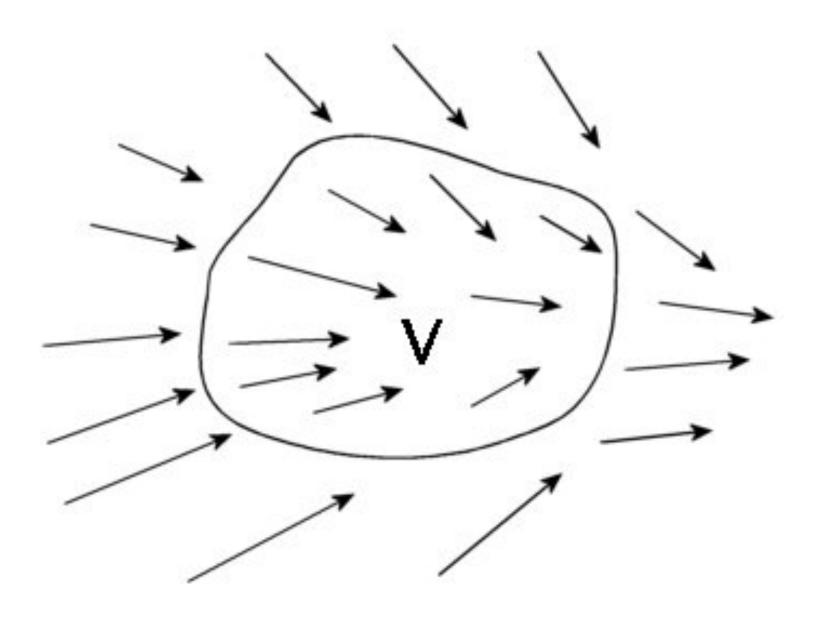


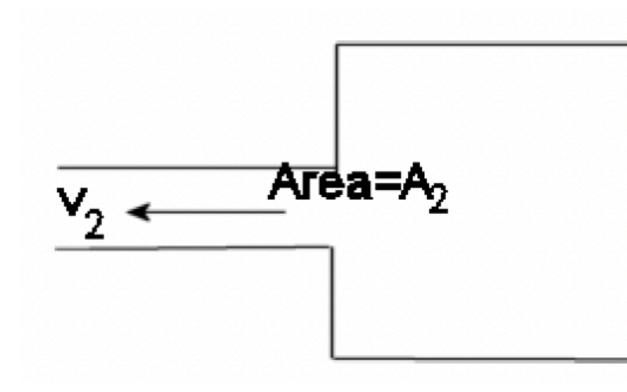


#### Integral Form of Continuity Equation

$$\begin{split} M &= \int_{V} \rho dV \\ \frac{dM}{dt} &= \int_{V} \frac{\partial \rho}{\partial t} dV = -\int_{V} \vec{\nabla} \cdot (\rho \vec{v}) dV \\ &= -\oint_{\partial V} \rho \vec{v} \cdot d\vec{S} \end{split}$$

Rate of increase in mass inside a volume V = net mass flow into the volume per unit time.





Consider air flowing from a tube with cross-sectional area  $A_1$  into a region with cross-sectional area  $A_2$ . In steady air flow, dM/dt = 0.

$$\rho v_1 A_1 = \rho v_2 A_2$$

$$v_2 = \frac{A_1}{A_2} v_1$$

#### **Example 1: Flow Tube**

Area=A1

### **Example 2: Water Leak**

There is a small hole at the bottom of a container and water leaks out from the hole at speed v.

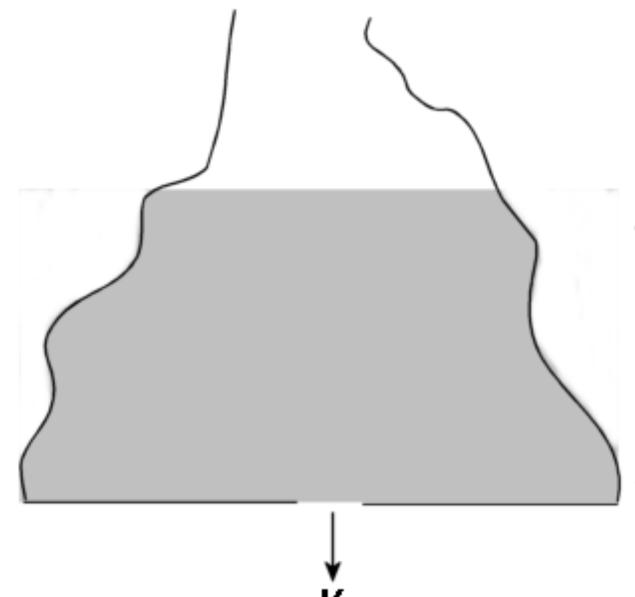
The water level y decreases slowly.

$$\frac{dM}{dt} = \frac{d(\rho V)}{dt} = -\rho v A_h$$

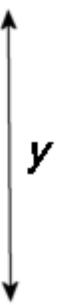
 $A_h$ : area of the hole. V = Volume of water inside the container.

$$\frac{dV}{dt} = A(y)\dot{y} \qquad A(y): \text{ cross-section}$$

$$\Rightarrow \qquad \dot{y} = -\frac{A_h}{A(y)}v$$



nal area at y



### **Momentum Equation**

Net force associated with pressure in *x*-direction:

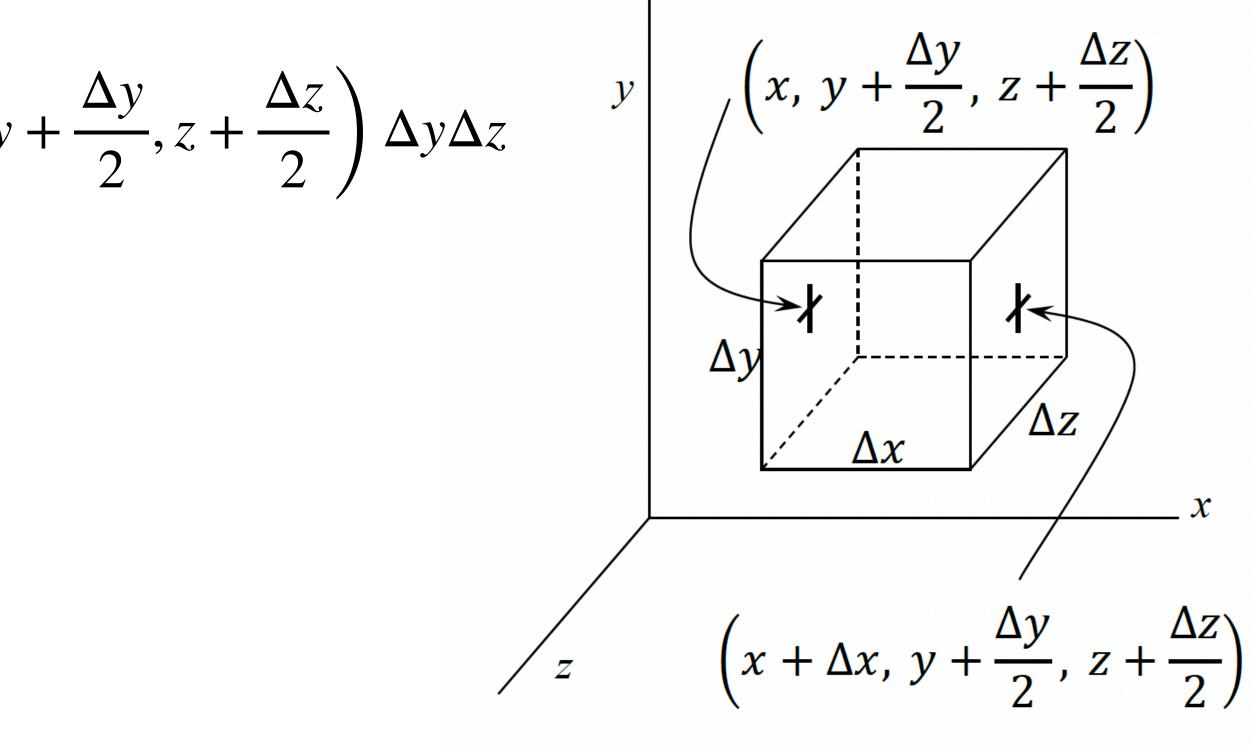
$$\Delta f_x = P\left(x, y + \frac{\Delta y}{2}, z + \frac{\Delta z}{2}\right) \Delta y \Delta z - P\left(x + \Delta x, y\right)$$
$$= -\frac{\partial P}{\partial x} \Delta x \Delta y \Delta z$$
$$= -\frac{\partial P}{\partial x} \Delta V$$

Similarly,  $\Delta f_y = -\frac{\partial P}{\partial v} \Delta V$ ,  $\Delta f_z = -\frac{\partial P}{\partial z} \Delta V$ 

Total net force associated with pressure:

$$\Delta \vec{f} = -\left(\frac{\partial P}{\partial x}\hat{x} + \frac{\partial P}{\partial y}\hat{y} + \frac{\partial P}{\partial z}\hat{z}\right)\Delta V = -\overrightarrow{\nabla}P.$$





 $\Delta V$ 



In addition to pressure, gravity also acts on the fluid:

$$\Delta \vec{f} = -\vec{\nabla} P \Delta V + (\rho \Delta V) \vec{g}$$

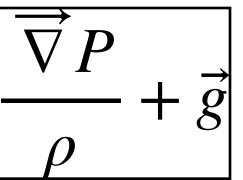
From Newton's second law:

$$(\rho \Delta V) \frac{d\vec{v}}{dt} = -\vec{\nabla} P \Delta V + \rho \vec{g} \Delta V$$
$$\frac{d\vec{v}}{dt} = \frac{d\vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} = -\vec{v}$$

This is also called Euler's equation.

It describes the conservation of momentum of an *ideal fluid* (i.e. without viscosity).

## **Momentum Equation (cont)**



$$\vec{v} \cdot \vec{\nabla} \vec{v} = v_x \frac{\partial \vec{v}}{\partial x} + v_y \frac{\partial \vec{v}}{\partial y} + v_z \frac{\partial \vec{v}}{\partial z}$$

$$= \left(v_x \frac{\partial v_x}{\partial x} + v_y \frac{\partial v_x}{\partial y} + v_z \frac{\partial v_x}{\partial z}\right) \hat{x} + \left(v_x \frac{\partial v_y}{\partial x} + v_y \frac{\partial v_y}{\partial y} + v_z \frac{\partial v_y}{\partial z}\right) \hat{y} + \left(v_x \frac{\partial v_z}{\partial x} + v_y \frac{\partial v_z}{\partial y} + v_z \frac{\partial v_z}{\partial z}\right)$$

If  $\vec{v}$  is represented by a row vector,  $\vec{\nabla} \vec{v}$  represented by a 3 × 3 matrix,  $\vec{v} \cdot \vec{\nabla} \vec{v}$  can be represented by a row vector by

$$\vec{v} \cdot \vec{\nabla} \vec{v} = (v_x \quad v_y \quad v_z) \begin{pmatrix} \frac{\partial v_x}{\partial x} & \frac{\partial v_y}{\partial x} & \frac{\partial v_z}{\partial x} \\ \frac{\partial v_x}{\partial y} & \frac{\partial v_y}{\partial y} & \frac{\partial v_z}{\partial y} \\ \frac{\partial v_x}{\partial z} & \frac{\partial v_y}{\partial z} & \frac{\partial v_z}{\partial z} \end{pmatrix}$$

The Meaning of  $\vec{v} \cdot \vec{\nabla} \vec{v}$ 



### **Hydrostatics**

Momentum equation:  $\frac{d\vec{v}}{dt} = -\frac{\vec{\nabla}P}{\rho} + \vec{g}$ 

Hydrostatics:  $\vec{v} = 0 \Rightarrow \vec{\nabla} P = \rho \vec{g}$ 

Pressure gradient is parallel to  $\vec{g} \Rightarrow$  surface of constant *P* (isobar) is perpendicular to  $\vec{g}$ .  $0 = \overrightarrow{\nabla} \times \overrightarrow{\nabla} P = \overrightarrow{\nabla} \rho \times \vec{g}$ 

 $\Rightarrow$  density gradient is parallel to  $\vec{g} \Rightarrow$  surface of constant  $\rho$  is perpendicular to  $\vec{g}$ .

Let  $\vec{g} = g\hat{z}$  ( $\hat{z}$  points downward),  $P = P(z), \rho = \rho(z)$ .

 $\overrightarrow{\nabla}P = \frac{dP}{dz}\hat{z} = \rho g\hat{z}$ 

### Hydrostatics (cont)

$$\frac{dP}{dz} = \rho g$$
$$P(z) = \int \rho(z)gdz$$

Consider a cylinder with cross-sectional area A and height z.

$$P(z) = \frac{1}{A} \left( \int \rho(z) A dz \right) g = \frac{M_f(z)g}{A}$$

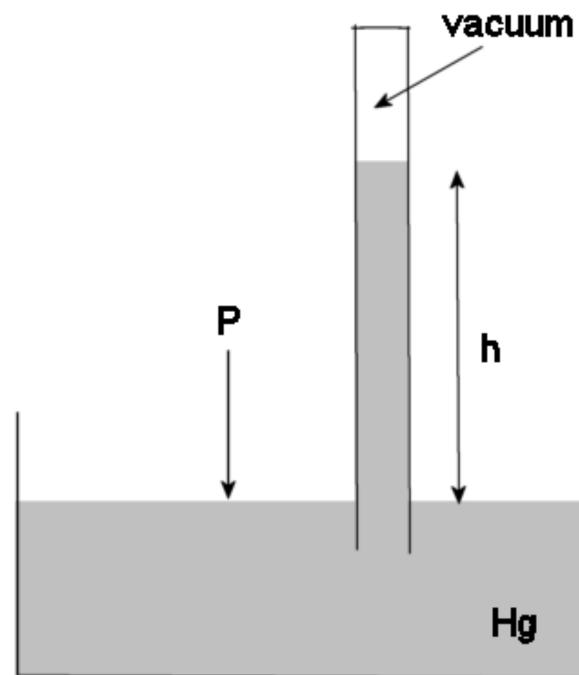
Pressure at depth z is the weight of the fluid per unit area above z.

For incompressible fluid,  $\rho(z) = \rho$  is constant,

 $P(z) = \rho g z$ 

Ζ

#### **Mercury Barometer**



$$P = \rho_{\rm Hg} g h$$

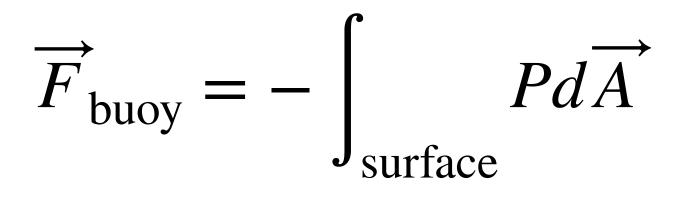
Standard atmospheric pressure = 101kPa  $\approx$ 760 mmHg

Hg

# **Archimedes'** Principle

Consider an object floating stationary in a fluid.

Buoyant force acting on the object:



Imagine removing the body and replacing it by fluid.

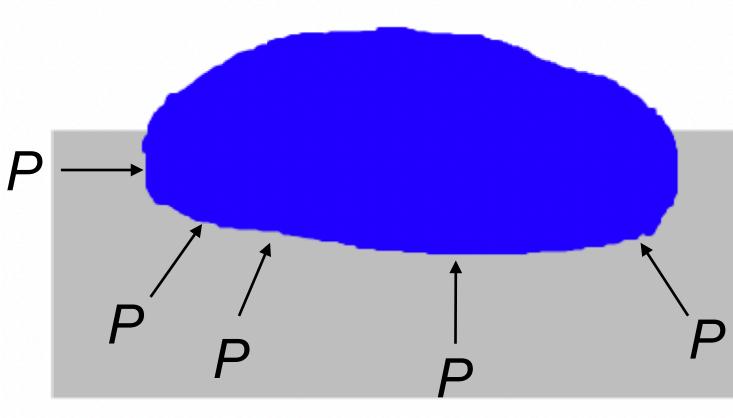
Pressure P(z) and density  $\rho(z)$  remain the same.

Hydrostatic eq:  $\overrightarrow{\nabla}P = \rho \overrightarrow{g}$ 

$$\int_{V} \overrightarrow{\nabla} P dV = \int \rho \vec{g} dV \quad \Rightarrow \quad \int_{\text{surface}} P d\vec{A} = M_{f} \vec{g}$$

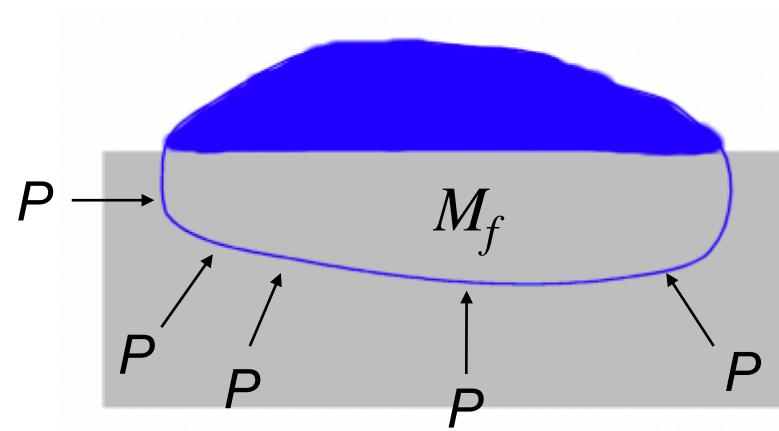
 $M_f$ : mass of the fluid displaced by the object.

Archimedes' principle:  $\vec{F}_{buoy} = -M_f \vec{g}$  (buoyant force = weight of fluid displaced by the object)















### **Tip of the lceberg**

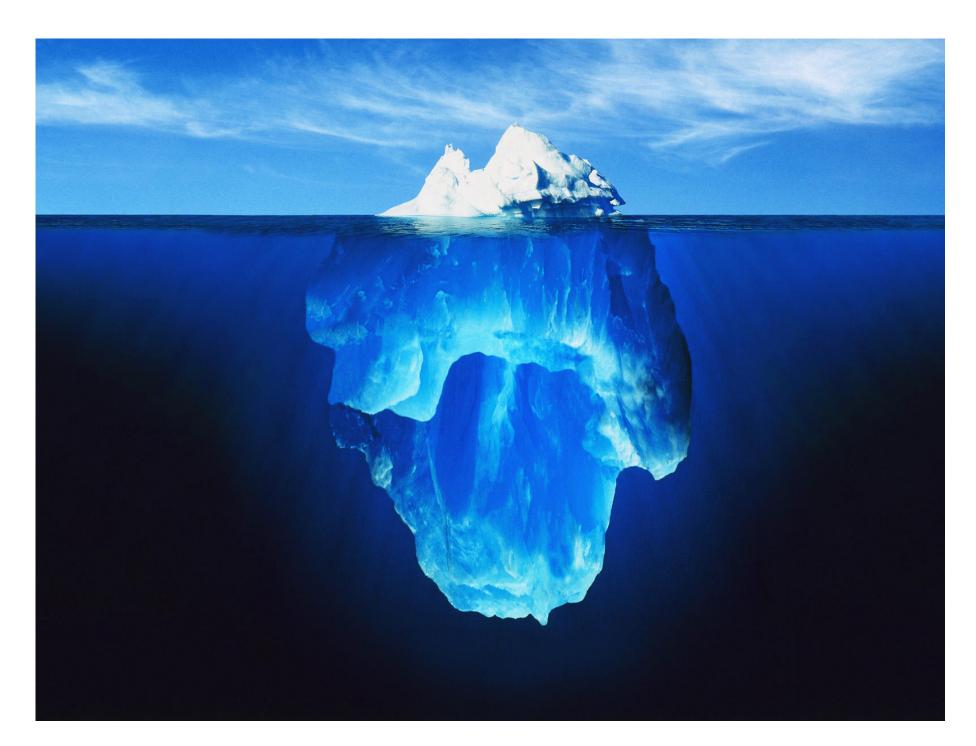
Density of ice  $\rho_i = 920 \text{ kg/m}^3$ Density of sea water  $\rho_w = 1027 \text{ kg/m}^3$  $V_a$ : volume of iceberg above water V: total volume of iceberg

In static state, weight of iceberg = buoyant force

$$\rho_i Vg = \rho_w (V - V_a)g$$

$$\frac{V_a}{V} = \frac{\rho_w - \rho_i}{\rho_w} = 0.10$$

Only 10% of the iceberg is above the sea water!



#### Credit: <u>clipground.com</u>

#### Earth's Atmosphere I

Earth's atmospheric pressure is closely approximated by the hydrostatic equilibrium.

Let 
$$\vec{g} = -g\hat{z}$$
 ( $\hat{z}$  points upward).  
 $\frac{dP}{dz} = -\rho g$  ideal gas law:  $P = nkT$ 

 $R = N_A k = 8.31 \text{J/(mol K)} = \text{gas constant}$ 

M: molar mass of air = 0.02896 kg/mol (78% N<sub>2</sub>, 21% O<sub>2</sub>, 0.9% Ar and small amount of other gases)

$$\frac{dP}{dz} = -\frac{Mg}{RT}P \quad \Rightarrow \quad \frac{dP}{P} = -\frac{Mg}{P}$$

$$P(z) = P_0 \exp\left(-\int_0^z \frac{Mg}{RT(z')}\right)$$

 $P_0$ : pressure at z=0.

 $r = \frac{\rho}{M}RT$ 

 $\frac{Mg}{RT}dz$ 

-dz'



#### Earth's Atmosphere II

\* If  $T = T_0$  = constant (isothermal)

$$P(z) = P_0 e^{-Mgz/RT_0}$$
 (isothermal)

\* If  $T = T_0 - Lz$  (*L* is called the temperature lapse rate):

$$P(z) = P_0 \left( 1 - \frac{Lz}{T_0} \right)^{Mg/RL}$$
 (lapse)

Recall:

$$\lim_{k \to \infty} \left( 1 + \frac{x}{k} \right)^k = \lim_{k \to \infty} \exp\left[ k \ln\left( 1 + \frac{x}{k} \right) \right] = \lim_{k \to \infty} \exp\left( k \cdot \frac{x}{k} \right) = e^x$$

The lapse equation reduces to the isothermal equation in the limit  $L \rightarrow 0$ .

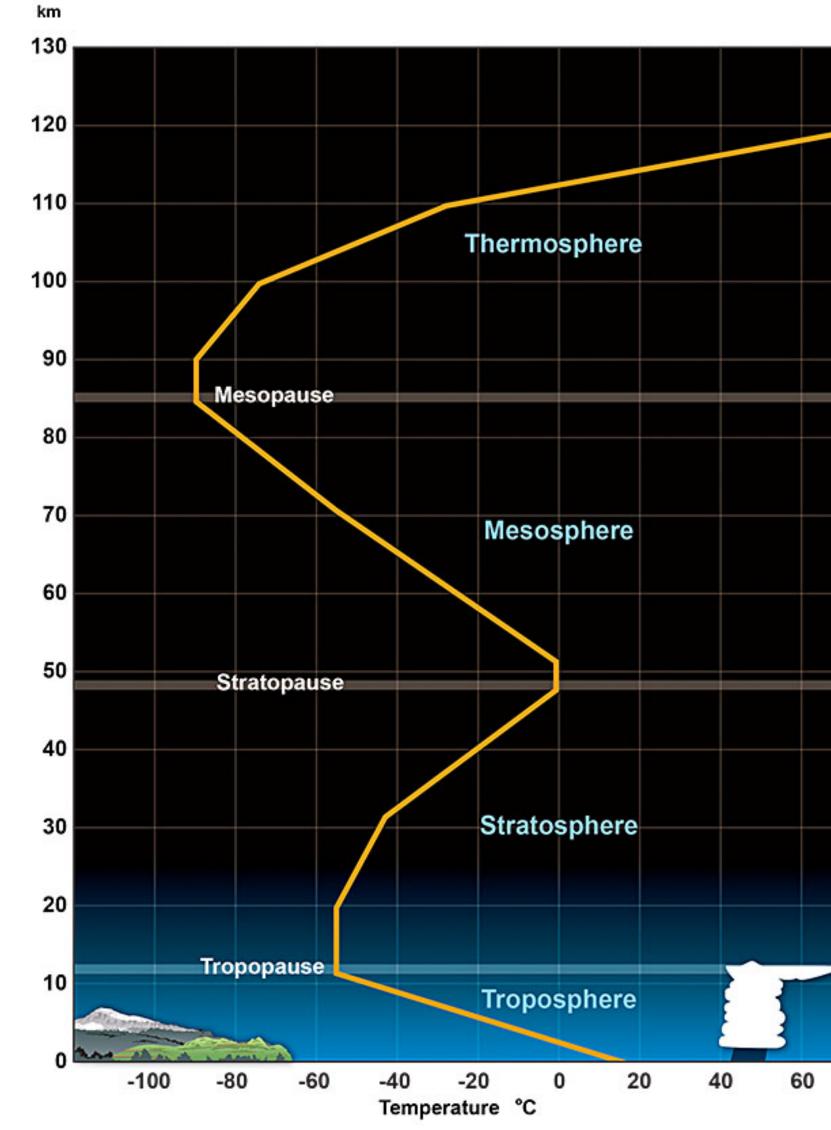
#### Earth's Atmosphere III

More realistic atmospheric model divides the atmosphere into several layers. Each layer has its own temperature lapse rate:

$$P(z) = P_b \left[ 1 - \frac{L_b(z - z_b)}{T_b} \right]^{Mg/RL_b}$$

 $P_b$ : pressure at the bottom of layer b.

- $T_b$ : temperature at the bottom of layer b.
- $L_b$  : temperature lapse rate in layer b.
- $z_b$  : altitude at the bottom of layer *b*.



Credit: NOAA



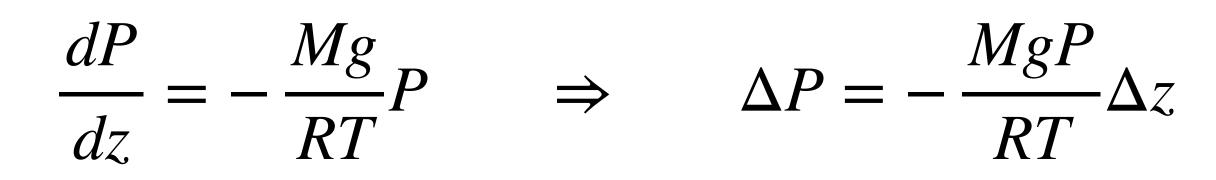
#### Earth's Atmosphere IV

Sub- script b	Geopotential height above mean Sea level (z)		Static pressure P <sub>b</sub>		Standard temperature (K)	Temperature lapse rate	
						$L_b$	
	<b>(m)</b>	<b>(ft)</b>	(Pa)	(inHg)	$T_b$	<b>(K/m)</b>	(K/ft)
0	0	0	101 325.00	29.92126	288.15	0.0065	0.0019812
1	11 000	36,089	22 632.10	6.683245	216.65	0.0	0.0
2	20 000	65,617	5474.89	1.616734	216.65	-0.001	-0.0003048
3	32 000	104,987	868.02	0.2563258	228.65	-0.0028	-0.00085344
4	47 000	154,199	110.91	0.0327506	270.65	0.0	0.0
5	51 000	167,323	66.94	0.01976704	270.65	0.0028	0.00085344
6	71 000	232,940	3.96	0.00116833	214.65	0.002	0.0006096

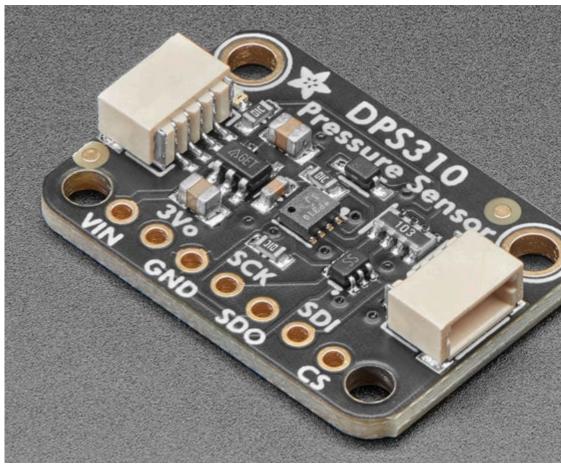
Credit: Wikimedia (https://en.wikipedia.org/wiki/Barometric\_formula)

#### **DPS 310 Pressure Sensor**

According to Adafruit, their DPS 310 pressure sensor can measure the change in pressure to an accuracy of 0.2 Pa.



 $\Delta P = 0.2$  Pa corresponds to  $\Delta z = 1.7$  cm for M = 0.02896 kg/mol, P = 101 kPa, and T = 300 K.

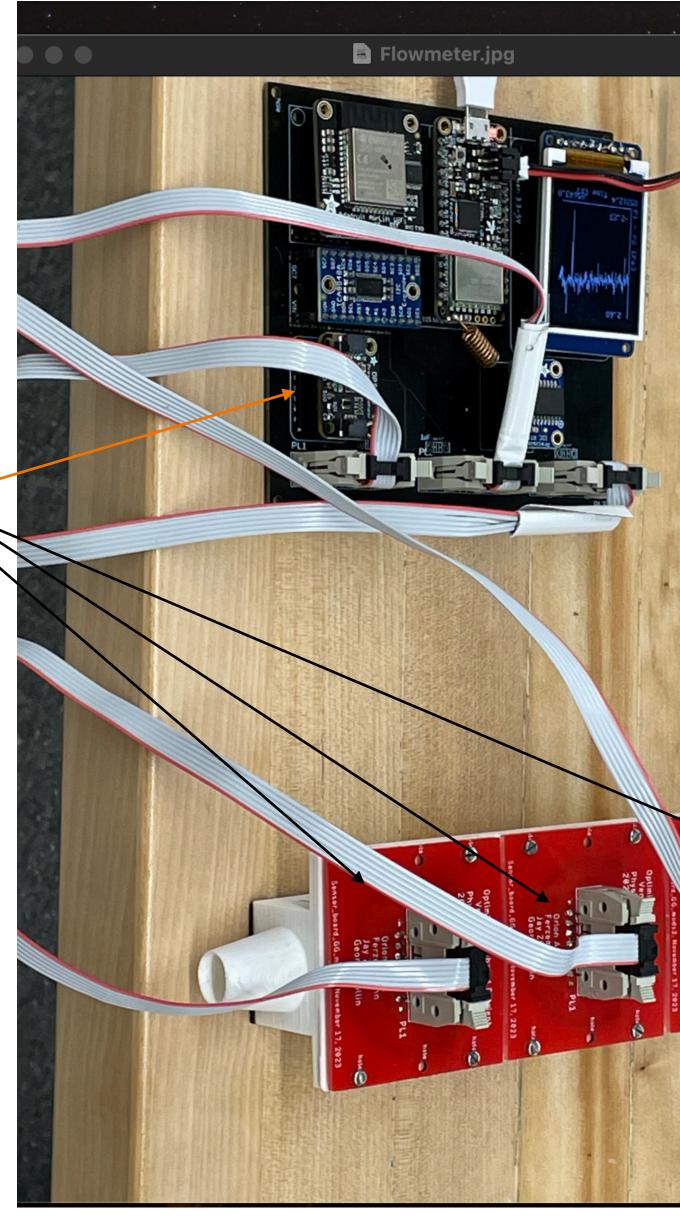


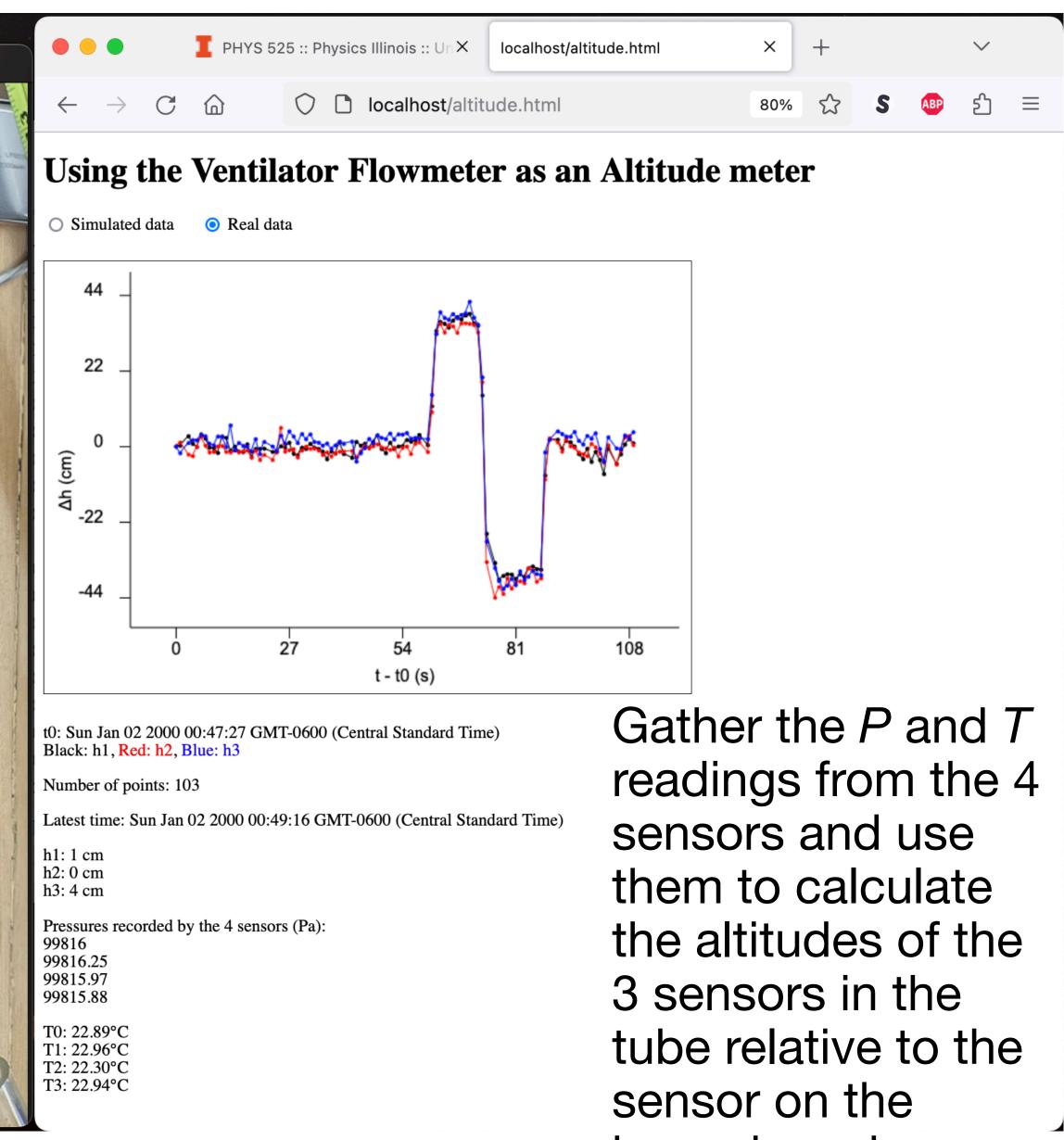
Credit: Adafruit



#### **Class Demonstration**

#### 4 DPS 310 sensors: 1 on the home board, 3 inside the flow tube





home board.

Momentum equation:  $\frac{d\vec{v}}{dt} = -\frac{\vec{\nabla}P}{\rho} + \vec{g}$  $\vec{v} \cdot \frac{d\vec{v}}{dt} = -\frac{\vec{v} \cdot \vec{\nabla} P}{\rho} + \vec{v} \cdot \vec{g} \quad , \qquad \vec{v} \cdot \cdot$ 

 $\vec{g} = -\vec{\nabla}U$ , U = gh is gravitational potential, h is height from a reference point.

Gravity is static near Earth's surface,  $\partial U/\partial t = 0$ .

$$\frac{dU}{dt} = \frac{\partial U}{\partial t} + \vec{v} \cdot \vec{\nabla} U = \vec{v} \cdot \vec{\nabla} U = -\vec{v} \cdot \vec{g}$$

$$\Rightarrow \quad \frac{d}{dt} \left( \frac{1}{2} v^2 + U \right) + \frac{\vec{v} \cdot \vec{\nabla} P}{\rho} = 0$$

#### **Energy Equation**

$$\frac{d\vec{v}}{dt} = \frac{1}{2}\frac{d}{dt}(\vec{v}\cdot\vec{v}) = \frac{d}{dt}\left(\frac{v^2}{2}\right)$$

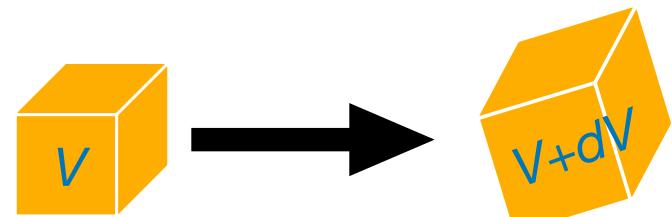
### First Law of Thermodynamics

Consider a fluid element in a small volume V.

Mass  $m = \rho V$ , internal energy is E. First law of thermodynamics: dE = dQ - PdV

dQ is the amount of heat added to the volume. In the absence of heat generation and heat flow, dQ=0. The system is said to be adiabatic and  $\frac{dE}{dt} = -P\frac{dV}{dt}$ . Divide the equation by the mass  $m = \rho V$  and write w = E/m (specific internal energy).

$$\frac{dw}{dt} = -\frac{P}{\rho V}\frac{dV}{dt} = -P\frac{d}{dt}\left(\frac{V}{\rho V}\right) = -P\frac{d}{dt}\left(\frac{1}{\rho}\right) = -\frac{d}{dt}\left(\frac{P}{\rho}\right) + \frac{1}{\rho}\frac{dP}{dt}$$
$$\frac{d}{dt}\left(w + \frac{P}{\rho}\right) = \frac{1}{\rho}\frac{dP}{dt} = \frac{1}{\rho}\frac{\partial P}{\partial t} + \frac{\vec{v}\cdot\vec{\nabla}P}{\rho}$$
$$V \longrightarrow V$$
$$\frac{\vec{v}\cdot\vec{\nabla}P}{\rho} = \frac{d}{dt}\left(w + \frac{P}{\rho}\right) - \frac{1}{\rho}\frac{\partial P}{\partial t}$$
$$V = \rho V = (\rho + d\rho)(V + \rho)$$



uid element dV

#### **Bernoulli's Equation**

Previous slides:

$$\frac{d}{dt}\left(\frac{1}{2}v^2 + U\right) + \frac{\vec{v}\cdot\vec{\nabla}P}{\rho} = 0 \quad , \quad \frac{\vec{v}\cdot\vec{\nabla}P}{\rho} = \frac{d}{dt}\left(w + \frac{P}{\rho}\right) - \frac{1}{\rho}\frac{\partial P}{\partial t}$$

Combine these two equations:

$$\frac{d}{dt}\left(\frac{1}{2}v^2 + \frac{P}{\rho} + U + w\right) = \frac{1}{\rho}\frac{\partial P}{\partial t}$$

In steady flow,  $\partial P/\partial t = 0$ , the resulting equation is called Bernoulli's equation.

$$\frac{d}{dt}\left(\frac{1}{2}v^2 + \frac{P}{\rho} + U + w\right) = 0$$

Recall: 
$$\frac{dw}{dt} = -P\frac{d}{dt}\left(\frac{1}{\rho}\right) = 0$$
 for incompressible fluid  $\Rightarrow \frac{d}{dt}\left(\frac{1}{2}v^2 + \frac{P}{\rho} + U\right) = 0$ 

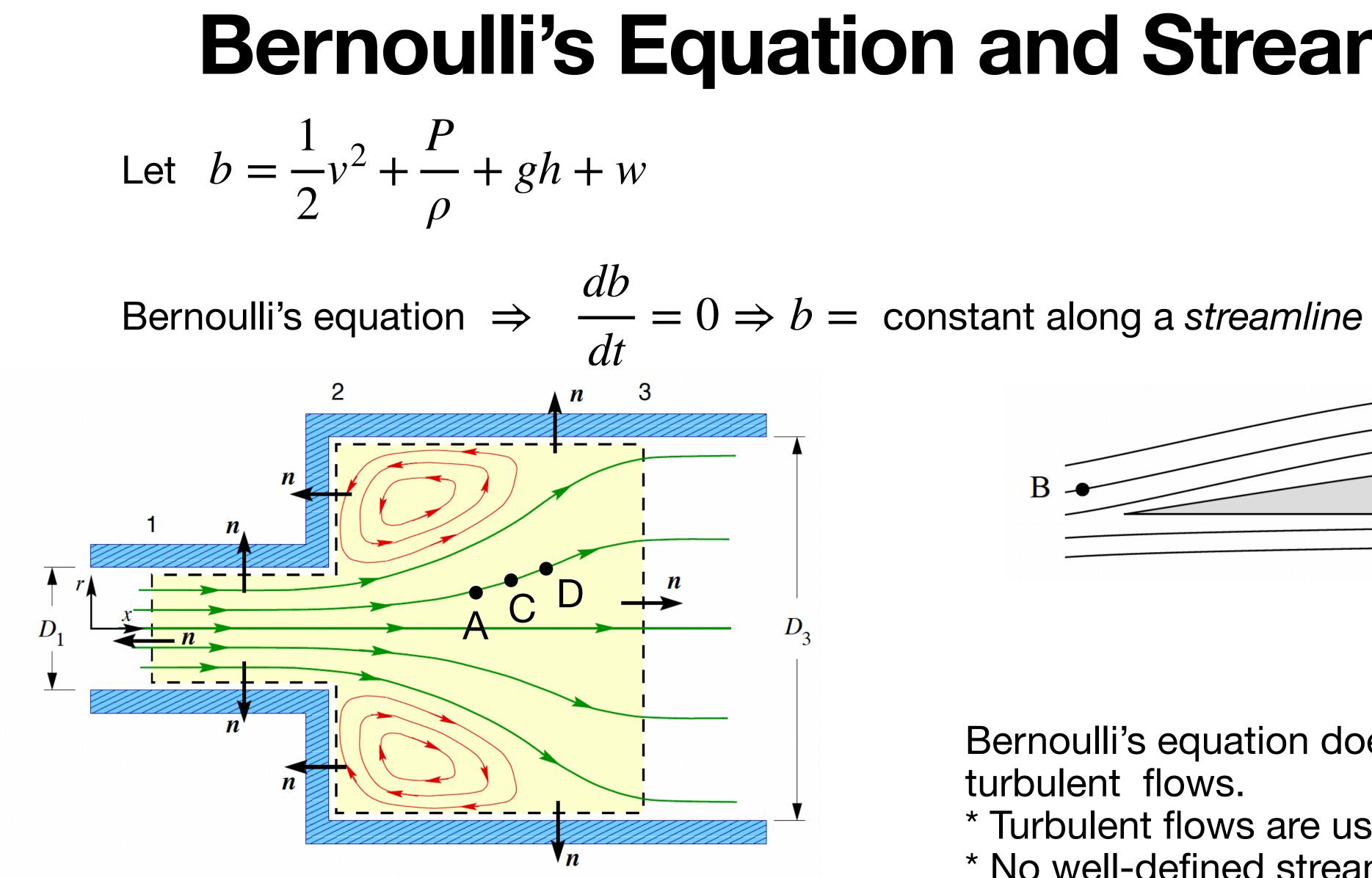
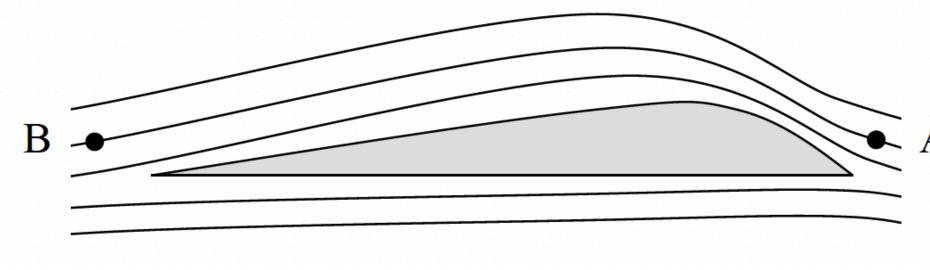


Figure 4.8: Flow through a rapidly-expanding pipe.

Figure credit: J.M. McDonough, Lectures In Elementary Fluid **Dynamics: Physics, Mathematics and Applications** 

#### **Bernoulli's Equation and Streamline**



Bernoulli's equation doesn't apply to turbulent flows.

\* Turbulent flows are usually not steady

- \* No well-defined streamlines
- \* Viscosity is important





Water is flowing out of a rectangular tank from a small hole at the bottom. How long does it take to excavate the water from the tank?

Apply Bernoulli's equation at the top and at the hole:

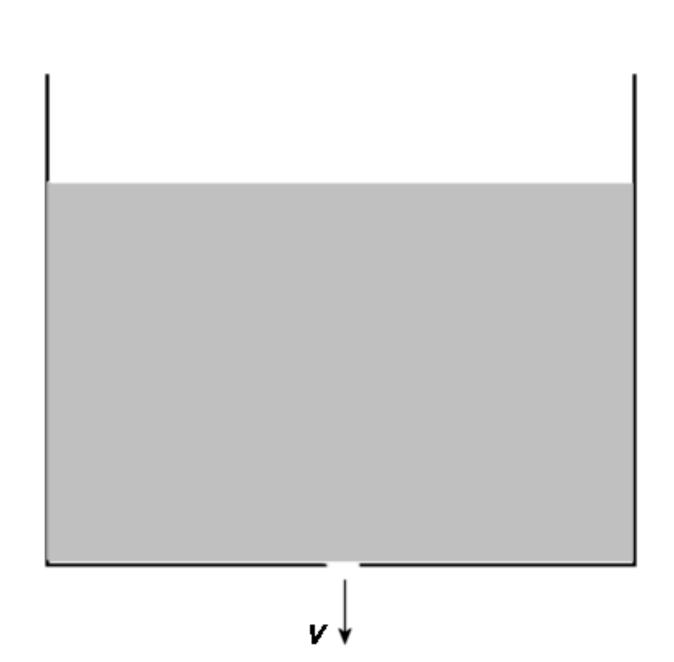
$$\frac{1}{2}\dot{y}^{2} + \frac{P}{\rho} + gy = \frac{1}{2}v^{2} + \frac{P}{\rho} \implies v^{2} - \dot{y}^{2} =$$
Previously, we find  $\dot{y} = -\frac{A_{h}}{A}v$ 

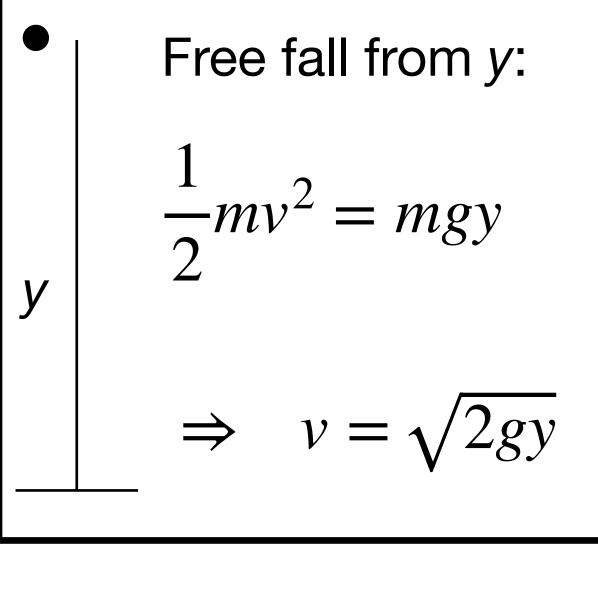
 $A_h$ : area of the hole, A: cross-sectional area of the tank.

$$\Rightarrow \left(1 - \frac{A_h^2}{A^2}\right) v^2 = 2gy \quad ,$$
$$v = \sqrt{2gy} \left(1 - \frac{A_h^2}{A^2}\right)^{-1/2} \approx \sqrt{2gy} \quad \text{for } A_h \ll A$$

This is the free-fall speed from y. As the water level drops, the speed also decreases.

#### Example









Rate of change of water level:  $\dot{y} = -\frac{A_h}{\Delta}v = -\frac{A_h}{\Delta}\sqrt{2gy}$ 

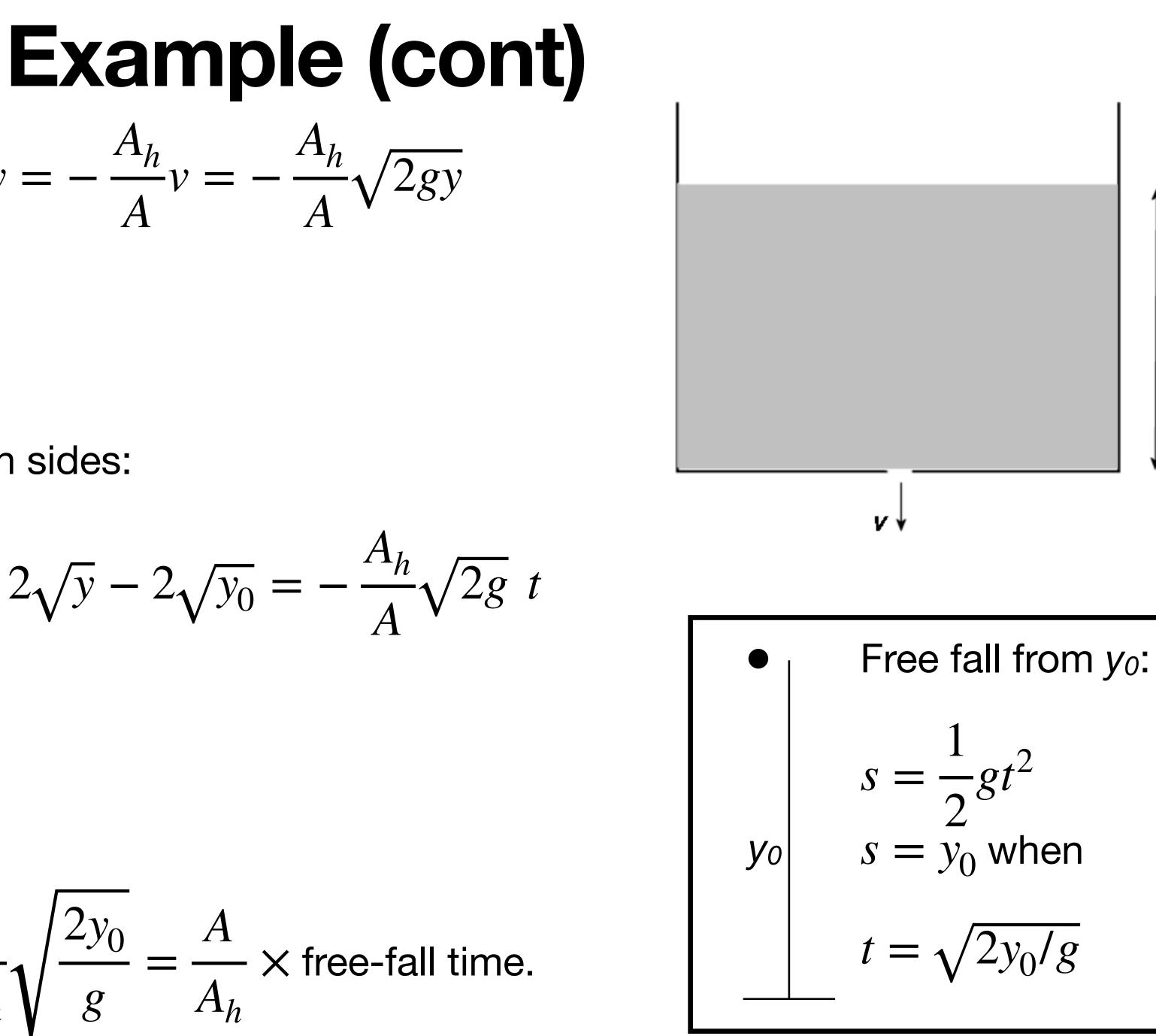
$$\frac{dy}{\sqrt{y}} = -\frac{A_h}{A}\sqrt{2g}dt$$

Let  $y_0 = y(t = 0)$ . Integrate both sides:

$$\int_{y_0}^{y} \frac{dy'}{\sqrt{y'}} = -\frac{A_h}{A}\sqrt{2g} \ t \qquad , \qquad 2\sqrt{y} - 2\sqrt{y}$$

$$y(t) = \left(\sqrt{y_0} - \frac{A_h}{A}\sqrt{\frac{g}{2}}t\right)^2$$

Setting y(T) = 0 gives  $T = \frac{A}{A_h} \sqrt{\frac{2y_0}{g}} = \frac{A}{A_h} \times$  free-fall time.









$$T = \frac{A}{A_h} \sqrt{\frac{2y_0}{g}}$$

For  $y_0 = 0.3 \text{ m}$ ,  $A/A_h = 40$ ,  $T \approx 10 \text{ s}$ .

Bernoulli's equation only applies to steady flow.

It's still a good approximation if the rate of change is sufficiently slow, which requires  $T \gg$  dynamical time scales.

Two dynamical time scales:

- (1) Time associated with pressure ~ time for sound to travel  $y_0$ :  $\tau = y_0/c_s$ . Sound speed in water  $\approx 1500$  m/s,  $\tau \approx 0.0002$  s  $\ll T$ .
- (2) Time associated with gravity ~ free-fall time.  $T = A/A_h \times$  free-fall time = 40 free-fall time. Relative error in estimated  $T \sim 1/40 = 2.5 \%$ .

## Example (cont)

*V* ∜



Vorticity is defined as  $\overrightarrow{\omega} = \overrightarrow{\nabla} \times \overrightarrow{v}$ . In Cartesian coordinates,  $\overrightarrow{\omega} = \left(\frac{\partial v_z}{\partial y} - \frac{\partial v_y}{\partial z}\right)\hat{x} + \left(\frac{\partial v_x}{\partial z} - \frac{\partial v_z}{\partial x}\right)\hat{y}$ 

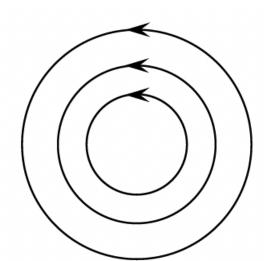
It describes the local spinning motion of fluid. Consider the velocity in the fluid near a vortex looks like this: The velocity field is given by  $\vec{v} = \vec{\Omega} \times \vec{r}$ , where  $\vec{\Omega}$  is a constant vector. In cylindrical coordinates with  $\overrightarrow{\Omega} = \Omega \hat{z}$ , we have  $v_{\phi} = \Omega r$  and  $v_r = v_z = 0$ .

$$\overrightarrow{\omega} = \frac{1}{r} \frac{\partial}{\partial r} (rv_{\phi})\hat{z} = 2\Omega\hat{z}$$

The fluid is irrotational if  $\vec{\omega} = 0$ .

## Vorticity

$$\hat{y} + \left(\frac{\partial v_y}{\partial x} - \frac{\partial v_x}{\partial y}\right)\hat{z}$$



### **Vector Derivatives in Cylindrical Coordinates**

CYLINDRICAL  $dI = dr \hat{r} + r d\phi \hat{\phi} + dz \hat{z}; d\tau = r dr d\phi dz$ Gradient.  $\nabla t = \frac{\partial t}{\partial r}\hat{r} + \frac{1}{r}\frac{\partial t}{\partial \phi}\hat{\phi} + \frac{\partial t}{\partial z}\hat{z}$ Divergence.  $\nabla \cdot \mathbf{v} = \frac{1}{r} \frac{\partial}{\partial r} (rv_r) + \frac{1}{r} \frac{\partial v_{\phi}}{\partial \phi} + \frac{\partial v_z}{\partial z}$ Curl.  $\nabla \times \mathbf{v} = \left[\frac{1}{r}\frac{\partial v_z}{\partial \phi} - \frac{\partial v_{\phi}}{\partial z}\right]\hat{r} + \left[\frac{\partial v_r}{\partial z} - \frac{\partial v_z}{\partial r}\right]\hat{\phi}$  $+\frac{1}{r}\left[\frac{\partial}{\partial r}(rv_{\phi})-\frac{\partial v_{r}}{\partial \phi}\right]\hat{z}$ Laplacian.

 $\nabla^2 t = \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial t}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 t}{\partial d^2} + \frac{\partial^2 t}{\partial z^2}$ 

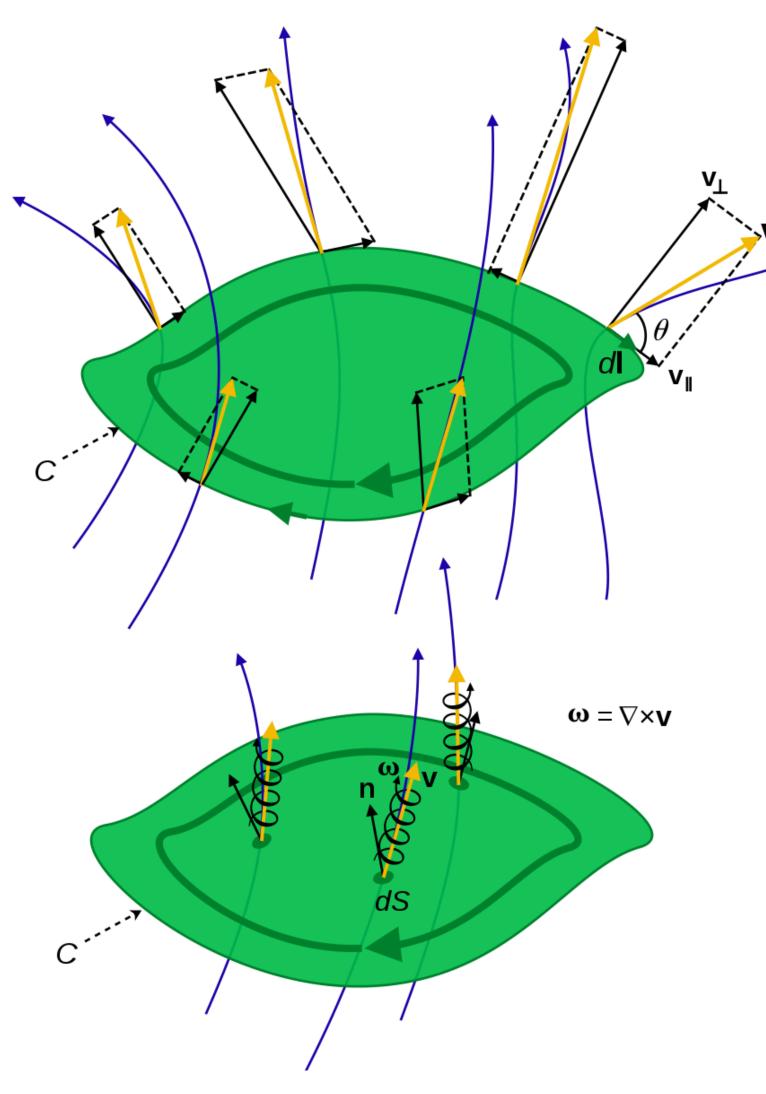


### Circulation

- Circulation is closely related to vorticity
- Circulation of a fluid around a closed loop is defined as  $\Gamma = \oint \vec{v} \cdot d\vec{l}$
- Stoke's theorem:

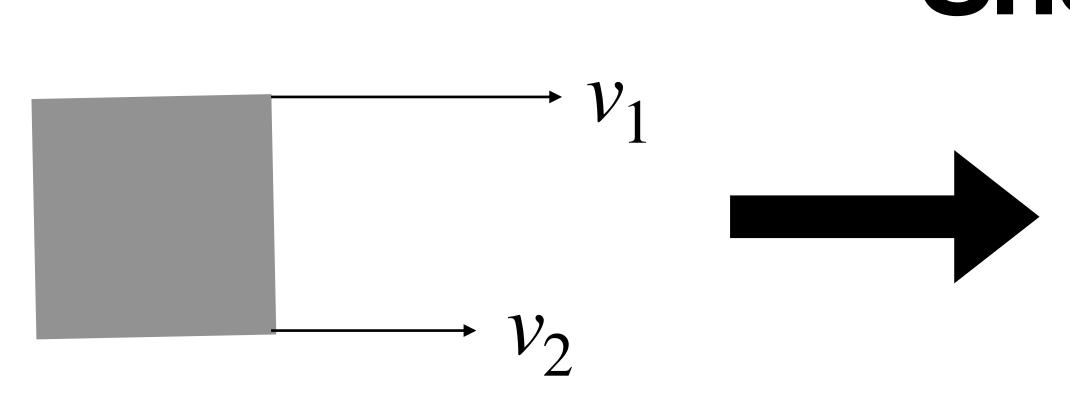
$$\Gamma = \int_{S} (\overrightarrow{\nabla} \times \overrightarrow{v}) \cdot d\overrightarrow{S} = \int_{S} \overrightarrow{\omega} \cdot d\overrightarrow{S}$$

• If the flow is irrotational,  $\overrightarrow{\omega} = 0 \Rightarrow \Gamma = 0$ .



Credit: Wikipedia

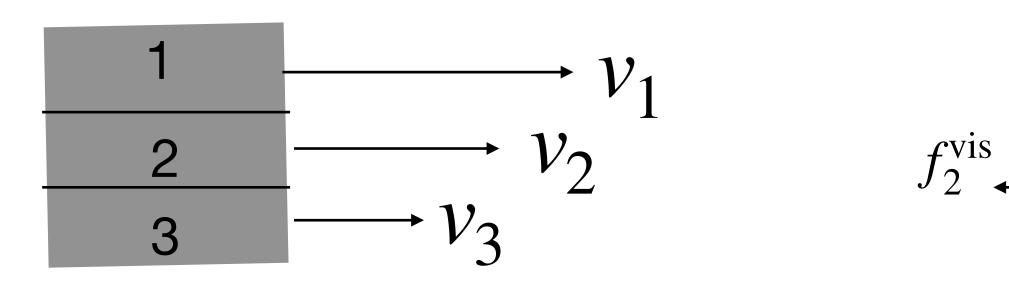




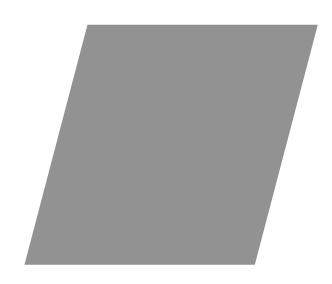
Shearing can occur when neighboring fluid moves with different velocities.

In the presence of viscosity, the shear motion develops a viscous stress that opposes the motion.

The stress acting on a fluid element can be characterized by a stress tensor T.



### Shearing



$$- f_1^{\text{vis}}$$

$$\begin{aligned} & \text{Simple Mod} \\ f_1^{\text{vis}} &= \mu \frac{\partial v_x(x, y + dy/2, z)}{\partial y} dx dz \\ f_2^{\text{vis}} &= -\mu \frac{\partial v_x(x, y - dy/2, z)}{\partial y} dx dz \end{aligned}$$

 $\mu$ : coefficient of shear viscosity

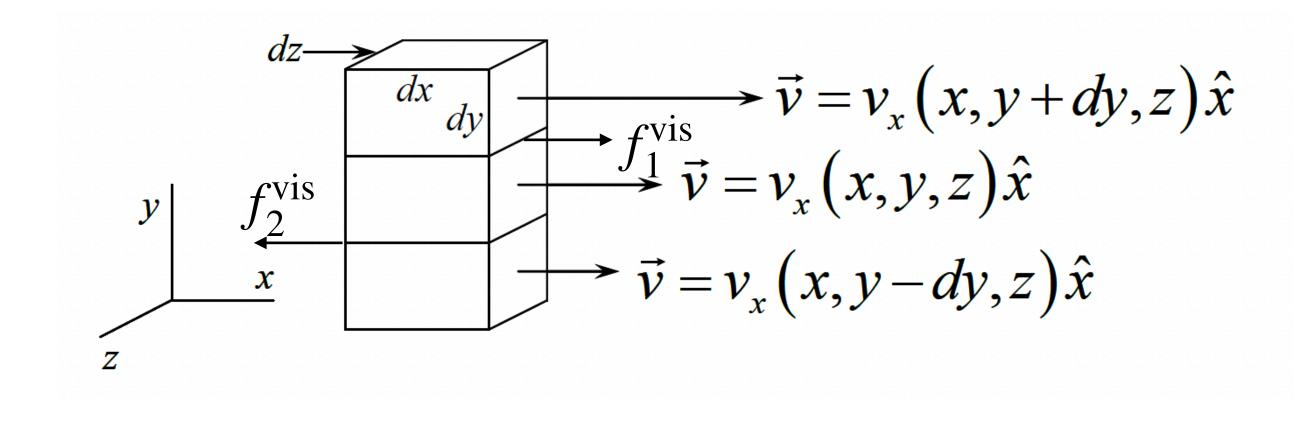
Net force 
$$f_x^{\text{vis}} = f_1^{\text{vis}} + f_2^{\text{vis}} = \mu \frac{\partial^2 v_x}{\partial y^2} dx dy dz =$$

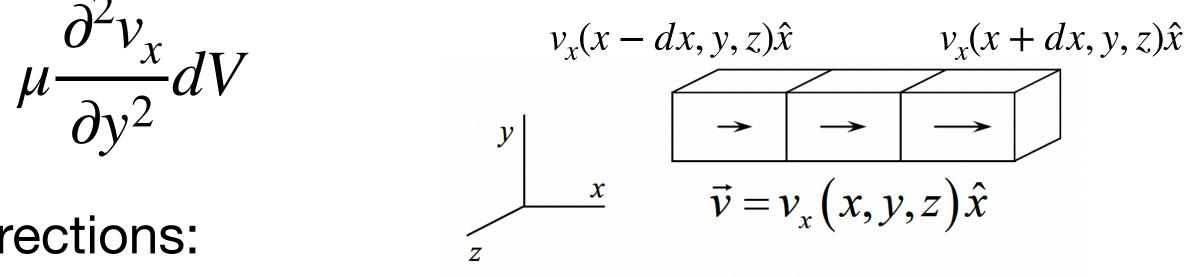
Adding the contributions from the other two directions:  $f_x^{\text{vis}} = \mu \left( \frac{\partial^2 v_x}{\partial x^2} + \frac{\partial^2 v_x}{\partial y^2} + \frac{\partial^2 v_x}{\partial z^2} \right) dV = \mu \nabla^2 v_x dV$ 

The y and z-components of the viscous force are obtained by changing  $v_x$  to  $v_y$  and  $v_z$ .

Viscous force: 
$$\vec{f}^{\text{vis}} = \mu \nabla^2 \vec{v} dV$$

#### del of Viscosity





#### **Stress Tensor**

• Stress tensor can be represented by a  $3 \times 3$  matrix. In Cartesian coordinates,

$$\dot{\overrightarrow{T}} = \begin{pmatrix} T_{xx} & T_{xy} & T_{xz} \\ T_{yx} & T_{yy} & T_{yz} \\ T_{zx} & T_{zy} & T_{zz} \end{pmatrix}$$

• Force acting on a small surface  $d\vec{A} = \hat{n}dA$  is given by

 $d\overrightarrow{F} = \overleftrightarrow{T} \cdot d\overrightarrow{A} = dA(T_{xx}n_x + T_{xy}n_y + T_{xz}n_z)\hat{x} + dA(T_{yx}n_x + T_{yy}n_y + T_{yz}n_z)\hat{y} + dA(T_{zx}n_x + T_{zy}n_y + T_{zz}n_z)\hat{z}$ 

$$= dA \begin{pmatrix} T_{xx} & T_{xy} & T_{xz} \\ T_{yx} & T_{yy} & T_{yz} \\ T_{zx} & T_{zy} & T_{zz} \end{pmatrix} \begin{pmatrix} n_x \\ n_y \\ n_z \end{pmatrix}$$

• It can be shown that  $\overleftarrow{T}$  must be symmetry: T

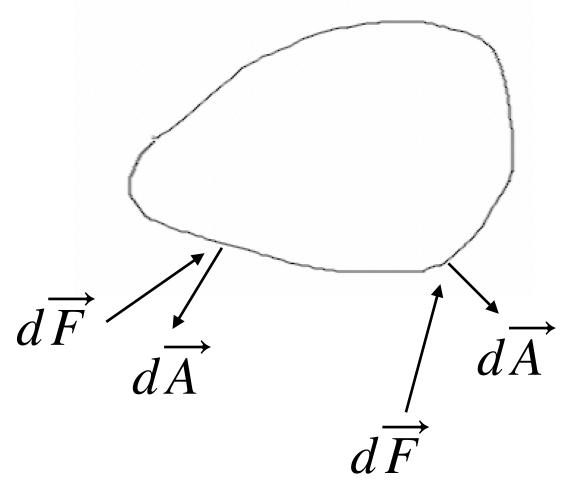
$$T_{ij} = T_{ji}$$

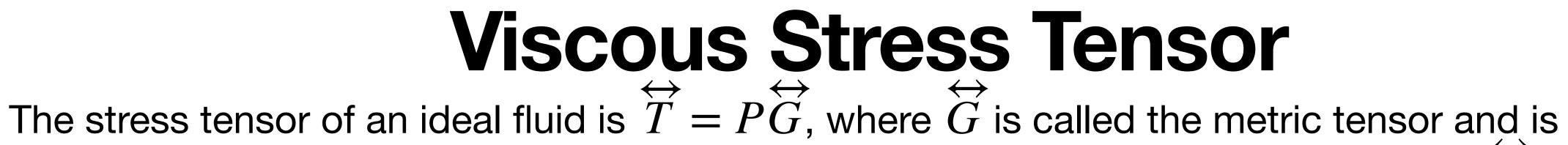
#### Force on Fluid

$$\overrightarrow{F} = -\int_{S} \overleftrightarrow{T} \cdot d\overrightarrow{A}$$

Note the negative sign since  $d\overrightarrow{A}$  points outward. Divergence theorem:

$$\vec{F} = -\int_{V} \vec{\nabla} \cdot \vec{T} dV \qquad d\vec{I}$$
$$\vec{\nabla} \cdot \vec{T} = \left(\frac{\partial T_{xx}}{\partial x} + \frac{\partial T_{yx}}{\partial y} + \frac{\partial T_{zx}}{\partial z}\right) \hat{x} + \left(\frac{\partial T_{xy}}{\partial x} + \frac{\partial T_{yy}}{\partial y} + \frac{\partial T_{zy}}{\partial z}\right) \hat{y} + \left(\frac{\partial T_{xz}}{\partial x} + \frac{\partial T_{yz}}{\partial y} + \frac{\partial T_{zz}}{\partial z}\right) \hat{z}$$
Force per unit volume:  $\vec{f} = -\vec{\nabla} \cdot \vec{T}$ 





represented by a diagonal matrix

$$\overleftrightarrow{T} = \begin{pmatrix} P & 0 & 0 \\ 0 & P & 0 \\ 0 & 0 & P \end{pmatrix}$$

every direction). Force per unit volume is

$$\vec{f} = -\overrightarrow{\nabla}\cdot\overrightarrow{T} = -\frac{\partial P}{\partial x}\hat{x} - \frac{\partial P}{\partial y}\hat{y} - \frac{\partial P}{\partial z}\hat{z} = -$$

In the presence of viscosity,  $\overleftarrow{T} = P\overrightarrow{G} + \overleftarrow{\tau}$ ,  $\overleftarrow{\tau}$  is called the viscous stress tensor.

Viscous force acting on a small ares is  $d\overrightarrow{F}_{vis} = \overleftrightarrow{\tau} \cdot d\overrightarrow{A}$ Viscous force per unit volume is  $\vec{f}_{vis} = -\vec{\nabla}\cdot\vec{\tau}$ 

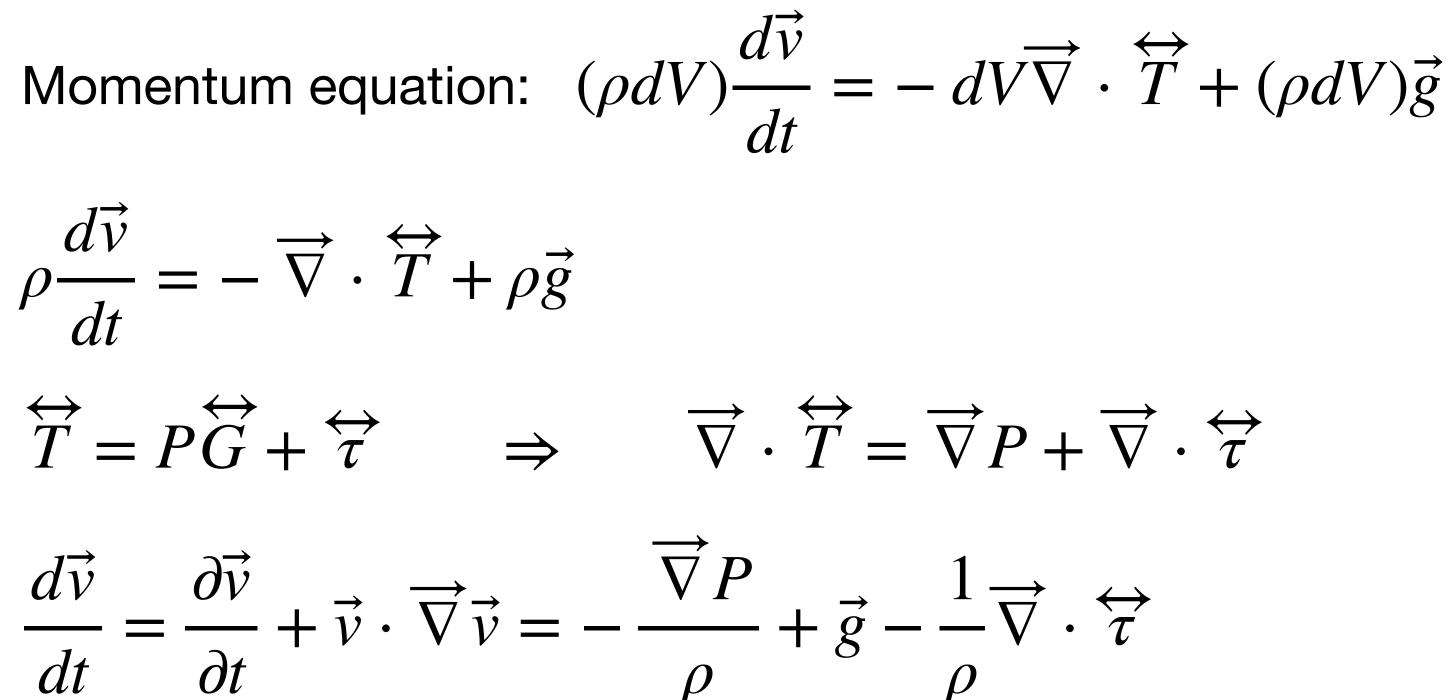
represented by an identity matrix in Cartesian coordinates. In Cartesian coordinates,  $\widetilde{T}$  is

Force acting on a small area is  $d\vec{F} = \vec{T} \cdot d\vec{A} = Pd\vec{A}$ . Force is isotropic (same magnitude in

 $\overrightarrow{\nabla} P$ 

 $d\vec{F} = Pd\vec{A}$ 

#### **Momentum Equation with Viscosity**



Need an expression for  $\overleftarrow{\tau}$  that depends on the velocity field  $\vec{v}$ .

 $\overleftarrow{\tau} \neq 0$  only for non-uniform  $\vec{v}$ , but  $\overleftarrow{\tau} = 0$  if the fluid is rigidly rotating.

## **Velocity Gradient Tensor**

The velocity gradient tensor  $\overrightarrow{\nabla} \overrightarrow{v}$  can be

 $\overleftarrow{\tau}$  is symmetric, but  $\overrightarrow{\nabla} \vec{v}$  is not. Cannot express  $\overleftarrow{\tau}$  in terms of  $\overrightarrow{\nabla} \vec{v}$  directly. Decompose  $\overrightarrow{\nabla} \overrightarrow{v}$  into 3 components: ( Expansion:  $\theta = Tr(\overrightarrow{\nabla} \overrightarrow{v}) = \overrightarrow{\nabla} \cdot \overrightarrow{v}$ Anti-symmetric part of  $\vec{\nabla} \vec{v}$ :  $r_{ij} = \frac{1}{2}$ Symmetric trace-free part of  $\overrightarrow{\nabla} \overrightarrow{v} : \sigma_{ii} =$ 

e represented by a matrix: 
$$\overrightarrow{\nabla} \overrightarrow{v} =$$

$$\vec{\nabla}\vec{v})_{ij} = \frac{\partial v_j}{\partial x_i} = \frac{1}{3}\theta\delta_{ij} + r_{ij} + \sigma_{ij}$$

$$\left(\frac{\partial v_j}{\partial x_i} - \frac{\partial v_i}{\partial x_j}\right)$$
$$= \frac{1}{2} \left(\frac{\partial v_j}{\partial x_i} + \frac{\partial v_i}{\partial x_j}\right) - \frac{1}{3} \theta \delta_{ij}$$

$$= \begin{pmatrix} \frac{\partial v_x}{\partial x} & \frac{\partial v_y}{\partial x} & \frac{\partial v_z}{\partial x} \\ \frac{\partial v_x}{\partial y} & \frac{\partial v_y}{\partial y} & \frac{\partial v_z}{\partial y} \\ \frac{\partial v_x}{\partial z} & \frac{\partial v_y}{\partial z} & \frac{\partial v_z}{\partial z} \end{pmatrix}$$

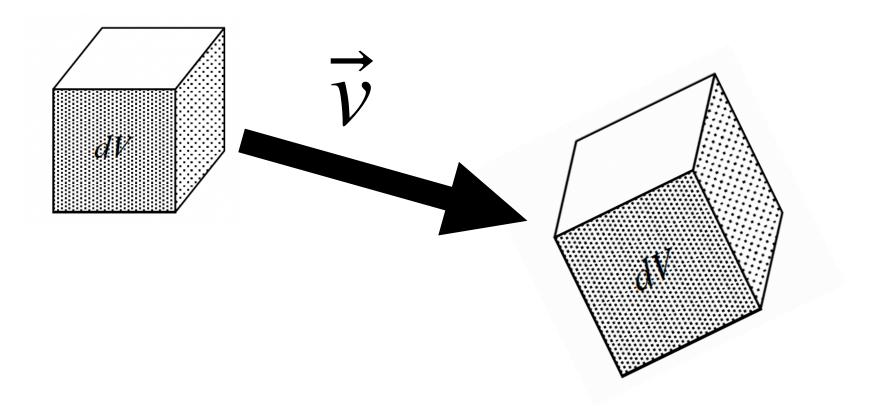
# **Physical Meaning of** $\theta$

Consider a small fluid element occupying a small volume  $\Delta V$  and mass  $\Delta m = \rho \Delta V$ .

Moving with the mass, we have

$$0 = \frac{d}{dt}(\rho\Delta V) = \Delta V \frac{d\rho}{dt} + \rho \frac{d\Delta V}{dt}$$
  
Continuity equation:  $\frac{d\rho}{dt} = -\rho \overrightarrow{\nabla} \cdot \overrightarrow{v} = -\rho\theta$   
 $-\rho\theta\Delta V + \rho \frac{d\Delta V}{dt} = 0$   
 $\theta = \frac{1}{\Delta V} \frac{d\Delta V}{dt}$ 

 $\theta$  is the fractional rate of increase of fluid element's volume.





$$r_{xx} = r_{yy} = r_{zz} = 0$$
,  $r_{xy} = -r_{yx} = \frac{1}{2} \left( \frac{\partial v_y}{\partial x} \right)$ 

Similarly,  $r_{yz} = -r_{zy} = \frac{1}{2}\omega_x$ ,  $r_{zx} = -r_{xz} = \frac{1}{2}\omega_y$ 

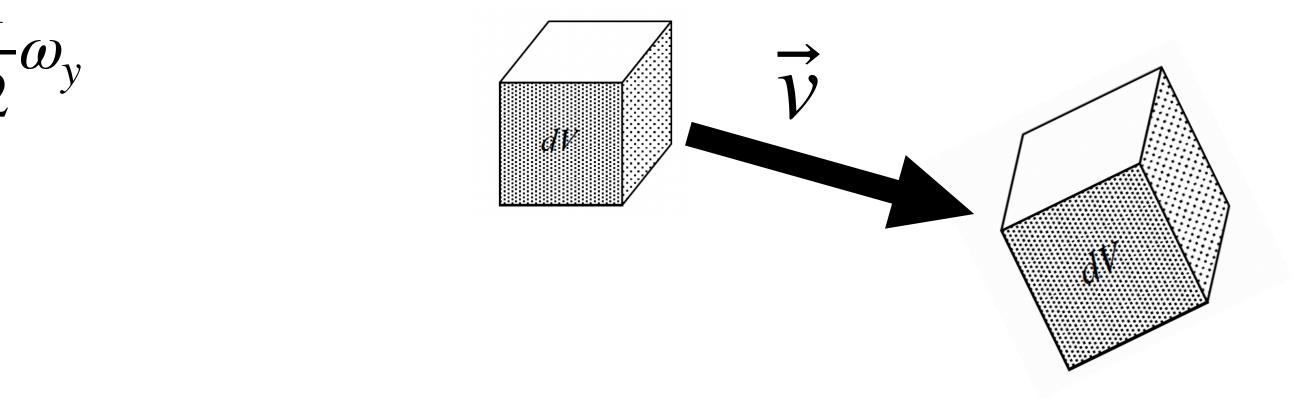
$$\dot{\vec{r}} = \frac{1}{2} \begin{pmatrix} 0 & \omega_z & -\omega_y \\ -\omega_z & 0 & \omega_x \\ \omega_y & -\omega_x & 0 \end{pmatrix}$$

 $\overrightarrow{r}$  describes the local rotation of fluid.  $\overleftrightarrow{\tau}$  is symmetric but  $\overleftrightarrow{r}$  is anti-symmetric.  $\overleftrightarrow{\tau}$  cannot depend on  $\overleftrightarrow{r}$ .

 $\leftrightarrow$ .  $\overleftarrow{\sigma}$  is symmetric and trace-free. It describes the shear motion of fluid.



 $-\frac{\partial v_x}{\partial y}\right) = \frac{1}{2}(\vec{\nabla} \times \vec{v})_z = \frac{1}{2}\omega_z$ 



#### **Bulk and Shear Viscosity**

Simple model of viscosity:  $\overleftarrow{\tau} = -\zeta \theta$  $\tau_{ii} = -\zeta \theta \delta_{ii} - 2\mu \sigma_{ii}$  $\zeta$ : coefficient of bulk viscosity,  $\mu$ : coefficient of shear viscosity. Bulk viscosity resists the fluid's expansion and contraction. Shear viscosity resists the fluid's shear motion. In general, bulk viscosity << shear viscosity. Another quantity is kinematic viscosit

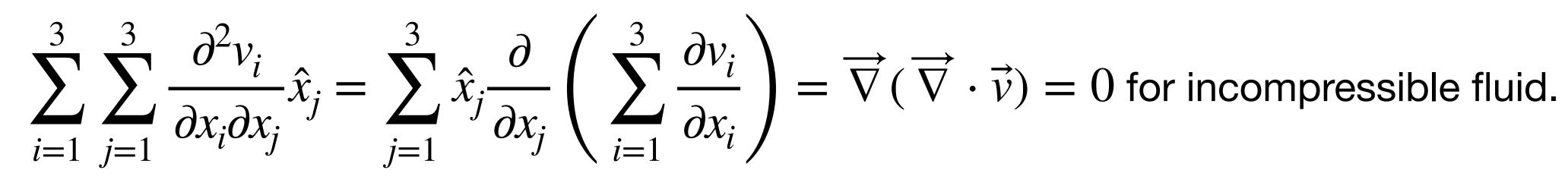
$$\partial \overleftrightarrow{G} - 2\mu \overleftrightarrow{\sigma}$$
 or in component form:

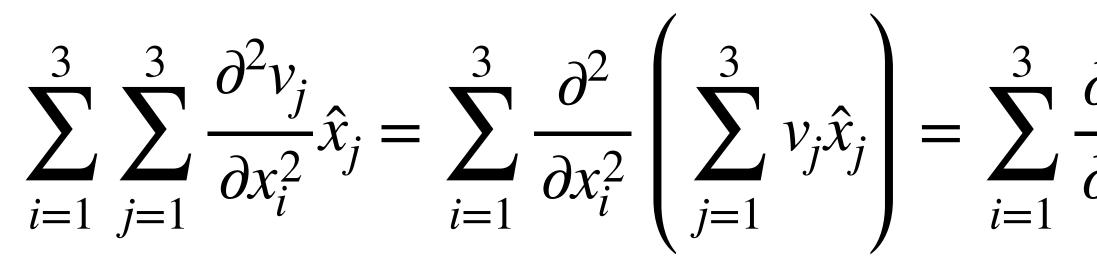
ty 
$$\nu = \mu / \rho$$

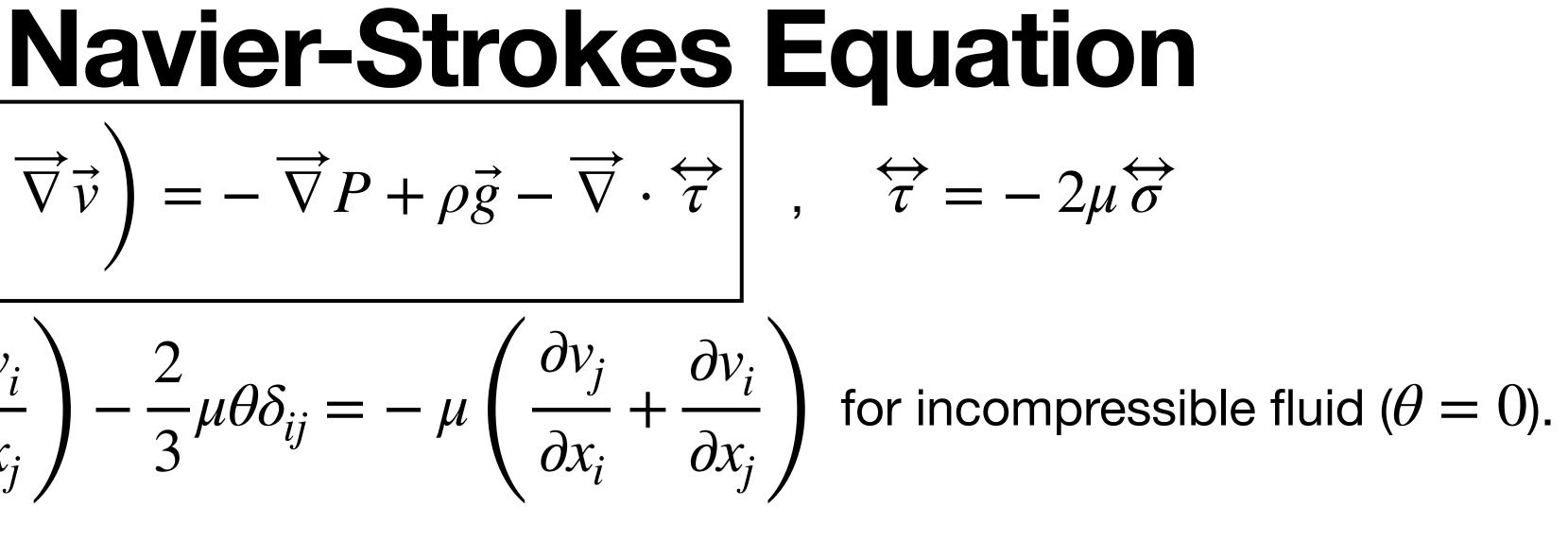
 $\rho \frac{d\vec{v}}{dt} = \rho \left( \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} \right) = -\vec{\nabla} P + \rho \vec{g} - \vec{\nabla} \cdot \overleftarrow{\tau} \quad , \quad \overleftarrow{\tau} = -2\mu \overleftarrow{\sigma}$ 

$$\tau_{ij} = -\mu \left( \frac{\partial v_j}{\partial x_i} + \frac{\partial v_i}{\partial x_j} \right) - \frac{2}{3}\mu \theta \delta_{ij} = -\mu \left( \frac{\partial v_j}{\partial x_i} + \frac{\partial v_j}{\partial x_j} \right)$$

$$\overrightarrow{\nabla} \cdot \overleftarrow{\tau} = \sum_{i=1}^{3} \frac{\partial}{\partial x_i} \left( \sum_{j=1}^{3} \tau_{ij} \widehat{x}_j \right) = -\mu \sum_{i=1}^{3} \sum_{j=1}^{3} \left( \frac{\partial^2 v_i}{\partial x_i \partial x_j} + \frac{\partial^2 v_j}{\partial x_i^2} \right) \widehat{x}_j$$







$$\frac{\partial^2 \vec{v}}{\partial x_i^2} = \nabla^2 \vec{v}$$

#### **Navier-Strokes Equation for Incompressible Fluid**

For incompressible fluid,  $\overrightarrow{\nabla} \cdot \overleftarrow{\tau} = -\mu \nabla^2 \vec{v}$ .

$$\rho \frac{d\vec{v}}{dt} = \rho \left( \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} \right) = -\vec{\nabla} P + \rho \vec{g} + \mu \nabla^2 \vec{v}$$
$$Or$$
$$\frac{d\vec{v}}{dt} = \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v} = -\frac{\vec{\nabla} P}{\rho} + \vec{g} + \nu \nabla^2 \vec{v}$$

 $\nu = \mu / \rho$  : kinematic viscosity

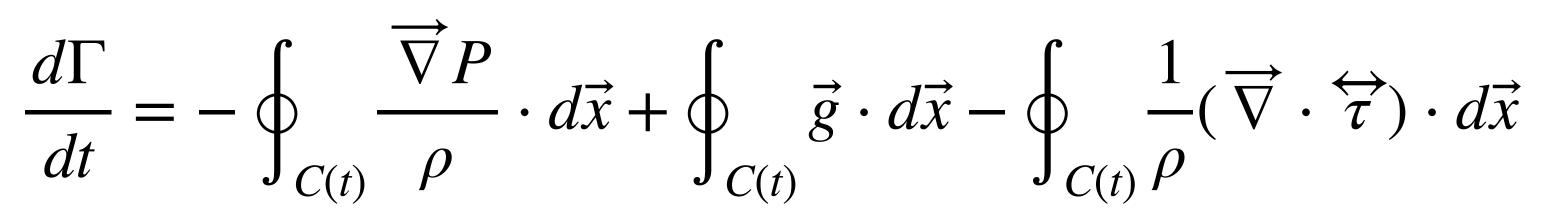


#### **Evolution of Circulation**

Circulation: 
$$\Gamma(t) = \oint_{C(t)} \vec{v} \cdot d\vec{x} = \int_{S(t)} \vec{\omega}$$

Suppose the loop C(t) follows the fluid's motion. Then

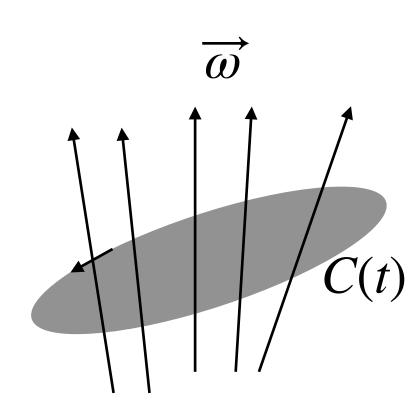
 $\frac{d\Gamma}{dt} = \oint_{C(t)} \frac{d}{dt} (\vec{v} \cdot d\vec{x}) = \oint_{C(t)} \frac{d\vec{v}}{dt} \cdot d\vec{x} +$  $\oint_{C(t)} \vec{v} \cdot d\left(\frac{d\vec{x}}{dt}\right) = \oint_{C(t)} \vec{v} \cdot d\vec{v} = \frac{1}{2} \oint_{C(t)} dv^2 = 0$ Navier-Stokes equation:  $\frac{d\vec{v}}{dt} = -\frac{\vec{\nabla}P}{c} + c$ 



 $\cdot d\vec{S}$ 

$$-\oint_{C(t)} \vec{v} \cdot d\left(\frac{d\vec{x}}{dt}\right)$$

$$\vec{g} - \frac{1}{\rho} \overrightarrow{\nabla} \cdot \overleftarrow{\tau}$$



#### **Kelvin's Circulation Theorem**

$$\begin{split} \oint_{C(t)} \vec{g} \cdot d\vec{x} &= \int_{S(t)} (\vec{\nabla} \times \vec{g}) \cdot d\vec{S} = -\int_{S(t)} (\vec{\nabla} \times \vec{\nabla} U) \cdot d\vec{S} = 0 \\ - \oint_{C(t)} \frac{\vec{\nabla} P}{\rho} \cdot d\vec{x} &= -\int_{S(t)} \left( \vec{\nabla} \times \frac{\vec{\nabla} P}{\rho} \right) \cdot d\vec{S} = \int_{S(t)} \frac{\vec{\nabla} \rho \times \vec{\nabla} P}{\rho^2} \cdot d\vec{S} \\ \frac{d\Gamma}{dt} &= \int_{S(t)} \frac{\vec{\nabla} \rho \times \vec{\nabla} P}{\rho^2} \cdot d\vec{S} - \oint_{C(t)} \frac{1}{\rho} (\vec{\nabla} \cdot \vec{\tau}) \cdot d\vec{x} \end{split}$$

If the fluid is barotropic:  $P = P(\rho), \ \overrightarrow{\nabla} P =$ 

$$\frac{d\Gamma}{dt} = 0$$
 for barotropic, in

$$\frac{dP}{d\rho} \overrightarrow{\nabla} \rho \text{ and so } \overrightarrow{\nabla} \rho \times \overrightarrow{\nabla} P = 0.$$

nviscid flow.



#### Water flowing through Cylindrical Pipe I Continuity equation: $\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{v}) = 0$

In cylindrical coordinates,

$$\frac{\partial \rho}{\partial t} + \frac{\partial (\rho v_r)}{\partial r} + \frac{1}{r} \frac{\partial (\rho v_\theta)}{\partial \theta} + \frac{\partial (\rho v_z)}{\partial z} = 0$$

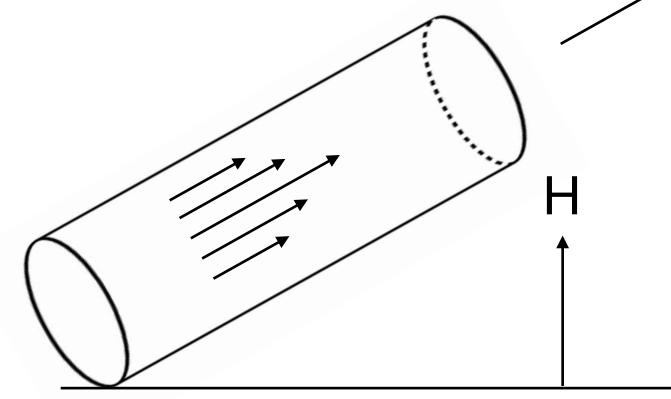
Looking for a steady solution ( $\partial \rho / \partial t = 0$ ), axisymmetric and  $v_r = v_{\theta} = 0$ 

$$\Rightarrow \frac{\partial v_z}{\partial z} = 0 , \Rightarrow v_z = v_z(r)$$

Navier-Stokes equation:

$$\rho\left(\frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \vec{\nabla} \vec{v}\right) = -\vec{\nabla}P + \rho \vec{g} + \mu \nabla$$

Set  $\partial \vec{v} / \partial t = 0$  and write  $P = \rho g H + P_1$ , where H is height from a reference point.

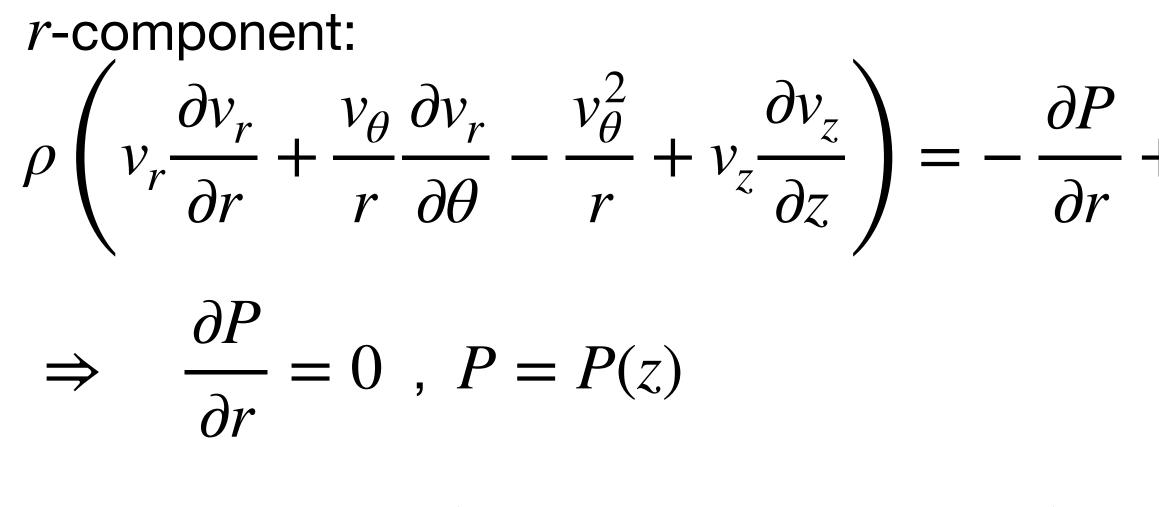




Water flowing through Cylindrical Pipe II  $P = \rho g H + P_1 \implies \overrightarrow{\nabla} P = \rho g \hat{H} + \overrightarrow{\nabla} P_1 = -\rho \vec{g} + \overrightarrow{\nabla} P_1$ 

Navier-Stokes equation becomes  $\rho \vec{v} \cdot \vec{\nabla} \vec{v} = -\vec{\nabla} P_1 + \mu \nabla^2 \vec{v}$ 

(pressure -  $\rho g H$ ).

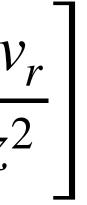


z-component:  $\rho\left(v_r\frac{\partial v_z}{\partial r} + \frac{v_\theta}{r}\frac{\partial v_z}{\partial \theta} + v_z\frac{\partial v_z}{\partial z}\right) =$ 

Gravity is eliminated by the ho gH term. In the following, I will drop the subscript 1. So P means  $P_1$ 

$$+\mu\left[\frac{\partial}{\partial r}\left(\frac{1}{r}\frac{\partial(rv_r)}{\partial r}\right)+\frac{1}{r^2}\frac{\partial^2 v_r}{\partial \theta^2}-\frac{2}{r^2}\frac{\partial v_\theta}{\partial \theta}+\frac{\partial^2 v_r}{\partial z^2}\right]$$

$$= -\frac{\partial P}{\partial z} + \mu \left[ \frac{1}{r} \frac{\partial}{\partial r} \left( r \frac{\partial v_z}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 v_z}{\partial \theta^2} + \frac{\partial^2 v_z}{\partial z^2} \right]$$



$$\frac{dP}{dz} = \frac{\mu}{r} \frac{d}{dr} \left( r \frac{dv_z}{dr} \right)$$

LHS is function of *z*, RHS is function of *r*.

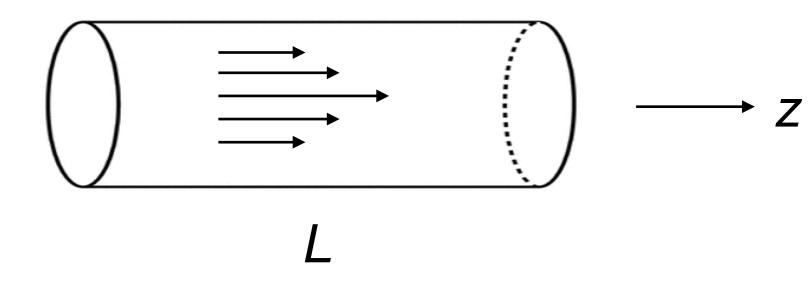
$$\Rightarrow \frac{dP}{dz} = \frac{\mu}{r} \frac{d}{dr} \left( r \frac{dv_z}{dr} \right) = k = \text{constant}$$

Let L be the length of the pipe. Integrating dP/dz = k from z = 0 to z = L gives

two ends of the pipe.

$$\frac{\mu}{r}\frac{d}{dr}\left(r\frac{dv_z}{dr}\right) = -\frac{\Delta P}{L} \implies r\frac{dv_z}{dr} = -\frac{\Delta P}{dr}$$
$$v_z = \int \left(-\frac{\Delta P}{2\mu L}r + \frac{C_1}{r}\right) dr = -\frac{\Delta P}{4\mu L}r^2 + \frac{C_1}{4\mu L}r^2$$

# ugh Cylindrical Pipe III



- $\Delta P = kL$  or  $k = -\Delta P/L$ , where  $\Delta P = P(0) P(L)$  is the pressure difference between the

 $\frac{\Delta P}{\mu L} \int r dr = -\frac{\Delta P}{2\mu L} r^2 + C_1$ 

 $C_1 \ln r + C_2$ 

Water flowing throw  

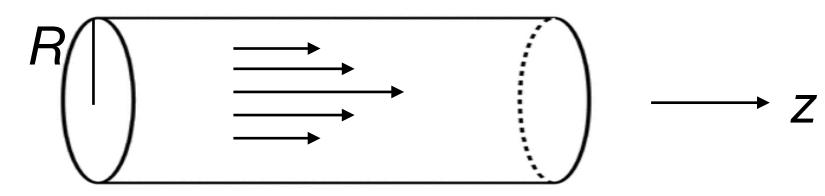
$$v_z(r) = -\frac{\Delta P}{4\mu L}r^2 + C_1 \ln r + C_2$$

Boundary conditions of  $v_7$ : (1) finite at  $r = 0 \Rightarrow C_1 = 0$ , (2)  $v_z = 0$  at the wall at  $r = R \Rightarrow C_2 = \frac{\Delta R}{A_{11}}$ 

$$v_z(r) = \frac{\Delta P}{4\mu L} R^2 \left( 1 - \frac{r^2}{R^2} \right) , \quad v_z(0) = \frac{A}{4\mu L} V_z(r) = \frac{2}{4\mu L} V_z(r) + \frac{2$$

Average flow velocity is  $\langle v_z \rangle = \frac{1}{\pi R^2} \int_0^R \int_0^{2\pi} \frac{\Delta P}{4\mu L} R^2 \left( 1 - \frac{r^2}{R^2} \right) r dr dt$  $\langle v_z \rangle = \frac{\Delta P R^2}{8\mu L} = \frac{1}{2} v_z(0)$ 

# ugh Cylindrical Pipe IV



$$\frac{P}{L}R^2$$

 $\frac{\Delta P}{4\mu L}R^2$ 

$$\theta = \frac{\Delta P}{2\mu L} \int_0^R \left( r - \frac{r^3}{R^2} \right) dr$$

## Water flowing through Cylindrical Pipe V

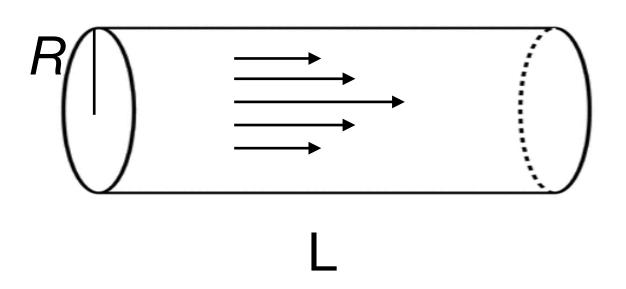
$$v_z(r) = \frac{\Delta P}{4\mu L} R^2 \left( 1 - \frac{r^2}{R^2} \right)$$

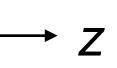
$$\langle v_z \rangle = \frac{\Delta P}{8\mu L} R^2$$

Flow rate:

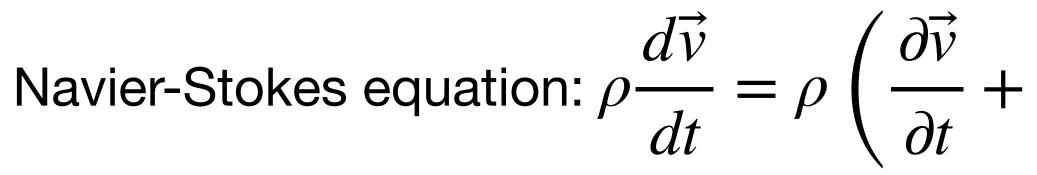
$$Q = \pi R^2 \langle v_z \rangle = \frac{\pi \Delta P R^4}{8\mu L}$$

This is called the Hagen-Poiseuille equation.





#### **Reynolds Number and Turbulence**



inertia	$\rho  d\vec{v}/dt $	$\rho u/T$	<i>ρ</i> и/(
viscosity	$\frac{1}{\mu}  \nabla^2 \vec{v} $	$\sim \mu u/L^2$	$\sim -\mu u$
Reynolds n	umber: Re =	<u>ρuL</u> μ	

L: characteristic length scale, u: characteristic speed. T = L/u: characteristic time.

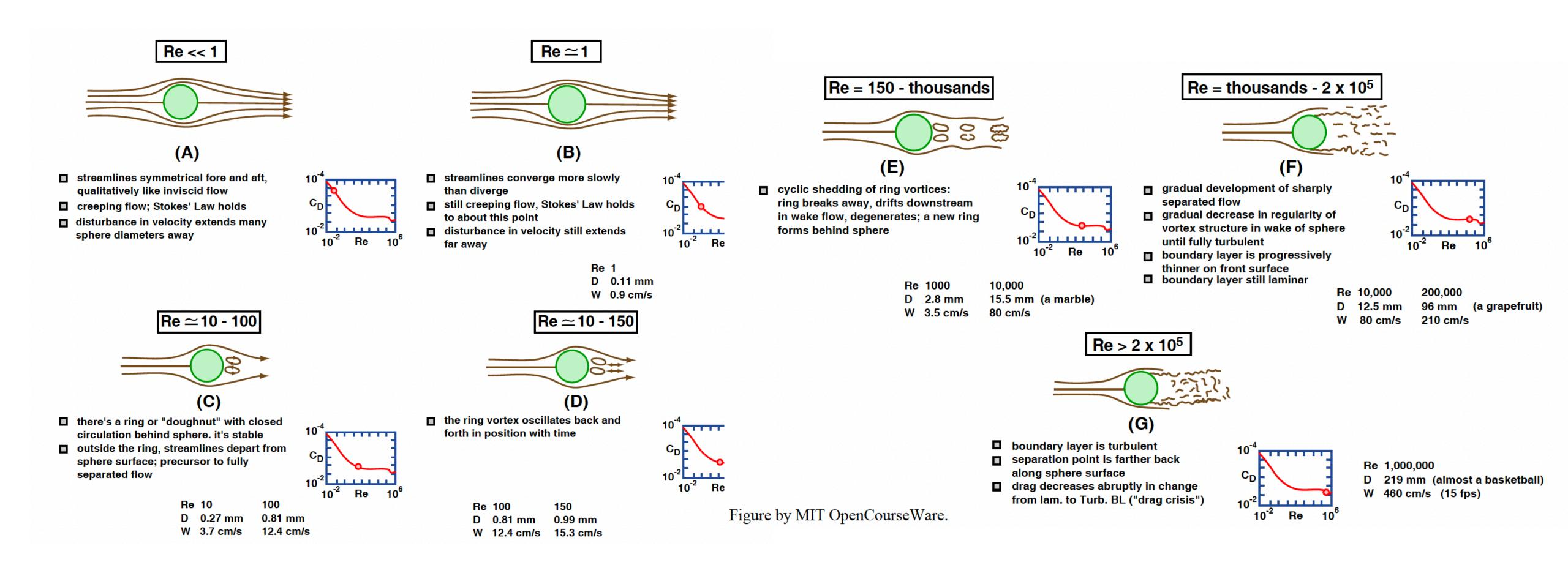
Low Reynolds number  $\rightarrow$  flow dominated by viscosity  $\rightarrow$  laminar

High Reynolds number  $\rightarrow$  flow dominated by inertia  $\rightarrow$  turbulence

Experiments show that pipe flow only remains laminar up to Re  $\sim 10^3 - 10^5$ , depending on the smoothness of pipe's entrance and roughness of its walls.

$$\vec{v} \cdot \vec{\nabla} \vec{v} = -\vec{\nabla} P + \rho \vec{g} + \mu \nabla^2 \vec{v}$$
$$\frac{(L/u)}{/L^2} = \frac{\rho u L}{\mu}$$

#### Flow around Sphere with Different Re's



#### Credit: <u>MIT OpenCourseWare</u>

#### **Darcy's Friction Factor and Head Loss**

Hagen-Poiseuille equation:  $\Delta P = \frac{8\mu L U_{avg}}{R^2} = \frac{32\mu L U_{avg}}{D^2}$ 

Here D = 2R is the pipe diameter,  $U_{avg} = \langle v_z \rangle$  is the average flow velocity in the pipe.

In the absence of viscosity, Bernoulli's equation:

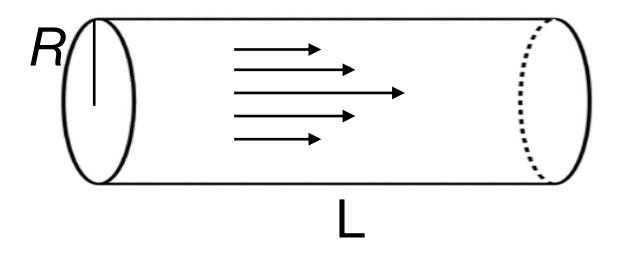
$$\frac{1}{2}\rho v_1^2 + P_1 + \rho g h_1 = \frac{1}{2}\rho v_2^2 + P_2 + \rho g h_2$$

For a horizontal and steady flow,  $\Delta P = P_1 - P_2 = 0$ .

In the presence of viscosity,  $\Delta P \propto L$ . Define a dimensionless parameter called *Darcy's friction factor*:

$$\frac{\Delta P}{L} = f \frac{\frac{1}{2} \rho U_{avg}^2}{D} \quad \text{or} \quad f = \frac{\Delta P}{\frac{1}{2} \rho U_{avg}^2} \left(\frac{D}{L}\right)$$
  
Head loss is defined as  $h_f \equiv \frac{\Delta P}{\rho g} \quad \Rightarrow \quad \left[h_f = f \frac{L U_{avg}^2}{2Dg}\right]$ 





(Darcy-Weisbach equation)



# **Darcy's Friction Factor and Head Loss (cont)**

For pipes with non-circular cross section, *f* and by the *hydraulic diameter*  $D_h \equiv \frac{4A}{P}$ .

A: cross-sectional area of the pipe, P: perimeter of the pipe.

For a duct with rectangular cross section with h

For laminar flow in a cylindrical pipe, Hagen-Poiseuille equation gives

$$f = \frac{64\mu}{\rho U_{avg}D} = \frac{64}{\text{Re}}$$

where the Reynolds number is calculated by  $R\epsilon$ 

In the presence of turbulence, f also depends on the surface roughness of the pipe  $\epsilon$ .

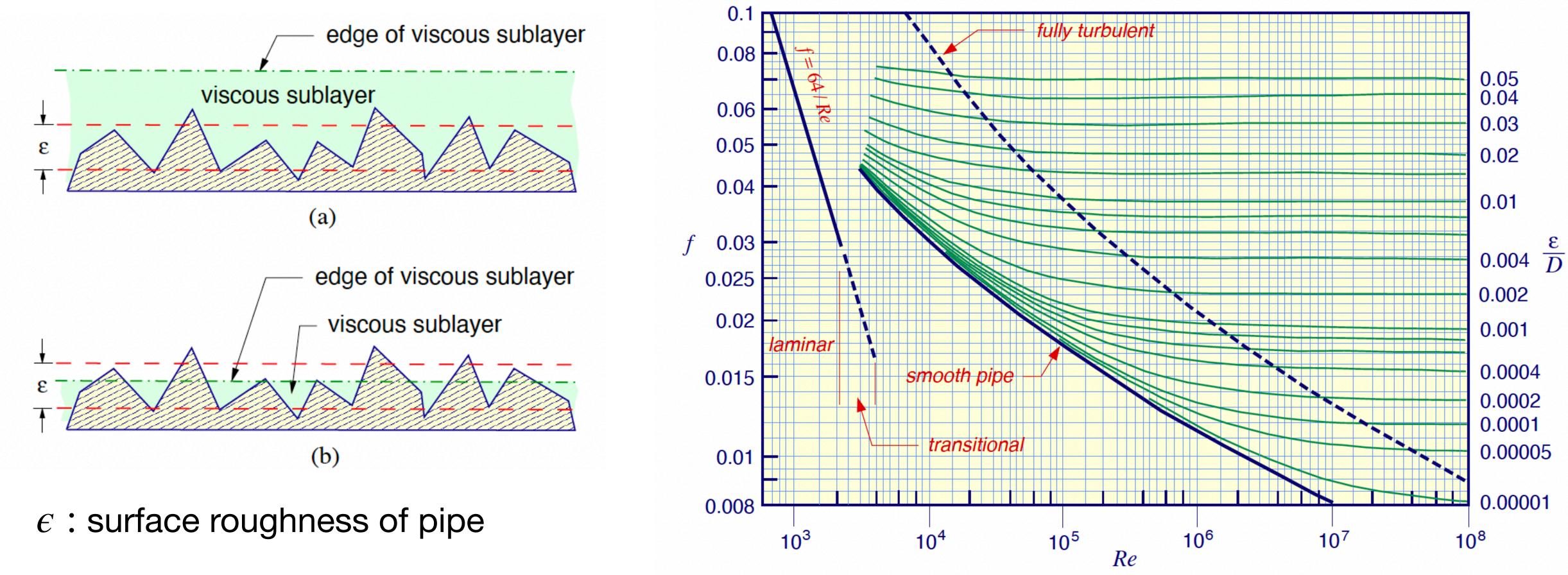
For pipes with non-circular cross section, f and  $h_f$  are defined by replacing the pipe diameter D

height *h* and width *w*, 
$$D_h = \frac{4wh}{2(w+h)}$$

$$e = \frac{\rho U_{avg} D}{\mu}$$



#### **Moody Diagram**



Credit: J.M. McDonough, Lectures In Elementary Fluid Dynamics: Physics, Mathematics and **Applications** 

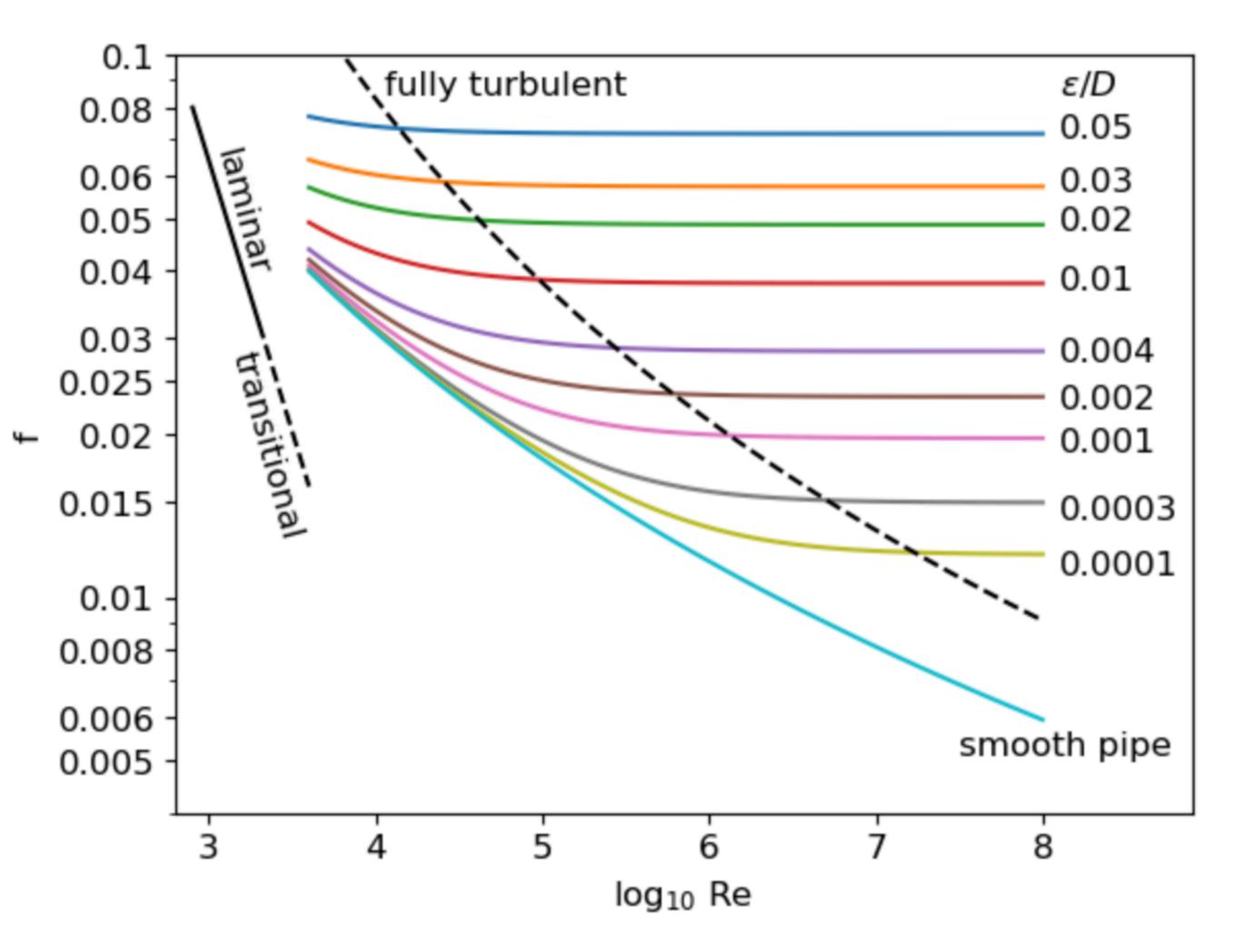
#### **Colebrook Formula**

For  $4 \times 10^3$  < Re <  $10^8$ , Darcy's friction factor may be computed by the Colebrook formula

$$\frac{1}{\sqrt{f}} = -2\log_{10}\left(\frac{\epsilon/D}{3.7} + \frac{2.51}{\text{Re}\sqrt{f}}\right)$$

f needs to be solved iteratively.

The calculated values of *f* differ from experimental results < 15%.



Moody diagram calculated by the Colebrook formula

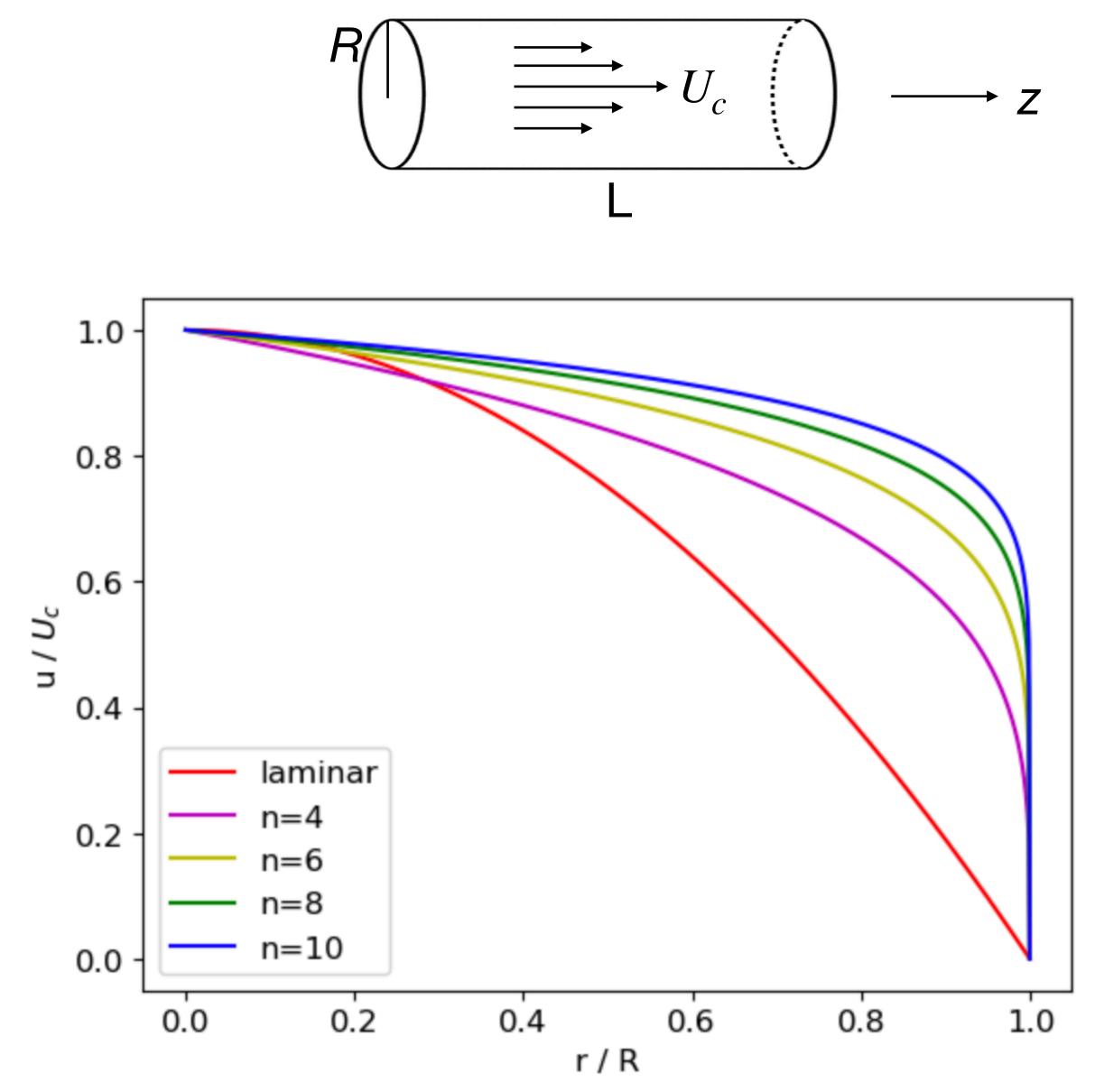


### Velocity Profile

Laminar flow: 
$$u = U_c \left(1 - \frac{r^2}{R^2}\right)$$
  
Turbulent flow:  $u = U_c \left(1 - \frac{r}{R}\right)^{1/n}$ 

$$n = 6$$
 when Re  $\approx 2 \times 10^4$   
 $n = 10$  when Re  $\approx 3 \times 10^6$ 

At high Re, velocity profile is relatively flat, but decreases rapidly to 0 near the wall.



#### **Practical Head Loss Equation**

Bernoulli's equation  $\frac{P_1}{\rho} + \frac{1}{2}v_1^2 + gz_1 = \frac{P_2}{\rho}$ replaced by:

$$\frac{P_1}{\rho g} + \alpha_1 \frac{U_1^2}{2g} + z_1 + h_{pump} = \frac{P_2}{\rho g} + \alpha_2 \frac{U_2^2}{2g} + z_2 + h_f + h_{turbine}$$

 $U_1, U_2$ : average flow speeds,  $\alpha_1, \alpha_2$ : correction factor for KE.

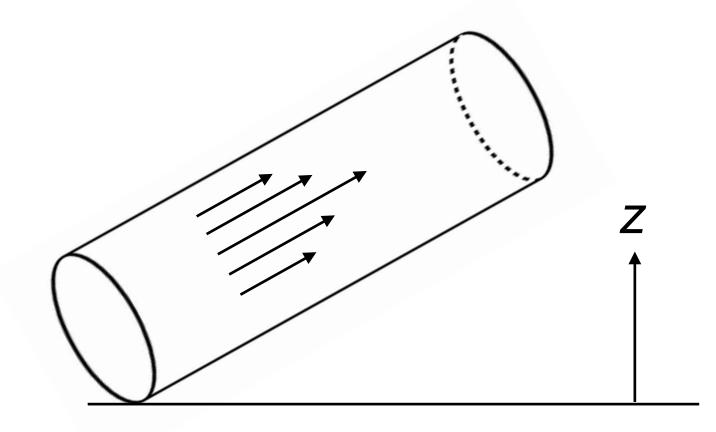
 $\alpha = 2$  for laminar flows,  $\alpha \approx 1$  for turbulent flows.

 $h_f$ : head loss caused by viscosity,

 $h_{pump}$ : head gain by a pump (if present),

 $h_{turbine}$ : head loss by driving a turbine (if present).

$$+\frac{1}{2}v_2^2 + gz_2$$
 is





Oil, with  $\rho = 900 \text{ kg/m}^3$ , and  $\nu = 10^{-5} \text{ m}^2/\text{s}$ , flows at  $Q = 0.2 \text{ m}^3/\text{s}$  through 500 m of 0.2m-diameter cast iron pipe (roughness  $\epsilon = 0.26$  mm). Determine the head loss and pressure drop if the pipe slopes down at  $10^{\circ}$ .

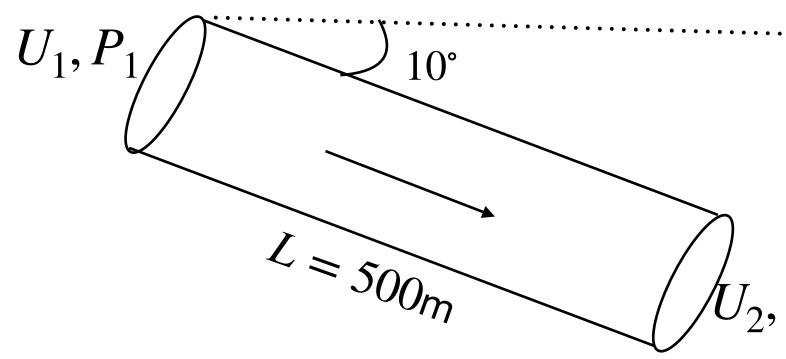
Flow speeds 
$$U_1 = U_2 = \frac{Q}{\pi D^2/4} = 6.37$$
 m/s  
Re  $= \frac{\rho UD}{\mu} = \frac{UD}{\nu} = 1.27 \times 10^5$ 

The flow is turbulent. Using Colebrook formula with  $\epsilon/D = 0.26/200$  and the above Re, I get f = 0.0227. The head loss is given by the Darcy-Weisbach equation:

$$h_f = f \frac{LU^2}{2Dg} = 117 \text{m. } \alpha \approx 1 \text{ for turbulent flows. } \frac{P_1}{\rho g} + \frac{U_1^2}{2g} + z_1 = \frac{P_2}{\rho g} + \frac{U_2^2}{2g} + z_2 + h_f$$
$$\frac{P_1 - P_2}{\rho g} = h_f - (z_1 - z_2) = 117 \text{m} - (500 \text{m}) \sin 10^\circ = 30 \text{m.}$$

Pressure drop  $\Delta P = \rho g(30m) = 2.65 \times 10^5$  Pa.

#### **Example 1**





#### **Example 2**

The pipe in the previous example is connected to a horizontal pipe of length 100 m. The pipe is also made of cast iron but with diameter D = 0.25m. Suppose the flow rate remains the same (Q = 0.2m<sup>3</sup>/s). Calculate the head loss and pressure difference in the second pipe.

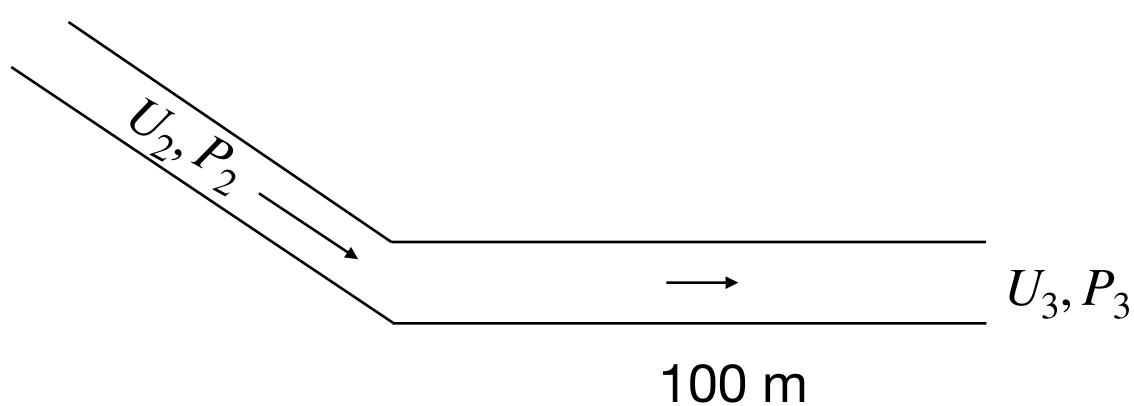
$$U_{3} = \frac{Q}{\pi D^{2}/4} = 4.07 \text{ m/s}$$
  
Re =  $\frac{U_{3}D}{\nu} = 1.02 \times 10^{5}$ ,  $\epsilon/D = 0.26/250$ .

The Colebrook formula gives f = 0.0223.

Head loss: 
$$h_f = f \frac{LU_3^2}{2Dg} = 7.54$$
 m.

Horizontal pipe  $\Rightarrow z_2 = z_3$ ,  $\frac{P_2}{\rho g} + \frac{U_2^2}{2g} = \frac{P_3}{\rho g} + \frac{U_2}{2g}$ 

$$\Rightarrow P_2 - P_3 = \rho g h_f + \rho (U_3^2 - U_2^2)/2 = 5.6 \times 10^4 \,\mathrm{F}$$



$$\frac{V_3^2}{g} + h_f$$
,  $U_2 = 6.37$  m/s from previous calculation.

Pa