

Gauge invariance and spin-1

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Recall our scalar Lagrangian from last time:

$$\mathcal{L}[\Phi] = \partial_\mu \Phi^\dagger \partial^\mu \Phi - m^2 \Phi^\dagger \Phi - \lambda (\Phi^\dagger \Phi)^2$$

We saw that $\delta \Phi = iQ\alpha \Phi$ was a symmetry. What if we let $\alpha = \alpha(x^\mu)$ depend on spacetime position? This is a local transformation because it's a different action at each point, in contrast to global which is the same everywhere.

The spacetime dependence doesn't affect the second and third terms, which remain invariant, but it does change the first one:

$$\begin{aligned} \delta(\partial_\mu \Phi^\dagger \partial^\mu \Phi) &= \partial_\mu \delta \Phi^\dagger \partial^\mu \Phi + \partial_\mu \Phi^\dagger \partial^\mu (\delta \Phi) \\ &= \partial_\mu (-iQ\alpha(x) \Phi^\dagger) \partial^\mu \Phi + \partial_\mu \Phi^\dagger \partial^\mu (iQ\alpha(x) \Phi) \\ &= -iQ \partial_\mu \alpha \Phi^\dagger \partial^\mu \Phi + iQ \partial^\mu \alpha \partial_\mu \Phi^\dagger \Phi \neq 0 \end{aligned}$$

Not invariant anymore!

We can fix this with a trick: swap out all instances of ∂_μ with

$D_\mu \equiv \partial_\mu - igQ A_\mu(x)$ (covariant derivative) where g is called a coupling constant.

We define A_μ to have the transformation rule $A_\mu \rightarrow A_\mu + \frac{1}{g} \partial_\mu \alpha$ for both finite and infinitesimal α

Then $D_\mu \Phi = \partial_\mu \Phi - igQ A_\mu \Phi$ transforms as

$$\begin{aligned} D_\mu \Phi &\rightarrow \partial_\mu (e^{iQ\alpha} \Phi) - igQ (A_\mu + \frac{1}{g} \partial_\mu \alpha) e^{iQ\alpha} \Phi \\ &= \cancel{iQ \partial_\mu \alpha} e^{iQ\alpha} \Phi + e^{iQ\alpha} \partial_\mu \Phi - igQ A_\mu e^{iQ\alpha} \Phi - \cancel{iQ \partial_\mu \alpha} e^{iQ\alpha} \Phi \\ &= e^{iQ\alpha} (\partial_\mu \Phi - igQ A_\mu \Phi) = e^{iQ\alpha} D_\mu \Phi \end{aligned}$$

Transformation of A_μ cancels extra term from derivative of local symmetry parameter

$$\Rightarrow D_\mu \Phi^\dagger D^\mu \Phi \rightarrow (e^{-iQ\alpha} D_\mu \Phi^\dagger) (e^{iQ\alpha} D^\mu \Phi) = D_\mu \Phi^\dagger D^\mu \Phi, \text{ invariant under local symmetry}$$

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So, we can promote a global symmetry $\Phi \rightarrow e^{iQ\alpha} \Phi$ to a local symmetry $\Phi \rightarrow e^{iQ\alpha(x)} \Phi$, at the cost of introducing another field A_μ which has its own non-homogeneous transformation rule $A_\mu \rightarrow A_\mu + \frac{1}{g} \partial_\mu \alpha$.

Why in the world would we do this?

- Turns out this is the correct way to incorporate interactions with spin-1 fields: A_μ will be the photon, and Q is the electric charge. (The coupling constant is $g = \sqrt{4\pi\alpha}$ where $\alpha \approx 1/137$ is the fine-structure constant you saw in Q.M.)
- In fact, this transformation rule for A_μ is required for a consistent, unitary theory of a massless spin-1 particle: invariance under this local transformation is known as gauge invariance.

Let's put Φ aside for now and just consider what form the Lagrangian for A_μ must take.

- Lorentz invariance: A_μ is a Lorentz vector, so $A_\mu(x) \rightarrow \Lambda_\mu^\nu A_\nu(\Lambda^{-1}x)$. So the "principle of contracted indices" holds: $A_\mu A^\mu$ is Lorentz-invariant, as is $(\partial_\mu A_\nu)(\partial^\mu A^\nu)$, etc.
- Gauge invariance: we want \mathcal{L} to be invariant under $A_\mu \rightarrow A_\mu + \frac{1}{g} \partial_\mu \alpha$

Try writing down a mass term:

$$\begin{aligned} \delta\left(\frac{1}{2} m^2 A_\mu A^\mu\right) &= \frac{1}{2} m^2 (\delta A_\mu A^\mu + A_\mu \delta A^\mu) \\ &= \frac{m^2}{g} \partial_\mu \alpha A^\mu \neq 0 \end{aligned}$$

Surprise! A mass term is not allowed by gauge invariance.

What about terms with derivatives? Something like $\partial_\mu A_\nu$ will pick up $\partial_\mu \partial_\nu \alpha$. Can cancel this with a compensating term $\partial_\nu \partial_\mu \alpha$, which comes from $\partial_\nu A_\mu$. This leads to $\mathcal{L}_a = -\frac{1}{4} (\partial_\mu A_\nu - \partial_\nu A_\mu)(\partial^\mu A^\nu - \partial^\nu A^\mu)$

↑
conventional $F_{\mu\nu}$, field strength tensor

With $A_m = (\phi, \vec{A})$, the electromagnetic potentials, you will find that \mathcal{L} is none other than the Maxwell Lagrangian, $\frac{1}{2}(\vec{E}^2 - \vec{B}^2)$.

But the photon has 2 polarizations, i.e. 2 independent components of A_m , which is a 4-vector. How do we get rid of the 2 extraneous components? Two-step process:

1. Note that A^0 has no time derivatives: $\partial_0 A_0$ never appears in Lagrangian, so its equation of motion doesn't involve time. Therefore A_0 is not a propagating degree of freedom: this follows immediately from writing $\langle [F_{\mu\nu}]$. Can solve for A^0 in terms of $\vec{A} \Rightarrow$ 3 components left.

2. Choose a gauge, for example $\vec{\nabla} \cdot \vec{A} = 0$. Solve for one component of \vec{A} in terms of the other two, and what's left are the two propagating degrees of freedom, whose equations of motion are $\square A^{(1,2)} = 0$.

The counting is fairly straightforward as above, but not Lorentz invariant; under a Lorentz transformation, A^0 mixes with \vec{A} , $\vec{\nabla} \cdot \vec{A} = 0$ is not preserved, etc.

Repeat the above analysis using unitary representations of the Lorentz group.

A 4-vector A_m must have some Hilbert space representation $|A_m\rangle$, so we can write a state $|\psi\rangle$ as a linear combination of the components:

$$|\psi\rangle = c_0 |A_0\rangle + c_1 |A_1\rangle + c_2 |A_2\rangle + c_3 |A_3\rangle$$

This state must have positive norm:

$$\langle \psi | \psi \rangle = |c_0|^2 + |c_1|^2 + |c_2|^2 + |c_3|^2 > 0.$$

But if the components of A_m change under a Lorentz transformation, we can change the norm, which is bad; the Lorentz transformation matrices are not unitary!

Alternatively, we could redefine the norm to be Lorentz-invariant,

$$\langle \Psi | \Psi \rangle = |c_0|^2 - |c_1|^2 - |c_2|^2 - |c_3|^2, \text{ but this is not positive-definite!}$$

Solution in two steps: (1) use fields as the representation, which do have unitary (infinite-dimensional) representations, and (2) project out the wrong-sign component. Since vectors live in the $(\frac{1}{2}, \frac{1}{2})$ representation, which has $j=0$ and $j=1$ components, this is equivalent to projecting out the $j=0$ component, leaving $j=1$ as appropriate for spin-1.

Write A_m in Fourier space; $A_m(x) = \int \frac{d^4 p}{(2\pi)^4} E_m(p) e^{-ip \cdot x}$
momentum-dependent polarization vector

A Lorentz transformation will act on this field as

$$A_m(x) \rightarrow \Lambda^{\nu}_{\mu} A_{\nu}(\Lambda^{-1}x) = \int \frac{d^4 p}{(2\pi)^4} \Lambda^{\nu}_{\mu} E_{\nu}(p) e^{-ip \cdot (\Lambda^{-1}x)}$$

polarization vectors rotate, but p_m (a dummy integration variable) does not.

This explains why we pick eigenstates of P^{μ} before defining action of W_m .

Use equations of motion to count independent polarizations:

$$\square A_m - \partial_m(\partial^{\nu} A_{\nu}) = 0 \quad (\text{HW})$$

Choose a gauge such that $\partial^{\nu} A_{\nu} = 0$. (can always do this: if $\partial^{\nu} A_{\nu} = X$, take $A_{\nu} \rightarrow A_{\nu} + \frac{1}{\square} \partial_{\nu} \lambda$, $\partial^{\nu} A_{\nu} \rightarrow X + \frac{1}{\square} \partial^2 \lambda$. Solve for λ to cancel X .)
 \Rightarrow in Fourier space, $p^2 = 0$ and $p \cdot \epsilon = 0$. The latter is an algebraic constraint which is Lorentz-invariant, so it projects out spin-0 as desired. Reduces four polarizations $E_m^{(0)} = (1, 0, 0, 0)$, $E_m^{(1)} = (0, 1, 0, 0)$, ... to three. But we have one more gauge transformation left!

Can still have $A_m = \partial_m \lambda$ consistent with $\partial^{\mu} A_{\mu} = 0$ if $\partial^2 \lambda = 0$.

In this case, A_m is gauge-equivalent to 0 (or pure gauge) and not physical. After Fourier-transforming, this means the polarization proportional to 4-momentum ($E_m \propto p_m$) is unphysical.

We are thus left with two independent polarization vectors:

in a frame where $p_\mu = (E, 0, 0, E)$. (recall this was our "standard vector" for massless particles)

$E_m^{(1)} = (0, 1, 0, 0)$ } linear polarization

$E_m^{(2)} = (0, 0, 1, 0)$

or

$E_m^{(L)} = \frac{1}{\sqrt{2}}(0, 1, -i, 0)$ } circular polarization

$E_m^{(R)} = \frac{1}{\sqrt{2}}(0, 1, i, 0)$

In QFT, these polarization vectors represent physical states, so we can take linear combinations of them:

e.g. $|E\rangle = c_1|1\rangle + c_2|2\rangle$. Define $\langle i|j\rangle = -E_\mu^{(i)} E^{\mu(j)}$

$\langle E|E\rangle = |c_1|^2 \langle 1|1\rangle + |c_2|^2 \langle 2|2\rangle + c_1^* c_2 \langle 1|2\rangle + c_1 c_2^* \langle 2|1\rangle$

$-(E_\mu^{(i)})^* E^{\mu(i)} = 1$

= 0 since $E_\mu^{(1)}$ and $E_\mu^{(2)}$ are orthogonal

$= |c_1|^2 + |c_2|^2$

This inner product is Lorentz-invariant because the basis vectors change under Lorentz, but not $|c|^2$! Moreover, gauge invariance let us get rid of the states with non-positive norm!

$E_\mu^{(0)} = (1, 0, 0, 0) \Rightarrow \langle 0|0\rangle = -1$, bad!

$E_\mu^{(f)} = (1, 0, 0, 1) \Rightarrow \langle f|f\rangle = 0$, unphysical (cancels out of any computation)
(forward, or longitudinal, polarization)

Including the Lagrangian for A_μ , our spin-0 and spin-1 Lagrangian is now

$\mathcal{L} = |D_\mu \Phi|^2 - m^2 \Phi^+ \Phi - \lambda (\Phi^+ \Phi)^2 - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$

Note: $[A_\mu] = [D_\mu] = 1$ from covariant derivative, so $[F_{\mu\nu} F^{\mu\nu}] = 4$, as required.

The derivative term in the Lagrangian for Φ with only the global symmetry, $\partial_\mu \Phi^\dagger \partial^\mu \Phi$, gave rise to the equations of motion for non-interacting (free) scalar fields. Once promoted to a covariant derivative, $|D_\mu \Phi|^2$ contains interactions between Φ and A_μ .

$$|D_\mu \Phi|^2 = (\partial_\mu \Phi^\dagger + igQ A_\mu \Phi^\dagger)(\partial^\mu \Phi - igQ A^\mu \Phi)$$

$$= \partial_\mu \Phi^\dagger \partial^\mu \Phi - A_\mu \underbrace{(-igQ(\Phi^\dagger \partial^\mu \Phi - \partial^\mu \Phi^\dagger \Phi))}_{\text{in QM, this would be the probability current for the wavefunction. In QFT, it's literally the electric current for a charged scalar particle.}}$$

in QM, this would be the probability current for the wavefunction. In QFT, it's literally the electric current for a charged scalar particle.

$\Rightarrow \mathcal{L}$ contains $-\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - A_\mu J^\mu$, which is exactly how you would write Maxwell's equations with an external source $J^\mu = (\rho, \vec{J})$! So Φ sources currents, which create \vec{E} and \vec{B} fields from A_μ , which back-reacts on Φ . These coupled equations are impossible to solve exactly, so starting in 2 weeks we will use perturbation theory in the coupling strength gQ to approximate the solutions.

Massive spin-1 fields

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As we saw, a mass term for a vector field is not gauge invariant. However, there are several massive spin-1 particles in nature, which are either composite particles (the ρ meson, for example) or which acquire a mass through the Higgs mechanism (the W and Z gauge bosons). So, we should understand what their Lagrangians should look like without assuming any gauge invariance conditions.

Luckily, the story is still quite simple. We still need to get rid of 1 extraneous degree of freedom, and this will restrict the form of the Lagrangian.

We want a Lagrangian whose equations of motion will yield $(\square + m^2)A_\mu = 0$ in order to satisfy the relativistic dispersion $p^2 = m^2$. So we can have quadratic terms with 0 or 2 derivatives. The most general such Lagrangian is

$$\mathcal{L} = \frac{a}{2} A^\mu \square A_\mu + \frac{b}{2} A^\mu \partial_\mu \partial^\nu A_\nu + \frac{1}{2} m^2 A^\mu A_\mu$$
 with a, b, m arbitrary coefficients. (Note that $[\mathcal{L}] = 4$ if $[A] = 1$, a and b are dimensionless, and $[m] = 1$.)

The equations of motion are **[HW]**

$$a \square A_\mu + b \partial_\mu \partial^\nu A_\nu + m^2 A_\mu = 0.$$

Take ∂^μ of this to get

$$((a+b)\square + m^2)(\partial^\mu A_\mu) = 0.$$

We are on the right track if we can enforce $\partial^\mu A_\mu = 0$: this is a scalar (i.e. spin-0) constraint so it projects out $j=0$ as desired. To do this, take $a=1, b=-1$:

$$\begin{aligned}
\mathcal{L} &= \frac{1}{2} A^\mu \square A_\mu - \frac{1}{2} A^\mu \partial_\mu \partial_\nu A^\nu + \frac{1}{2} m^2 A^\mu A_\mu \\
&= -\frac{1}{2} (\partial^\nu A^\mu \partial_\nu A_\mu - \partial^\nu A^\mu \partial_\mu A_\nu) + \frac{1}{2} m^2 A^\mu A_\mu \quad (\text{integrating by parts}) \\
&= -\frac{1}{4} (\partial_\mu A_\nu - \partial_\nu A_\mu)(\partial^\mu A^\nu - \partial^\nu A^\mu) + \frac{1}{2} m^2 A^\mu A_\mu \quad (\text{rearranging}) \\
&= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} m^2 A^\mu A_\mu \quad \leftarrow \text{Proca (massive spin-1) Lagrangian}
\end{aligned}$$

The field strength $F_{\mu\nu}$ just appeared without having to invoke gauge invariance! The equations of motion are now

$$(\square + m^2)A_\mu = 0 \quad \text{and} \quad \partial^\mu A_\mu = 0.$$

We can now find the 3 linearly-independent polarization vectors as before, but now in a frame where $p^\mu = (m, 0, 0, 0)$

Since the Poincaré Casimir $p^2 = m^2$.

In Fourier space, have $p^2 = m^2$ and $p \cdot \epsilon = 0$. So can take

$$\epsilon_\mu^1 = (0, 1, 0, 0), \quad \epsilon_\mu^2 = (0, 0, 1, 0), \quad \text{and} \quad \epsilon_\mu^3 = (0, 0, 0, 1).$$

These satisfy $\epsilon^{\mu\nu} \cdot \epsilon_\nu = -1$ as did the massless polarizations, and they are all physical.

In a boosted frame with $p^\mu = (E, 0, 0, p_z)$ ($p_z^2 = E^2 - m^2$),

we have

$$\epsilon_\mu^1 = (0, 1, 0, 0), \quad \epsilon_\mu^2 = (0, 0, 1, 0), \quad \epsilon_\mu^3 = \left(\frac{p_z}{m}, 0, 0, \frac{E}{m} \right).$$

The third polarization is called longitudinal because it has a spatial component along the direction of motion.

Note that for ultra-relativistic energies $E \gg m$,

$$\epsilon_\mu^3 \rightarrow \frac{E}{m} (1, 0, 0, 1).$$

This will cause problems in QFT, and is why massive spin-1 must either be composite or arise from a Higgs mechanism.