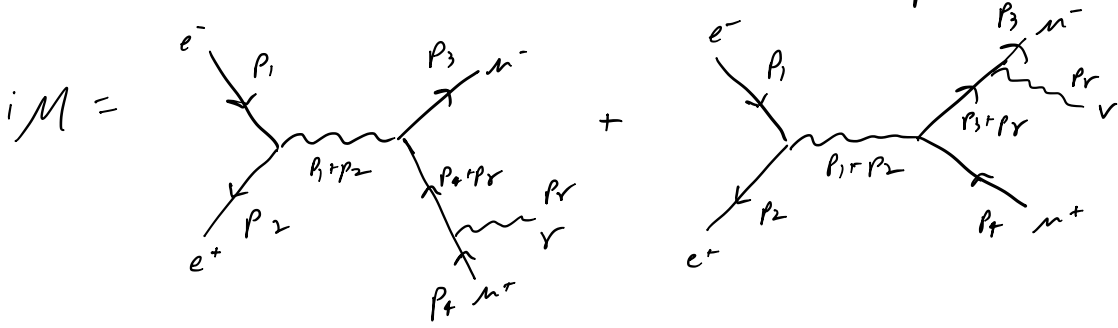


Photon emission:  $e^+e^- \rightarrow m^+m^-\gamma$

We now consider an  $\mathcal{O}(\alpha)$  correction to the process we studied last week.



Assume  $Q^2 = (p_1+p_2)^2 \gg m_m^2$  so we can ignore  $m_e, m_m$ .

$$iM = i \frac{e^2}{Q^2} \bar{v}(p_2) \gamma^\mu u(p_1) \bar{u}(p_3) \left[ \gamma_\mu \frac{-i(\not{p}_4 + \not{p}_r)}{(p_4+p_r)^2} (-ie\gamma^\alpha) + (-ie\gamma^\alpha) \frac{i(\not{p}_3 + \not{p}_r)}{(p_3+p_r)^2} \gamma_\mu \right] v(p_4) \epsilon_\alpha^\beta(p_r)$$

internal fermion propagators defined with momentum along arrows, so need a minus sign here

$$\text{Let } S^{\mu\alpha} = -ie \left[ \gamma^\alpha \frac{i(\not{p}_3 + \not{p}_r)}{(p_3+p_r)^2} \gamma_\mu - \gamma_\mu \frac{i(\not{p}_4 + \not{p}_r)}{(p_4+p_r)^2} \gamma^\alpha \right]$$

(yes I know the index heights are wrong, but  $S^{\mu\alpha} \neq S^{\alpha\mu}$ , so this lets us keep track of order better)

Cross section after averaging over initial and summing over final spins is

$$\sigma_r = \frac{1}{2Q^2} \int d\pi \langle |M|^2 \rangle = \frac{e^4}{2Q^6} L^{\mu\nu} X_{\mu\nu}$$

$\frac{1}{2} \rightarrow \frac{1}{2} \times 2$      $d\pi = E_1 + E_2$     Contains momentum-conserving  $\delta$ -function

$L^{\mu\nu}$  is left half of the diagram:

$$L^{\mu\nu} = \frac{1}{4} \sum_{s_1, s_2} \bar{v}_{s_2}(p_2) \gamma^\mu u_{s_1}(p_1) \bar{u}_{s_1}(p_1) \gamma^\nu v_{s_2}(p_2) = \frac{1}{4} \text{Tr}[\not{p}_2 \gamma^\mu \not{p}_1 \gamma^\nu] = p_1^\mu p_2^\nu + p_1^\nu p_2^\mu - \frac{1}{2} Q^2 g^{\mu\nu}$$

$X_{\mu\nu}$  is right half, involving the photon:

$(\bar{u} \Gamma v)^\dagger = \bar{v} \tilde{\Gamma} u$  where  $\tilde{\Gamma}$  reverses order of  $\gamma$  matrices: check this using  $(\gamma^\mu)^\dagger = \gamma^0 \gamma^\mu \gamma^0$

$$X_{\mu\nu} = \int d\pi \sum_{\substack{s_3, s_4 \\ \text{pols.}}} [\bar{u}_{s_3}(p_3) S^{\mu\alpha} v_{s_4}(p_4) \bar{v}_{s_4}(p_4) S^{\beta\nu} u_{s_3}(p_3) \epsilon_\alpha^\beta(p_r) \epsilon_\rho^\sigma(p_r)]$$

Use  $\sum_{\text{pols.}} \epsilon_\alpha^\beta(p_r) \epsilon_\rho^\sigma(p_r) \rightarrow -\delta_\alpha^\rho$  :  $X_{\mu\nu} = - \int d\pi \text{Tr}[\not{p}_3 S^{\mu\alpha} \not{p}_4 S^{\alpha\nu}]$

Here, we are integrating over 3-body phase space,

$$d\mathcal{T}_3 = \frac{d^3p_3}{(2\pi)^3} \frac{d^3p_4}{(2\pi)^3} \frac{d^3p_r}{(2\pi)^3} \frac{1}{2E_3} \frac{1}{2E_4} \frac{1}{2E_r} (2\pi)^4 \delta(Q - p_3 - p_4 - p_r)$$

where  $Q = p_1 + p_2$

By the Ward identity, we know  $Q_\mu X^{\mu\nu} = 0$ . After  $\int d\mathcal{T}_3$ ,  $X$  is a function of  $Q$  only, symmetric in  $\mu \rightarrow \nu$  (because  $L^{\mu\nu}$  is),

$$\text{so } X_{\mu\nu} = (Q_\mu Q_\nu - Q^2 \eta_{\mu\nu}) X(Q^2)$$

only symmetric tensors  
built out of  $Q$       scalar function  
of  $Q^2$

In this form,  $\eta^{\mu\nu} X_{\mu\nu} = (Q^2 - 4Q^2) X(Q^2)$ , so  $X(Q^2) = -\frac{1}{3Q^2} \eta^{\mu\nu} X_{\mu\nu}$

Plug in for  $L^{\mu\nu}$ :  $L^{\mu\nu} X_{\mu\nu} = (p_1^\mu p_2^\nu + p_1^\nu p_2^\mu - \frac{1}{2} Q^2 \eta^{\mu\nu}) (Q_\mu Q_\nu - Q^2 \eta_{\mu\nu}) X(Q^2)$   
 $= (2(p_1 \cdot Q)(p_2 \cdot Q) - \frac{1}{2} Q^4 - 2Q^2(p_1 \cdot p_2) + 2Q^4) X(Q^2)$

Now,  $Q^2 = (p_1 + p_2)^2 = 2p_1 \cdot p_2$  (assuming  $m_e = 0$ ), and similarly,

$$p_1 \cdot Q = p_1 \cdot (p_1 + p_2) = p_1 \cdot p_2 = \frac{Q^2}{2} = p_2 \cdot Q$$

$$L^{\mu\nu} X_{\mu\nu} = (2(\frac{Q^2}{2})(\frac{Q^2}{2}) - \frac{1}{2} Q^4 - Q^4 + 2Q^4) X(Q^2) = Q^2 X(Q^2) = -\frac{Q^2}{3} \eta^{\mu\nu} X_{\mu\nu}$$

$$\Rightarrow \sigma_r = \frac{e^4}{2Q^6} L^{\mu\nu} X_{\mu\nu} = -\frac{e^4}{6Q^4} \eta^{\mu\nu} X_{\mu\nu}$$

Recall from last week that  $\frac{d\sigma_{e^+e^- \rightarrow \mu^+\mu^-}}{d\theta} = \frac{e^4}{32\pi Q^2} (1 + \cos^2\theta)$ , so integrating over  $\theta$ ,

$$\sigma_0 \equiv \sigma_{e^+e^- \rightarrow \mu^+\mu^-} = \frac{e^4}{12\pi Q^2}$$

Thus we can write  $\sigma_r = \sigma_0 \left( \frac{-2\pi}{Q^2} \eta^{\mu\nu} X_{\mu\nu} \right)$ .

There is a nice way to interpret this result. Let's write

$$\sigma_\gamma = \frac{4\pi\alpha_0}{Q} \times \frac{1}{2Q} x^{\mu\nu} (-\eta_{\mu\nu}), \text{ where } Q = \sqrt{Q^2}. \text{ The decay rate}$$

of a particle of mass  $m$  is given by  $\Gamma = \frac{1}{2M} \int d\pi \langle |M|^2 \rangle$ .

So we can interpret the rate for  $e^+e^- \rightarrow \mu^+\mu^-\gamma$  as the product of the rate for  $e^+e^- \rightarrow \gamma^*$ , a virtual photon of "mass"  $Q$ , times the decay rate of that photon,  $\gamma^* \rightarrow \mu^+\mu^-\gamma$ , summed over polarizations of the final-state photon. This is a special case of the narrow-width approximation, which is a general statement about the factorization of Feynman diagrams through an intermediate state. (More on this in HW 6!)

Let's parameterize the phase space of  $\gamma^* \rightarrow \mu^+\mu^-\gamma$  using Mandelstam

$$\text{Variables as } s = (p_3 + p_4)^2 \equiv Q^2(1-x_\gamma)$$

$$t = (p_3 + p_\gamma)^2 \equiv Q^2(1-x_1)$$

$$u = (p_4 + p_\gamma)^2 \equiv Q^2(1-x_2)$$

From HW 4,  $s+t+u = \sum m_i^2 \approx Q^2$  (you derived it for  $p_1+p_2 \rightarrow p_3+p_4$ , but a similar result holds with appropriate minus signs for  $Q \rightarrow p_3+p_4+\gamma$ )

$\Rightarrow x_\gamma + x_1 + x_2 = 2$ , take  $x_\gamma = 2 - x_1 - x_2$  so  $x_1$  and  $x_2$  are independent.

Limits of integration:  $t = 2p_3 \cdot p_\gamma = 2E_3 E_\gamma (1 - \cos\theta_{3\gamma})$ .  $t_{\min} = 0$  when  $E_\gamma = 0$ ;

$t_{\max} = 4E_3 E_\gamma$  when  $\cos\theta_{3\gamma} = -1$ . If  $E_4 = 0$ ,  $E_3 = E_\gamma = \frac{Q}{2}$ , so  $t_{\max} = Q^2$

$$\Rightarrow x_{1,\min} = 0, x_{1,\max} = 1$$

$$\int d\pi = \frac{Q^2}{128\pi^3} \int_0^1 dx_1 \int_{1-x_1}^1 dx_2$$

$$\text{Tr}[\not{p}_3 \not{S}^{\mu\alpha} \not{p}_4 \not{S}^{\nu\beta}] \eta_{\mu\nu} = \frac{8e^2(x_1^2 + x_2^2)}{(1-x_1)(1-x_2)}$$

(HW)

Trace tricks:  
Schwartz appendix A.4  
Peskin & Schroeder Section 5.1  
You will need these!

This diverges logarithmically ( $\int \frac{1}{x} dx$ ) at  $x_1, x_2 = 1$ .

By the analysis above,  $x_1 = 1$  corresponds to  $2E_3 E_r (1 - \cos \theta_{3r}) = 0$ .

This can happen either if  $E_r = 0$  (a soft singularity), or  $\cos \theta_{3r} = 0$  (a collinear singularity). This behavior is generic in QFT: massless particles prefer to be emitted with low energies and along the directions of charged particles.

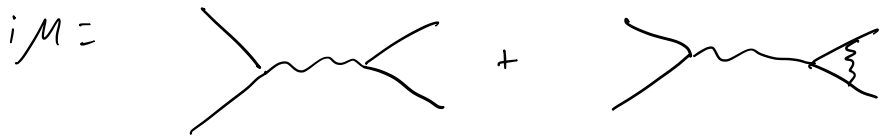
If we pretend that the photon has a mass  $m_\gamma$ , and let  $\beta = \frac{m_\gamma^2}{Q^2}$ ,

the limits of integration change to  $\int d\pi = \int_0^{1-\beta} dx_1 \int_{1-x_1-\beta}^{1-\frac{\beta}{1-x_1}} dx_2$  (HW)

Doing the integral,  $\int_0^{1-\beta} dx_1 \int_{1-x_1-\beta}^{1-\frac{\beta}{1-x_1}} dx_2 \frac{x_1^2 + x_2^2}{(1-x_1)(1-x_2)} = \ln^2 \beta + 3 \ln \beta - \frac{\pi^2}{3} + 6$

double  
logarithms  
from  $x_1$  and  $x_2$

However, these singularities are not physical! It turns out they cancel exactly against the interference terms from



The result is  $6 - \frac{\pi^2}{3} \rightarrow \frac{3}{2}$  for the finite pieces, so

$$\Gamma(\gamma^0 \rightarrow \mu^+ \mu^- \gamma) = \frac{e^2}{2Q} \frac{Q^2}{128\pi^3} \left(8 \times \frac{3}{2}\right) = \frac{3Qe^2}{64\pi^3}$$

$$\sigma_\gamma = \frac{4\pi\sigma_0}{Q} \left(\frac{3Qe^2}{64\pi^3}\right) = \sigma_0 \left(\frac{3e^2}{16\pi^2}\right) \leftarrow \text{quantum correction to } \mu^+ \mu^- \text{ production cross section.}$$

What this result tells us is that  $e^+e^- \rightarrow \mu^+ \mu^-$  with an arbitrarily low energy photon, or one emitted along one of the muon directions, is indistinguishable from just  $\mu^+ \mu^-$  in the final state.

Charged particles are accompanied by clouds of photons.

More concrete interpretation: any real experiment will have a finite energy resolution  $E_{res}$  and angular resolution  $\theta_{res}$ . Instead of cutting off the integral with  $m_V$ , use  $E_{res}$  and  $\theta_{res}$  instead.

This is technically complicated, so we will just quote the answer:

$$\sigma(e^+e^- \rightarrow \mu^+\mu^- \gamma) \Big|_{\substack{E_V > E_{res} \\ \theta_{V\mu} > \theta_{res}}} = \sigma_0 \frac{e^2}{8\pi^2} \left( \ln \frac{1}{\theta_{res}} \left[ \ln \left( \frac{Q}{2E_{res}} - 1 \right) + \dots \right] + \dots \right)$$

Focus on  $\ln \frac{Q}{2E_{res}}$ . If  $Q \gg E_{res}$ , could be in a situation where

$$\ln \left( \frac{Q}{2E_{res}} \right) > \frac{8\pi^2}{e^2}, \text{ and perturbation theory breaks down.}$$

Solution: Consider  $e^+e^- \rightarrow \mu^+\mu^- + N\gamma$ , and don't restrict to a fixed number of photons. This is no longer at a fixed order in the coupling  $e$ , but corresponds better to the physical situation where distinguishing 2 vs. 3 vs. 4 very low-energy photons isn't possible in practice. Will see this again with quarks and gluons in QCD!

Lessons from this week:

- QFT gives infinities when you ask it dumb (unphysical) questions. By relating amplitudes to a physically measurable quantity, we always get finite results.
- Singularities tend to appear beyond the lowest-order diagrams. Resolving them may require summing over several amplitudes coherently.
- Not all loop diagrams suffer from this complication: electron magnetic moment is one example.