

Hyper fine structure (+ general "nuclear effects")

Lecture #5
 PHYS 598 A00
 Fall 2017

so far, we've examined:

1) gross structure - $\left\{ \begin{array}{l} \text{electrostatic interactions} \\ \text{Hydrogen-like} \rightarrow \text{just } e^- \text{ charge} + \text{charge of nucleus} \\ \text{more generally} \rightarrow e^-e^- \text{ interactions (shielding)} \end{array} \right.$ also plays a big role

2) fine structure

- relativistic corrections
- spin-orbit interactions $\vec{S} \cdot \vec{L}$

relates to magnetic moment \leftarrow spin of electron

$|\vec{\mu}| \sim \mu_B = \frac{e\hbar}{2m_e}$

for single unpaired electron $A_{LS} = -\vec{\mu} \cdot \vec{B}$

$\vec{S} \cdot \vec{L}$ orbital ang. moment

for $l \neq 0$, effective B-field due to motion in $V(r)$ electro potential

recall $\vec{J} = \vec{L} + \vec{S}$

for $s = \frac{1}{2}$, have $j = l - \frac{1}{2}$ and $j = l + \frac{1}{2}$ if $l \neq 0$

$E(p_{3/2}) - E(p_{1/2}) = \Delta E_{s-o}$

general

$l-s$ to $l+s$
 for $l > s$;

$s-l$ to $s+l$
 for $s > l$

3) hyper fine - shifts + energy splitting due to nuclear effects

(smaller than fine structure)

nucleus is not a point-like charge,

but has structure and spin

we'll treat energy corrections due to ~~finite~~ finite size, variations in the mass of nucleus shortly

② the hyperfine structure, in terms of hyperfine energy levels, comes about due to the fact that the nucleus can have a spin angular momentum, and thus a magnetic moment.

We're going to try to give a rough sketch for how the nuclear spin gives rise to hyperfine structure. For a rigorous treatment based on a relativistic, quantum treatment, see lecture notes of W. Johnson

Let's start w/ the semi-classical picture as presented in Foot, which works well for hydrogen and will set the stage for a more general result.

⑨ <https://www3.nd.edu/~johnson/Publications/book.pdf>
Ch.#5

The nucleus has a magnetic moment that can be related to the nuclear spin \mathbf{I} as

$$\vec{\mu}_I = g_I \mu_N \mathbf{I} \rightarrow$$

where $\mu_N = \mu_B \frac{m_e}{m_p}$, i.e.

The nuclear magneton is lower than the Bohr magneton by the electron-to-proton mass ratio.

Different conventions
Note: you'll see several different formulations, depending on the source. Some have positive g_I , some have negative. Some absorb μ_N/μ_B factors into g_I , and express as $\vec{\mu}_I = g_I \mu_B \mathbf{I}$

$$I, (m_I = -I, -I+1, \dots, I)$$

3) Given this nuclear magnetic moment $\vec{\mu}_I$, there is an energy correction ~~due~~ in the presence of a magnetic field (from generic contributions) \vec{B}

$$H_{HFS} = -\vec{\mu}_I \cdot \vec{B} = -g_I \mu_N \vec{I} \cdot \vec{B}$$

while $g_s \approx 2$ for the electron spin
 g_I value will depend on nuclear structure

In the absence of an external field, it is the effective field due to the presence of electrons that shifts the energy. There are ~~two~~ ^{three} basic contributions to this field \vec{B}_e :

a) "field" produced by fact that electron has spin, and thus a magnetic moment. This can also be thought of in terms of the dipole-dipole interactions between $\vec{\mu}_I$ and $\vec{\mu}_e$ (when they're spatially separated), i.e.

$$\frac{\mu_0}{4\pi r^3} \left[\vec{\mu}_I \cdot \vec{\mu}_e - 3(\vec{\mu}_I \cdot \hat{r})(\vec{\mu}_e \cdot \hat{r}) \right]$$

this can also be recast as the field due to $\vec{\mu}_e$,

where $\vec{r} = \vec{r}_e - \vec{r}_N \approx \vec{r}_e$

$$\vec{B}_{e,s} = \frac{\mu_0}{4\pi} \left[\frac{\vec{\mu}_e}{r^3} - 3 \frac{(\vec{\mu}_e \cdot \hat{r}) \hat{r}}{r^3} \right]$$

$$= \frac{2\mu_0 \mu_B}{4\pi r^3} \left[\vec{S} - 3(\vec{S} \cdot \hat{r}) \hat{r} \right]$$

Note: This is just the classical expression for point-like dipoles, and we'll use an alternative approach for s-orbitals that overlap with the nucleus

b) the effective field that the nuclear spin feels due to the orbit of the electron. For states w/ $l \neq 0$, the "current" of the electron orbiting the nucleus leads to an effective field

contribution

$$\vec{B}_{e,l} \approx \frac{2\mu_0 \mu_B}{4\pi r^3} \vec{L}$$

(c) For the s-states, the contribution due to "dipole-dipole" averages to zero [anisotropic], but a Fermi contact interaction remains, which can be described in terms of the effective field

$$\vec{B}_{e,Fermi} = -\frac{2}{3} \mu_0 g_s \mu_B \vec{S} \delta^3(\vec{r}) \approx -\frac{2}{3} \mu_0 g_s \mu_B |\psi(0)|^2 \vec{S}$$

④ So, summing up,

$$H_{\text{HFS}} = -\vec{\mu}_{\text{I}} \cdot \vec{B} \quad \text{and} \quad \vec{\mu}_{\text{I}} = g_{\text{I}} \mu_{\text{N}} \vec{I}$$

$$\vec{B}_{\text{e}} = -\frac{2\mu_0 \mu_{\text{B}}}{4\pi} \left[\underbrace{\frac{\vec{L}}{r^3}}_{\text{orbit term}} - \underbrace{\frac{(\vec{S} - 3(\vec{S} \cdot \hat{r})\hat{r})}{r^3}}_{\text{dipole-dipole}} + \underbrace{\frac{4\pi}{3} g_{\text{S}} |\psi(0)|^2 \vec{S}}_{\text{Fermi contact}} \right]$$

For an s-orbital, the first two contributions are both zero, and we have for the $l=0$ state

$$H_{\text{HFS}} = \frac{2\mu_0 \mu_{\text{B}} g_{\text{I}} \mu_{\text{N}}}{4\pi} \frac{4\pi}{3} g_{\text{S}} |\psi(0)|^2 \vec{I} \cdot \vec{S}$$

for hydrogenic atoms w/ s-orbital $|\psi_{n0}(0)|^2 = \frac{1}{\pi} \left(\frac{Z}{na_0}\right)^3$

$$\Rightarrow H_{\text{HFS}} = \frac{2}{3\pi} \mu_0 g_{\text{I}} g_{\text{S}} \mu_{\text{B}} \mu_{\text{N}} \frac{Z^3}{n^3 a_0^3} \vec{I} \cdot \vec{S}$$

for this $l=0$ s-state, $\vec{J} = \vec{S}$ and the electron spin and nuclear spin add together as $\vec{F} = \vec{I} + \vec{J}$ w/ F-values of $I-J$ to $I+J$ (in steps of 1) for $I > J$

~~so~~ so $\vec{F} = \vec{I} + \vec{S}$

recall our treatment of $\vec{L} \cdot \vec{S}$ w/ $\vec{J} = \vec{L} + \vec{S}$

and $J-I$ to $J+I$ for $J > I$

note: mp values from $-F, -F+1, \dots, F-1, F$

$$\Rightarrow \hat{F}^2 = \hat{I}^2 + \hat{S}^2 + 2\vec{I} \cdot \vec{S}$$

expectation values

$$\hookrightarrow \vec{I} \cdot \vec{S} = \frac{1}{2} [\hat{F}^2 - \hat{I}^2 - \hat{S}^2] \rightarrow [F(F+1) - I(I+1) - S(S+1)]$$

~~4/15~~

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this gives us an energy correction to hydrogenic s-state levels of

$$\Delta E = \frac{2}{3\pi} \mu_0 g_I g_S \mu_B \mu_N \frac{Z^3}{n^3 a_0^3} \left[\frac{F(F+1) - I(I+1) - S(S+1)}{2} \right]$$

we expect that this should work fairly well for hydrogen + hydrogenic atoms, so let's check.

for the hydrogen nucleus (i.e. a proton), $Z=1$
 $g_I = 5.5883$
and $I = \frac{1}{2}$

for the electron we have $S = \frac{1}{2}$ and $g_S \approx 2$ (we'll look @ $n=1$)

The allowed F-values are $F = 0, 1$.

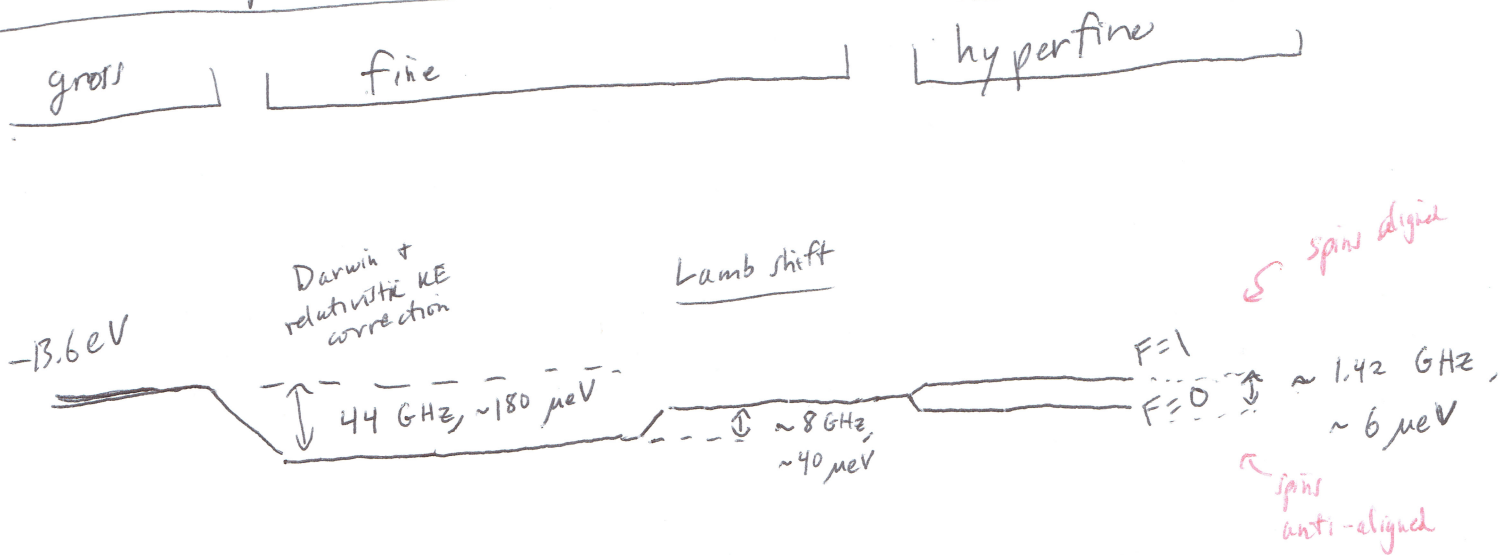
$$\Delta E_{F=0} = -\frac{3}{2} \left[\frac{2}{3\pi} (5.5883) \mu_0 \mu_B \mu_N \frac{1}{a_0^3} \right]$$

$$\Delta E_{F=1} = +\frac{1}{2} \left[\frac{2}{3\pi} \dots \dots \dots \right]$$

$$\left| \frac{\Delta E_{F=0}}{\Delta E_{F=1}} \right| = 3$$

$$\Delta E_{F=1} - \Delta E_{F=0} = \frac{4}{3\pi} g_I \mu_0 \mu_B \mu_N \frac{1}{a_0^3} \approx h \times 142 \text{ GHz}$$

this corresponds to $\lambda = 21 \text{ cm}$, used extensively in radio astronomy



⑥ This result fails for the more general case (multi-electron atoms, $l \neq 0$)

but a result having similar form follows from a proper treatment using relativistic quantum mechanics.

$$H_{\text{HFS}} = H_{\text{dip}} + H_{\text{quad}}$$

\uparrow magnetic dipole moment of nucleus
 \uparrow electric quadrupole moment of nucleus

(relevant for $I > \frac{1}{2}$ and $l \neq 0$, interaction of quadrupole moment w/ $\vec{\nabla}E$)

$$H_{\text{dip}} = A_{\text{hd}} \vec{I} \cdot \vec{J}$$

\uparrow same form as before

roughly we've calculated A for hydrogen-like atoms, but in general these values come from experiment.

$$H_{\text{quad}} = B_{\text{hfs}} \frac{3(\vec{I} \cdot \vec{J})^2 + \frac{3}{2}(\vec{I} \cdot \vec{J}) - I(I+1)J(J+1)}{2IJ(2I-1)(2J-1)} \quad \left\{ \begin{array}{l} \text{only for} \\ I \geq \frac{1}{2}, l \neq 0 \end{array} \right\}$$

F is a good quantum number here, as is its z -projection, m_F .
 ($\vec{F} = \vec{I} + \vec{J}$) \rightarrow will be good @ low B-fields

these $|I, J, F, m_F\rangle$ states ($|F, m_F\rangle$ for short) can be expressed in terms of the quantum #'s of the nuclear spin & electron spin as.

$$|F, m_F\rangle = \sum_{m_I, m_J} \langle F, m_F | I, m_I, J, m_J \rangle |I, m_I, J, m_J\rangle$$

~~$|I, J, F, m_F\rangle |I, m_I, J, m_J\rangle$~~

Clebsch-Gordan coefficient

$$\langle I, m_I, J, m_J | F, m_F \rangle = \sqrt{2F+1} (-1)^{I-J+m_F} \begin{pmatrix} I & J & F \\ m_I & m_J & -m_F \end{pmatrix}$$

Wigner 3j

⑦ The first-order energy shift is given by

$$\Delta E_{\text{HF}} = \frac{1}{2} A_{\text{HFS}} K + B_{\text{HFS}} \frac{\frac{3}{2} K(K+1) - 2I(I+1)J(J+1)}{2I(2I-1)2J(2J-1)}$$

where $K = F(F+1) - I(I+1) - J(J+1) = 2\langle \vec{I} \cdot \vec{J} \rangle$

see slides for some experimental values.

for some details on measurement of A_{HFS} & B_{HFS} constants, see RMP, 49 31 (1997) by Arimondo & Inguscio

Some notes: $A_{\text{HFS}} > 0$ is referred to as "regular",

whereas

$A_{\text{HFS}} < 0$, as in ^{40}K , is called "inverted"

For the alkalis, $S = \frac{1}{2}$, the ground state ($l=0$) Zeeman levels, i.e. the states $|F, m_F\rangle$ are constructed by superpositions of only 2 states in the $|I, m_I, J, m_J\rangle$ basis.

For $l=0$, we have $J=S$ such that we use $|I, m_I, S, m_S\rangle$, and call these states $|m_S, m_I\rangle$ for short.

recall that $m_F = m_S + m_I$, and $m_S = \pm \frac{1}{2}$, such that the allowed m_I values are $m_I = m_F - m_S = m_F \mp \frac{1}{2}$ for $m_S = \pm \frac{1}{2}$

In general, $|F, m_F\rangle = \alpha |m_S = \frac{1}{2}, m_I = m_F - \frac{1}{2}\rangle + \beta |m_S = -\frac{1}{2}, m_I = m_F + \frac{1}{2}\rangle$

α, β are Clebsch-Gordan coefficients.

This form will help us understand how we have states w/ vanishing magnetic moment ($m_F=0$) at low fields.

⑧

Other "nuclear" effects

We also get corrections to the energy levels due to details of the nucleus' mass, variations in its spin, and the non-zero size of the nucleus (not a point-like charge).

① Size effects → Volume shifts - in HW1, you calculated the energy corrections that one would expect due to the ~~modification~~ ^{deviation} of the nuclear charge distribution from that of a point charge, which led to a potential that was not proportional to $\frac{1}{r}$ for small r values

→ This effect is sometimes considered a hyperfine effect (comes from nuclear property) and sometimes absorbed as part of other effects (as in the discussion of proton size puzzle w/ Lamb shift measurements)

Simple toy model

(~ same as HW#1)
for hydrogen-like atom

Consider a sphere w/ uniform charge density, having total charge $Q_N = Ze$

and radius $R \approx A^{1/3} r_0$

where $A \approx 2Z$ is the atomic mass number (# of nucleons)

and $r_0 \approx 1.2$ fm (radius close to proton radius $r_p \approx 0.8$ fm)

The potential energy experienced by the electron

$$V(r) = \begin{cases} \frac{Ze^2}{4\pi\epsilon_0 R} \frac{1}{2} \left(\frac{r^2}{R^2} - 3 \right) & \text{for } r \leq R \\ = \frac{Ze^2}{4\pi\epsilon_0 r} & \text{for } r > R \end{cases}$$

④ This volume effect lead to an energy correction:

$$\Delta E = \langle H' \rangle \quad \text{where} \quad H' = V(r) - V_{\text{point charge}}$$

$$= \begin{cases} \frac{Ze^2}{4\pi\epsilon_0 R} \frac{1}{2} \left(\frac{r^2}{R^2} + \frac{2R}{r} - 3 \right) & \text{for } r \leq R \\ 0 & \text{for } r > R \end{cases}$$

$$\Delta E = \langle H' \rangle = \langle \psi_{nlm} | H' | \psi_{nlm} \rangle$$

because R is so small, this effect is only important (non-zero) for $l=0$ states

$$\Delta E = \frac{Ze^2}{4\pi\epsilon_0 2R} \int_0^R |R_{nl}(r)|^2 \left(\frac{r^2}{R^2} + \frac{2R}{r} - 3 \right) r^2 dr$$

$$\Delta E = \frac{Ze^2}{4\pi\epsilon_0 2R} |R_{n0}(0)|^2 \underbrace{\int_0^R \left(\frac{r^2}{R^2} + \frac{2R}{r} - 3 \right) r^2 dr}_{\frac{R^3}{5}}$$

$$= \frac{Ze^2}{10\epsilon_0} R^2 |R_{n0}(0)|^2 \rightarrow \frac{4Z^3}{n^3 a_0^3}$$

$$= \frac{Z^4}{a_0^3 n^3} \frac{e^2 R^2}{10\epsilon_0} \Rightarrow \text{Foot Eq (6.26)}$$

↳ result applied to alkalis, etc.
based on central field approximation

$$\frac{4Z^3}{n^3} a_0^3 Z$$

(10)

(2) Mass effects → we'll only discuss the reduced mass effect, ~~discuss~~

which requires a correction due to our assuming that $\frac{M_{\text{nuclear}}}{m_e} = \infty$

$$E_n = -\frac{e^2}{4\pi\epsilon_0 a_\mu} \frac{Z^2}{2n^2} = \underbrace{-\frac{e^2}{4\pi\epsilon_0 a_0} \frac{Z^2}{2n^2}}_{E_n^\infty} \frac{M_N}{m_e + M_N} = \frac{E_n^\infty}{\frac{m_e}{M_N} + 1} = E_n^\infty \left(1 + \frac{m_e}{M_N}\right)^{-1} \approx E_n^\infty \left(1 - \frac{m_e}{M_N}\right)$$

So, the correction for a finite nuclear mass

$$\text{ii } \Delta E_n = - E_n^\infty \frac{m_e}{M_N}$$

this correction always

$\propto -E_n^\infty$, such that it shifts energies up, and correction

term gets smaller as M_N increases.